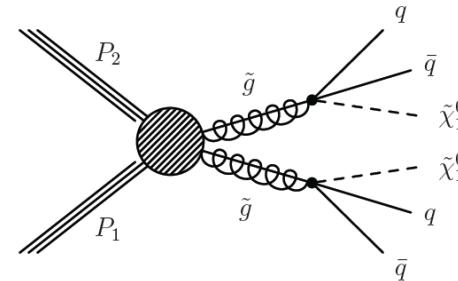


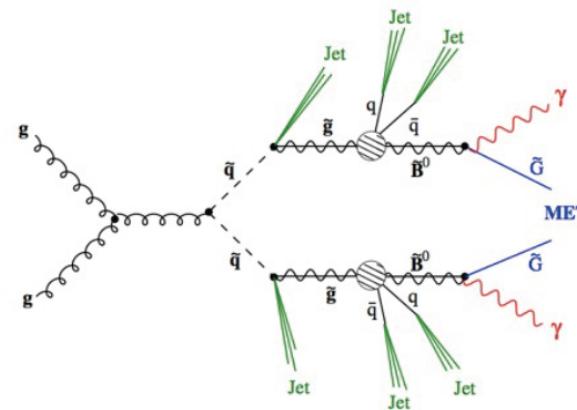
SUSY searches at the LHC

Purely hadronic:
Jets + missing energy

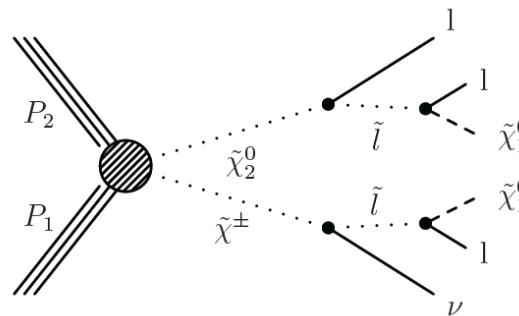


Photons:

- targets:
 - > gauge mediated SUSY
 - > dark matter
 - > large extra dimensions

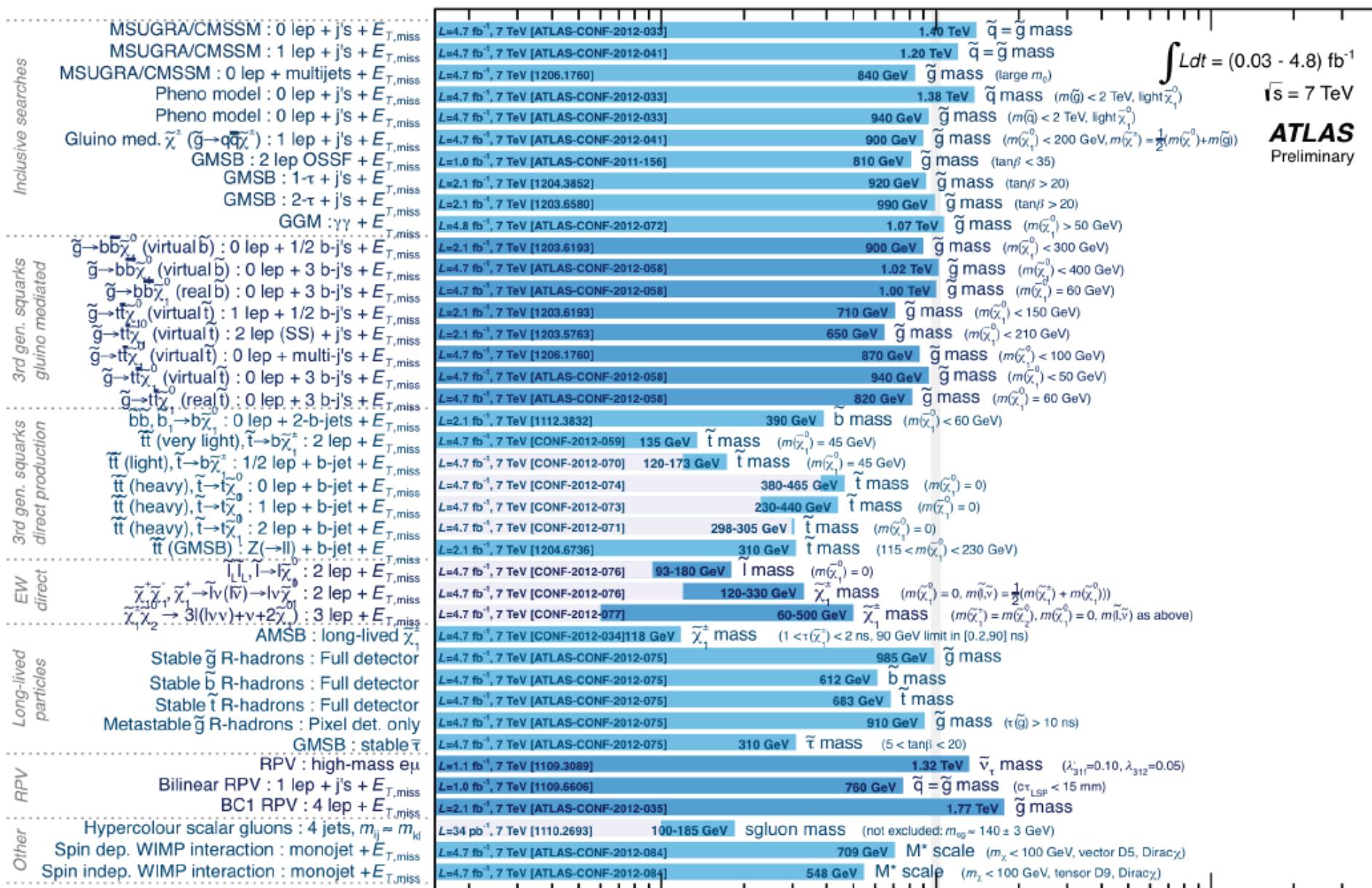


Leptonic:

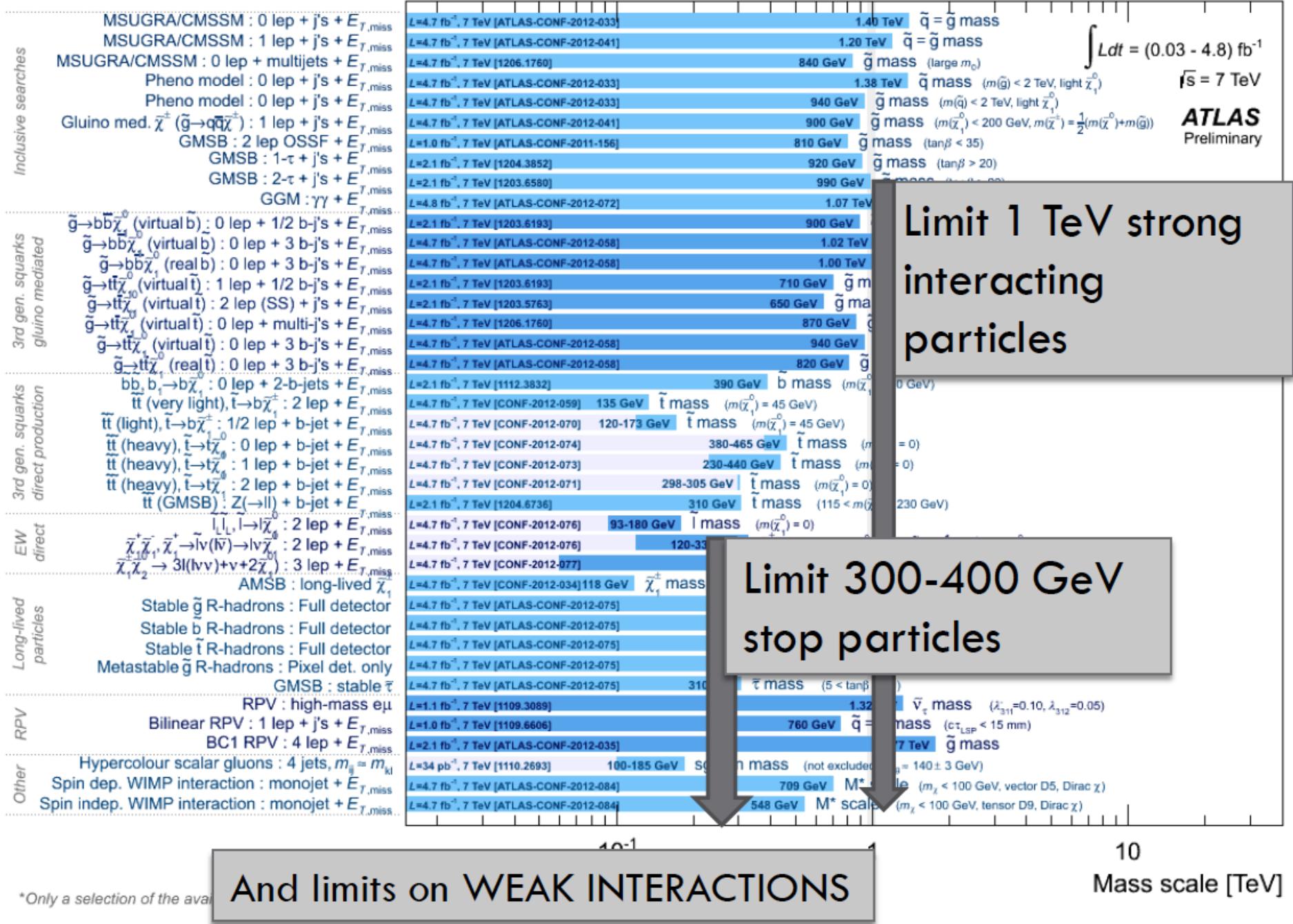


• • •

ATLAS SUSY Searches* - 95% CL Lower Limits (Status: ICHEP 2012)

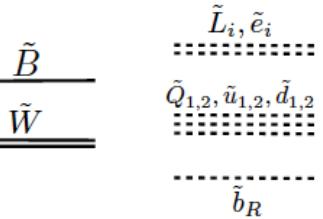
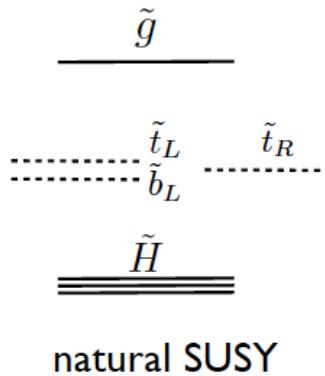
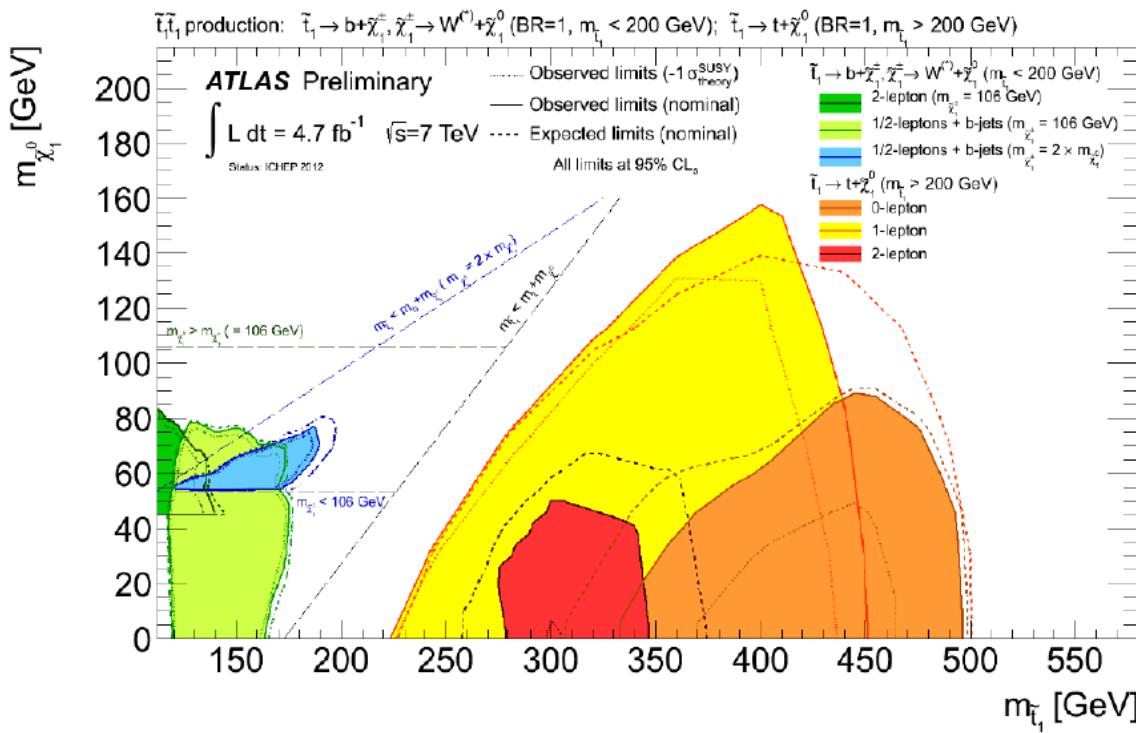


*Only a selection of the available mass limits on new states or phenomena shown

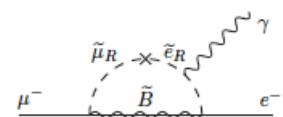
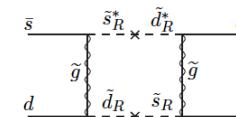


An exception: "Natural" SUSY

FCNC: 1,2 sgenerations heavy
 Hierarchy problem: 3rd sgeneration light



decoupled SUSY



Status of SUSY -little hierarchy problem

MSSM: 105 +(19) Parameters

$$M_Z^2 = \sum_{\tilde{q}, \tilde{l}} a_i \tilde{m}_i^2 + \sum_{\tilde{g}, \tilde{W}, \tilde{B}} \tilde{M}_i^2 + \dots$$

$$m_{\tilde{q}} > 0.6 - 1 TeV \Rightarrow \Delta > a \frac{\tilde{m}^2}{M_Z^2} \sim 100 \quad (\text{Unless light stop } m_{\tilde{t}, LHC} > 250 GeV)$$

⇒ Correlations between SUSY breaking parameters
and/or additional low-scale states

Status of SUSY -little hierarchy problem

MSSM: 105 +(19) Parameters

$$M_Z^2 = \sum_{\tilde{q}, \tilde{l}} \tilde{a}_i \tilde{m}_i^2 + \sum_{\tilde{g}, \tilde{W}, \tilde{B}} \tilde{M}_i^2 + \dots$$

$$m_{\tilde{q}} > 0.6 - 1 \text{TeV} \Rightarrow \Delta > a \frac{\tilde{m}^2}{M_Z^2} \sim 100 \quad (\text{Unless light stop } m_{\tilde{t},LHC} > 250 \text{ GeV})$$

⇒ Correlations between SUSY breaking parameters
and/or additional low-scale states

Fine Tuning measure:

$$\Delta(\gamma_i) = \left| \frac{\gamma_i}{M_Z} \frac{\partial M_Z}{\partial \gamma_i} \right|,$$

$$\Delta_{\max} = \text{Max}_{a_i} \Delta(\gamma_i)$$

$$\gamma_i = \{\tilde{m}_i, \tilde{M}_i, \dots\}$$

definite SUSY structure

Ellis, Enquist, Nanopoulos, Zwirner
Barbieri, Giudice

Fine Tuning measure:

$$\Delta(\gamma_i) = \left| \frac{\gamma_i}{M_Z} \frac{\partial M_Z}{\partial \gamma_i} \right|,$$

$$\Delta_{\max} = \text{Max}_{a_i} \Delta(\gamma_i)$$

Ellis, Enquist, Nanopoulos, Zwirner
 Barbieri, Giudice

Likelihood:

$$L(\text{data}|\gamma_i^0) = \frac{1}{\Delta_q} L(\text{data}|\gamma_i; v_0, \beta, \tilde{y}_t(\beta), \tilde{y}_b(\beta)) \Big|_{\beta=\beta_0(\gamma_i); \gamma_i=\gamma_i^0}$$

$$\chi^2(\gamma_i) = -2 \ln(L) = \chi^2_{old}(\gamma_i) + 2 \ln \Delta_q \quad \text{Ghilencea, Ross}$$

$$\Delta_q = \left(\sum_i \Delta_{\gamma_i}^2 \right)^{1/2}$$

$$\Delta_q = 100, \quad \delta\chi^2 / d.o.f. \sim 1$$

$$\Delta_q = 1000, \quad \delta\chi^2 / d.o.f. \sim 1.5$$

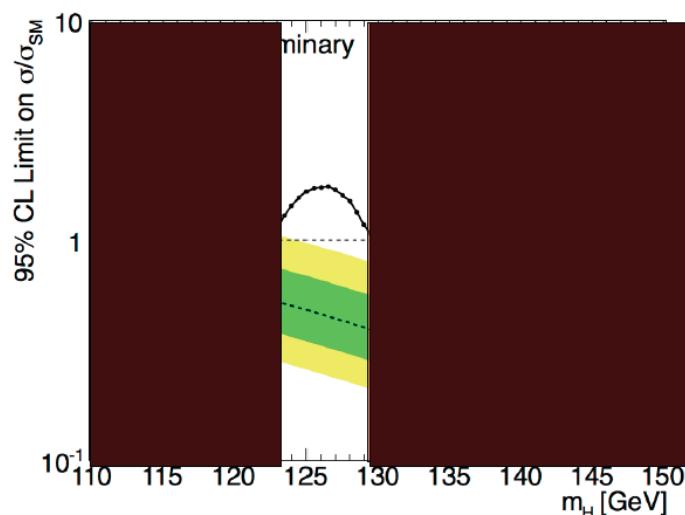
$$\chi^2 / d.o.f. = 2(2.5), \quad P = 3(0.6)\%$$

The Higgs mass in SUSY ?

$$M_S^2 = m_{q_3} m_{U_3} \geq (500 \text{ GeV})^2$$

$$M_{h^0}^2 = M_Z^2 \cos^2 2\beta + \frac{3M_t^2 h_t^2}{4\pi^2} \left(\ln\left(\frac{M_S^2}{M_t^2}\right) + \delta_t \right) + \dots \simeq 125 \text{ GeV (LHC)}$$

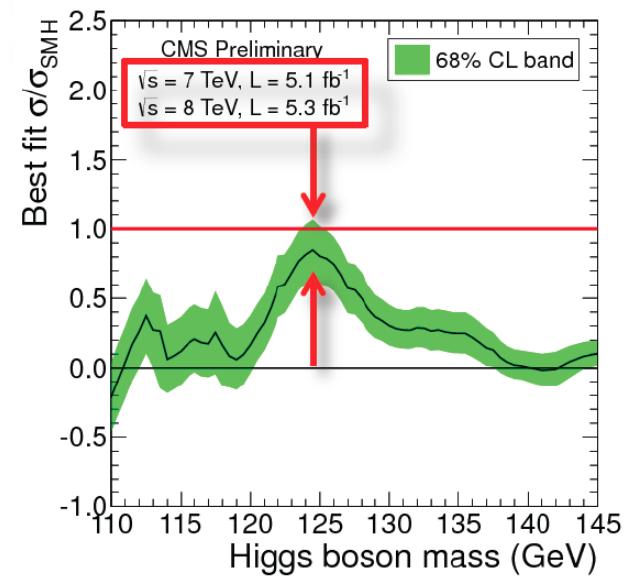
Atlas



Excluded at 95% CL

110-122.7 129-557 GeV

CMS



- The CMSSM

$$\mu_0, m_0, m_{1/2}, A_0, B_0$$


assume correlation between SUSY breaking parameters

● The CMSSM

$\mu_0, m_0, m_{1/2}, A_0, B_0$

$$\begin{aligned} V = & m_1^2 |H_1|^2 + m_2^2 |H_2|^2 - (m_3^2 H_1 \cdot H_2 + h.c.) \\ & + \frac{1}{2} \lambda_1 |H_1|^4 + \frac{1}{2} \lambda_2 |H_2|^4 + \lambda_3 |H_1|^2 |H_2|^2 + \lambda_4 |H_1 \cdot H_2|^2 \\ & + \left[\frac{1}{2} \lambda_5 (H_1 \cdot H_2)^2 + \lambda_6 |H_1|^2 (H_1 \cdot H_2) + \lambda_7 |H_2|^2 (H_1 \cdot H_2) + h.c. \right] \end{aligned}$$

Minimisation conditions:

$$\underline{v^2 = -m^2/\lambda}, \quad 2\lambda \frac{\partial m^2}{\partial \beta} = m^2 \frac{\partial \lambda}{\partial \beta}$$

$$\begin{aligned} m^2 &= m_1^2 \cos^2 \beta + m_2^2 \sin^2 \beta - m_3^2 \sin 2\beta \\ \lambda &= \frac{\lambda_1}{2} \cos^4 \beta + \frac{\lambda_2}{2} \sin^4 \beta + \frac{\lambda_{345}}{4} \sin^2 2\beta + \sin 2\beta (\lambda_6 \cos^2 \beta + \lambda_7 \sin^2 \beta) \end{aligned}$$

$$\Delta \equiv \max |\Delta_p|_{p=\{\mu_0^2, m_0^2, m_{1/2}^2, A_0^2, B_0^2\}}, \quad \Delta_p \equiv \frac{\partial \ln v^2}{\partial \ln p}$$

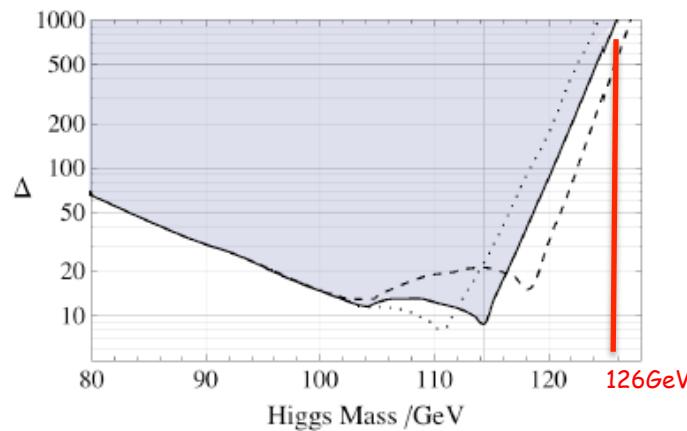
Couplings and masses evaluated to two loop (leading log) order

...enhanced sensitivity due to small tree-level $\lambda = \frac{1}{8} (g_1^2 + g_2^2) \cos^2 2\beta$

Cassel, Ghilencea, GGR
c.f. earlier work : Dimopoulos, Giudice, Chankowski, Ellis, Olechowski, Pokorski

• The CMSSM

$$\mu_0, m_0, m_{1/2}, A_0, B_0$$



Constraints

SUSY particle masses

$$3.20 < 10^4 \text{ Br}(b \rightarrow s\gamma) < 3.84$$

$$\text{Br}(b \rightarrow \mu\mu) < 1.8 \times 10^{-8}$$

$$\delta a_\mu < 292 \times 10^{-11}$$

$$-0.0007 < \delta\rho < 0.0012$$

Radiative EW breaking

Relic density unrestricted

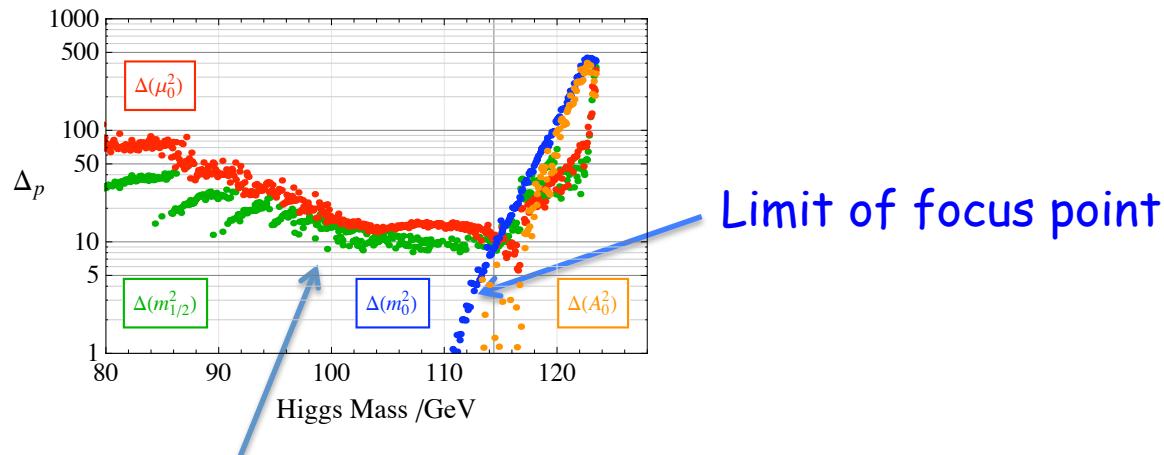
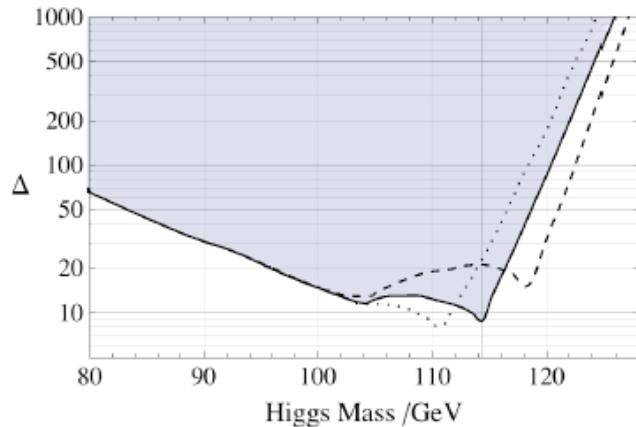
$$\Delta \equiv \max |\Delta_p|_{p=\{\mu_0^2, m_0^2, m_{1/2}^2, A_0^2, B_0^2\}}, \quad \Delta_p \equiv \frac{\partial \ln v^2}{\partial \ln p}$$

$$\Delta_{Min} = 9, \quad m_h = 114 \pm 2 \text{ GeV}$$

(No Higgs bound applied)

• The CMSSM

Constraints



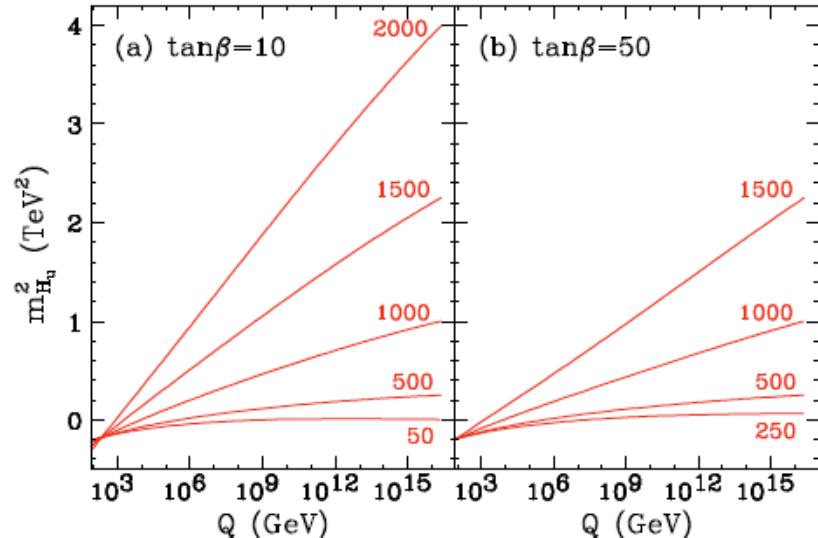
λ increase with m_H

$$v^2 = -\frac{m^2}{\lambda}$$

S.Cassel, D.Ghilencea, GGR

Focus Point

$$\begin{aligned}
 & 2 |y_t|^2 (m_{H_u}^2 + m_{Q_3}^2 + m_{u_3}^2) + 2 |a_t|^2 \\
 16\pi^2 \frac{d}{dt} m_{H_u}^2 &= 3X_t - 6g_2^2 |M_2|^2 - \frac{6}{5}g_1^2 |M_1|^2 \\
 16\pi^2 \frac{d}{dt} m_{Q_3}^2 &= X_t + X_b - \frac{32}{3}g_3^2 |M_3|^2 - 6g_2^2 |M_2|^2 - \frac{2}{15}g_1^2 |M_1|^2 \\
 16\pi^2 \frac{d}{dt} m_{u_3}^2 &= 2X_t - \frac{32}{3}g_3^2 |M_3|^2 - \frac{32}{15}g_1^2 |M_1|^2
 \end{aligned}$$



$$m_{H_u}^2(Q^2) = m_{H_u}^2(M_P^2) + \frac{1}{2} \left(m_{H_u}^2(M_P^2) + m_{Q_3}^2(M_P^2) + m_{u_3}^2(M_P^2) \right) \left[\left(\frac{Q^2}{M_P^2} \right)^{\frac{3y_t^2}{4\pi^2}} - 1 \right]$$

m_0^2 $3m_0^2$ $\simeq -\frac{2}{3}, Q^2 \simeq M_Z^2$

“Focus point”: $m_{H_u}^2(0) = m_{Q_3}^2(0) = m_{u_3}^2(0) \equiv m^2$

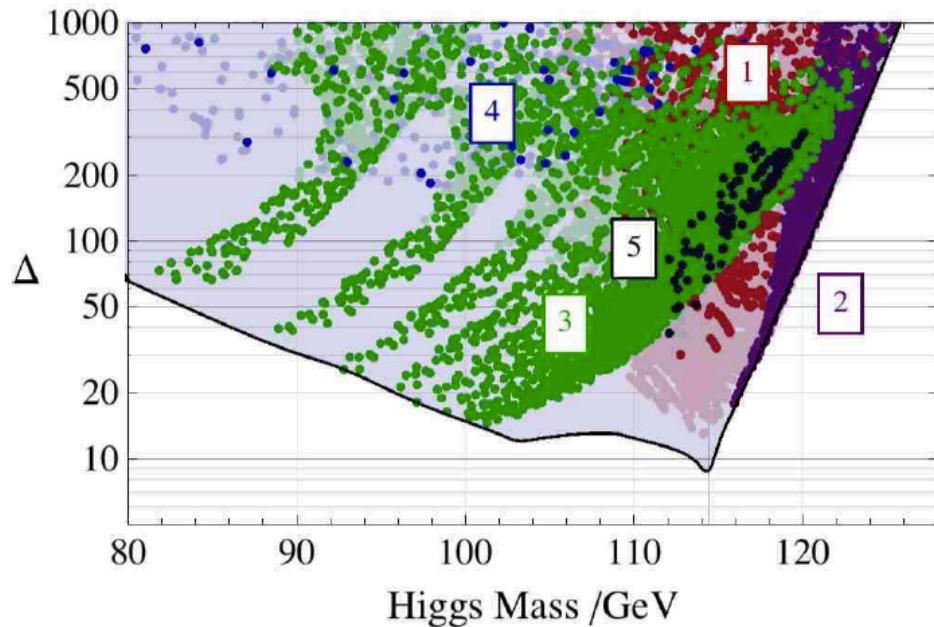
i.e. $m_{Q_3}^3, m_{u_3}^2 \gg M_Z^2$ possible

Natural choice

$$m_{H_u}^2(t_0) = a_0 m^2 + \dots, a_0 \leq 0.1$$

Feng, Matchev, Moroi
Chan, Chattopadhyay, Nath
Barbieri, Giudice
Feng, Sanford

Dark Matter structure



Relic density restricted

- 1 h^0 resonant annihilation
- 2 \tilde{h} t-channel exchange
- 3 $\tilde{\tau}$ co-annihilation
- 4 \tilde{t} co-annihilation
- 5 A^0 / H^0 resonant annihilation

Within 3σ WMAP:

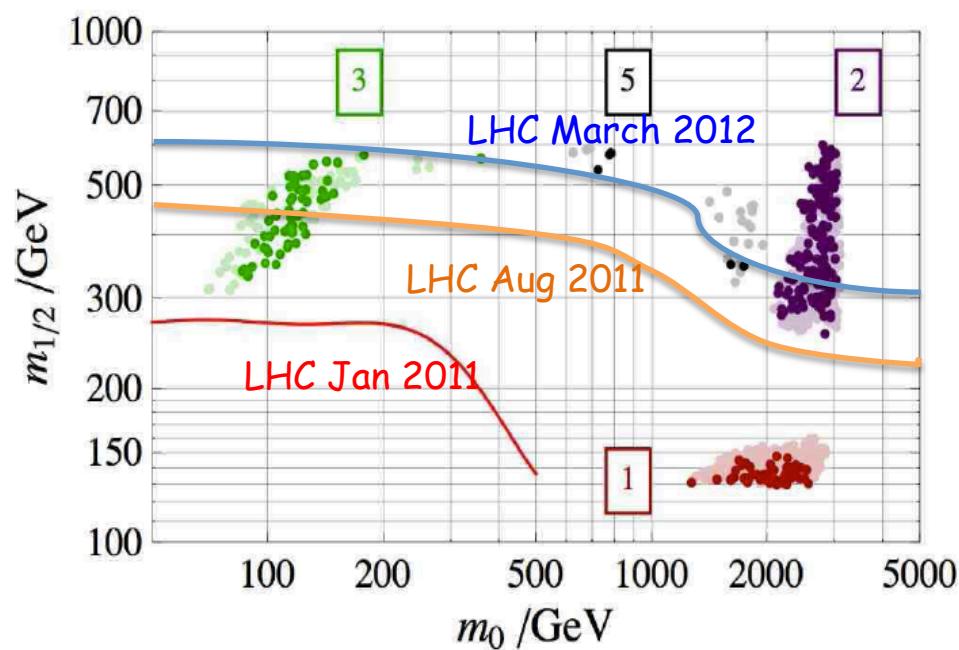
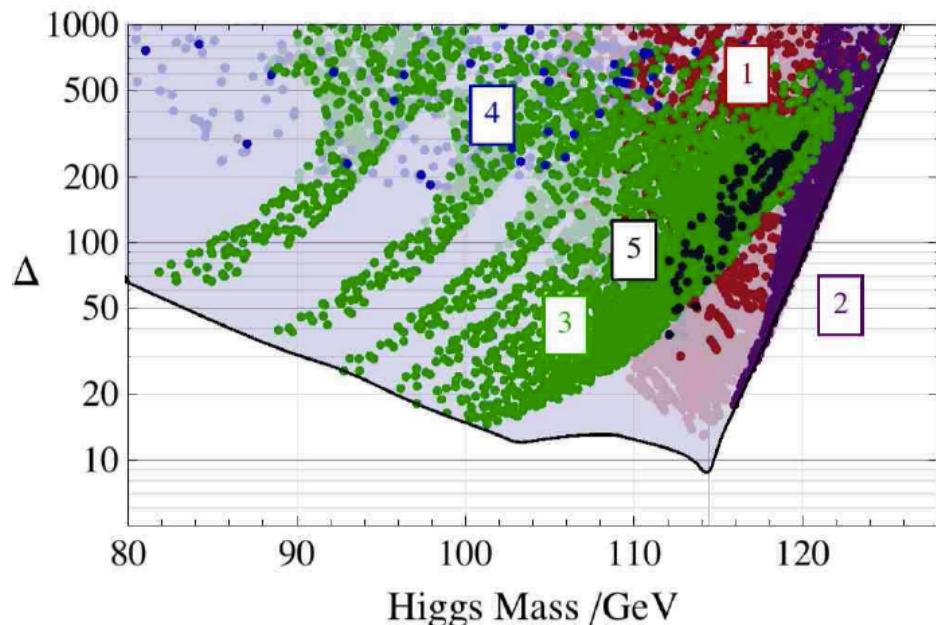
$$\Delta_{Min} = 15, \quad m_h = 114.7 \pm 2 \text{ GeV}$$

< 3σ WMAP:

$$\Delta_{Min} = 18, \quad m_h = 115.9 \pm 2 \text{ GeV}$$

Cassel, Ghilencea, GGR

Relic density restricted



- 1 h^0 resonant annihilation
- 2 \tilde{h} t-channel exchange
- 3 $\tilde{\tau}$ co-annihilation
- 4 \tilde{t} co-annihilation
- 5 A^0 / H^0 resonant annihilation

Within 3σ WMAP:

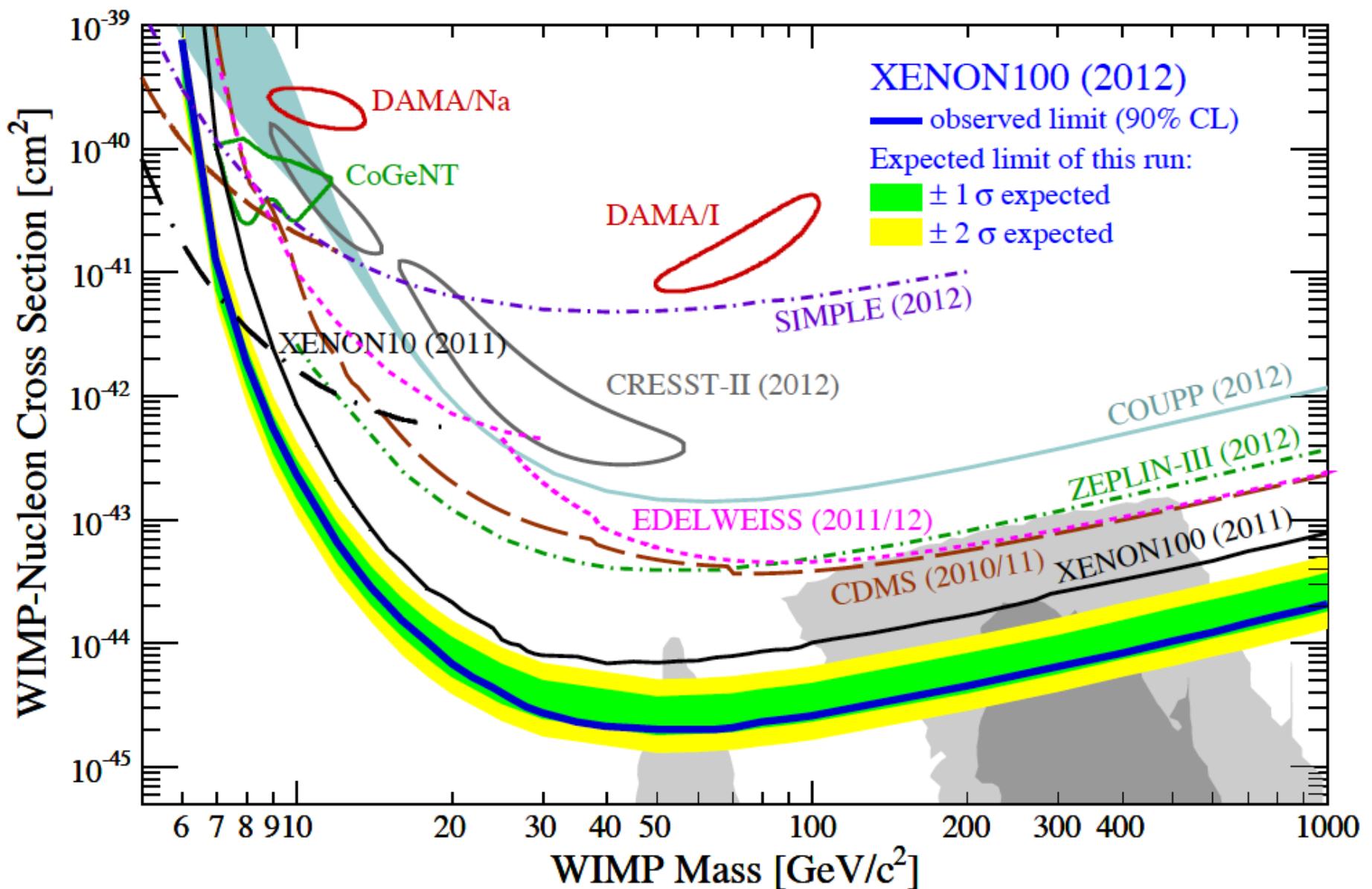
$$\Delta_{Min} = 15, \quad m_h = 114.7 \pm 2 \text{ GeV}$$

< 3σ WMAP:

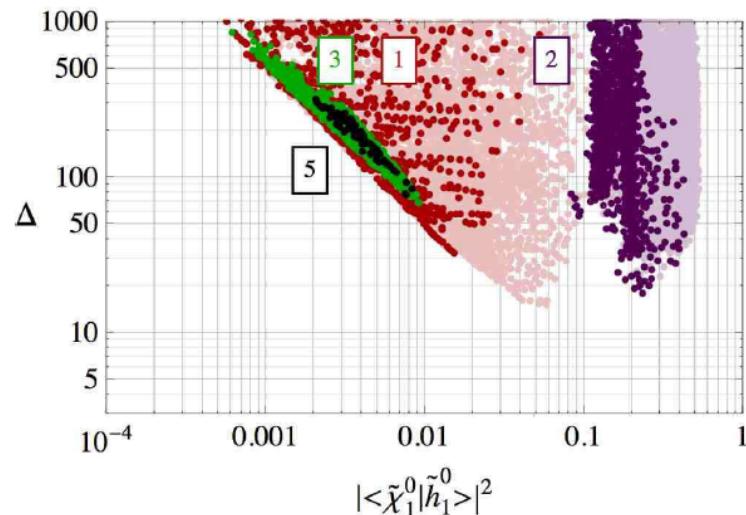
$$\Delta_{Min} = 18, \quad m_h = 115.9 \pm 2 \text{ GeV}$$

Cassel, Ghilencea, GGR

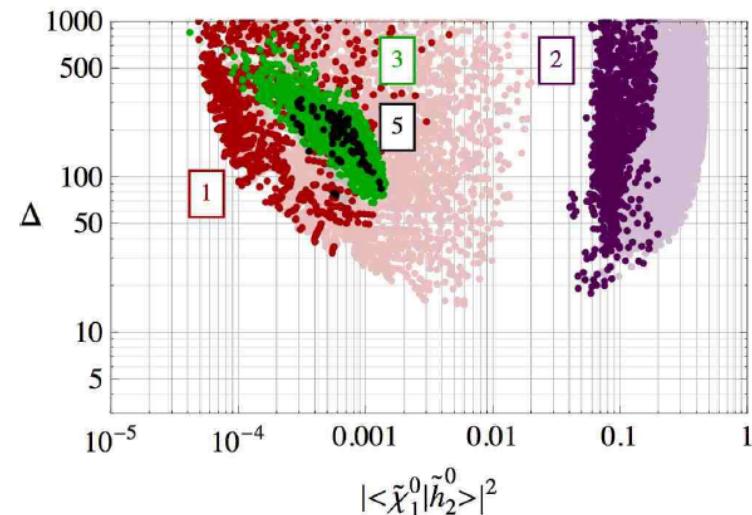
Direct dark matter searches: (spin independent)



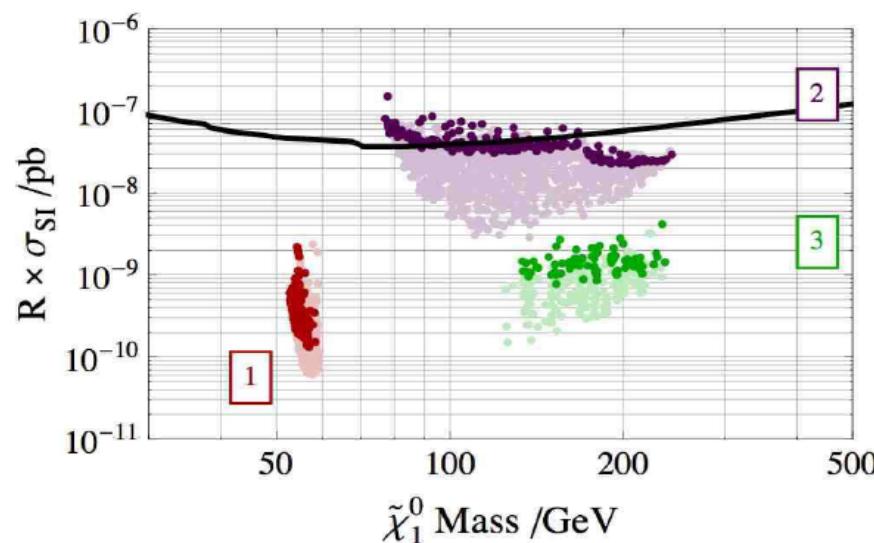
DM - Scaled spin independent cross section for LSP-proton scattering:



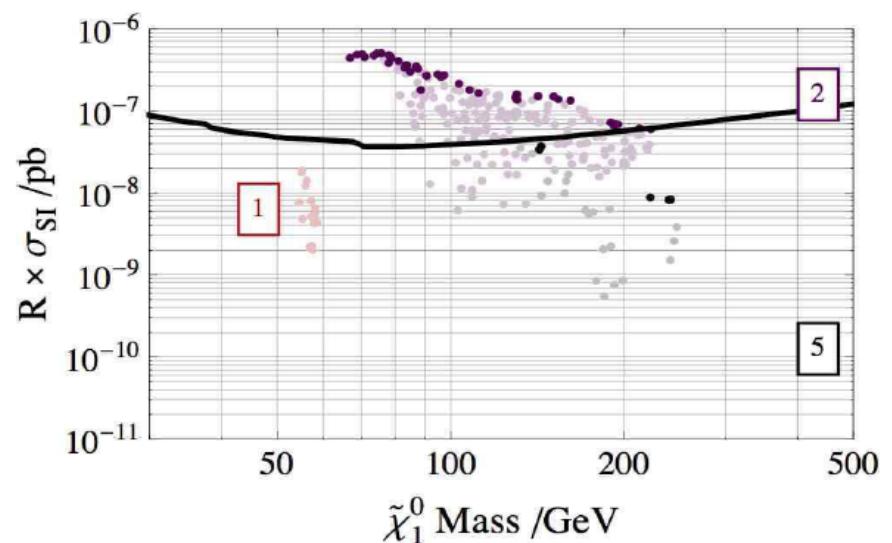
(a) LSP \tilde{h}_1^0 component



(b) LSP \tilde{h}_2^0 component



(a) $\tan \beta \leq 45$
 $\Delta < 100$



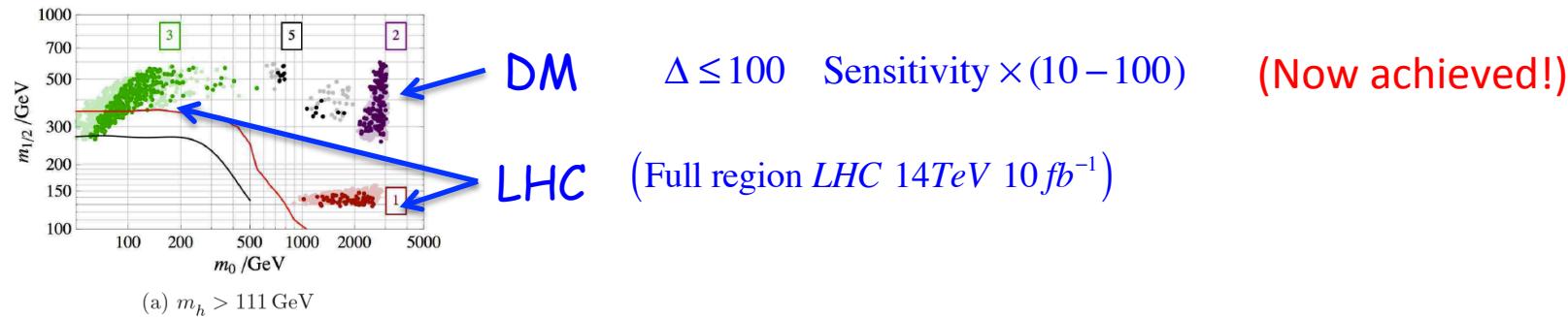
(b) $50 \leq \tan \beta \leq 55$
 $\Delta < 100$

CMSSM summary:

- Minimises MSSM fine tuning (focus point)

$$\text{Max}[\Delta_{EW}, \Delta_\Omega] = 15(29), \quad m_h = 114(116) \pm 2 \text{GeV}$$

- Complementary DM & LHC searches



- **BUT** $\Delta > 300$ for $m_H = 126 \text{ GeV}$

- (General) Gauge mediation in the MSSM

$$M_{\tilde{\lambda}_i}(M_{mess}) = k_i \frac{\alpha_i(M_{mess})}{4\pi} \Lambda_G$$

$$m_{\tilde{f}}^2(M_{mess}) = 2 \sum_{i=1}^3 C_i k_i \frac{\alpha_i^2(M_{mess})}{(4\pi)^2} \Lambda_S^2$$

$$k_i = \left(\frac{5}{3}, 1, 1\right)$$

$$k_i \alpha_i(M_{GUT}) = 1, \quad i = 1, 2, 3$$

No focus point

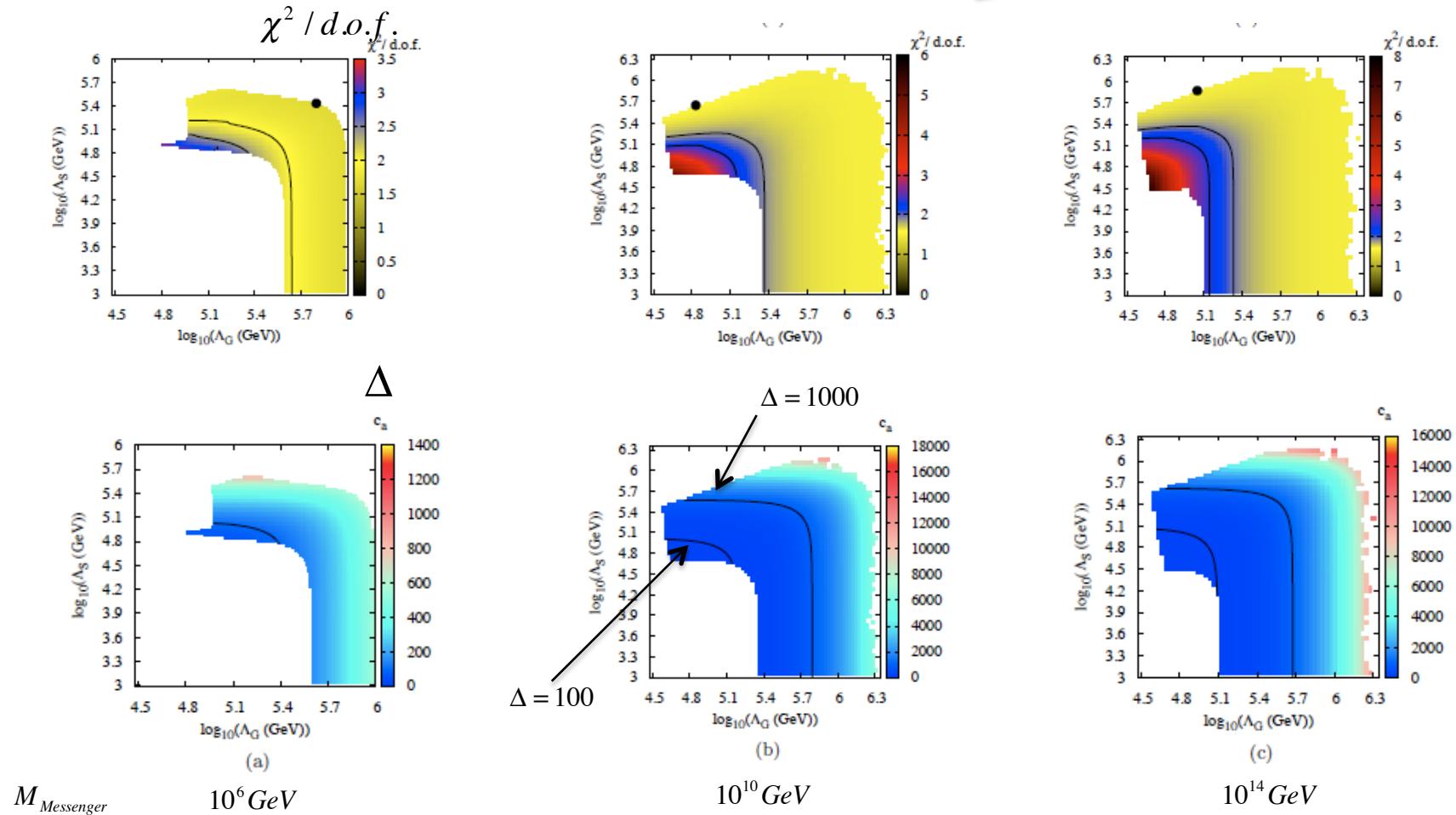


(Ordinary gauge mediation $\Lambda_G = \Lambda_S$)

Meade, Seiberg, Shih

Fine tuning in General Gauge Mediation

$B \rightarrow X_s \gamma, B \rightarrow \tau \mu, B \rightarrow \mu^+ \mu^-, B \rightarrow D \tau \mu,$
 $D_s \rightarrow \mu \nu, D_s \rightarrow \tau \nu, K \rightarrow \mu \nu / \pi \rightarrow \mu \nu, \Delta_0$



$\Delta >> 100$ no focus point

Abel, Dolan, Jaeckel, Khoze
(Giusti, Romanino, Strumia)

Reduced fine tuning (c.f. CMSSM)

- New degrees of freedom

- New focus points?

Gauginos:

$M_{\tilde{g}, \tilde{W}, \tilde{B}}$ Non-universal gaugino correlations

Kane, King
Lebedev, Nilles, Ratz...
Horton, GGR
...

Scalars:

M_0, A_0 correlations?

Feldman, Kane, Kuflik, Lu

$$M_0, A_0, B_0 \gg \mu, m_a$$

$$m_{h_u^2}^2(t) = f_{M_0} M_0^2 - f_{A_0} A_0^2$$

$$f_{M_0} \sim f_{A_0} \sim 0.1 + M_0 \sim A_0 \Rightarrow$$

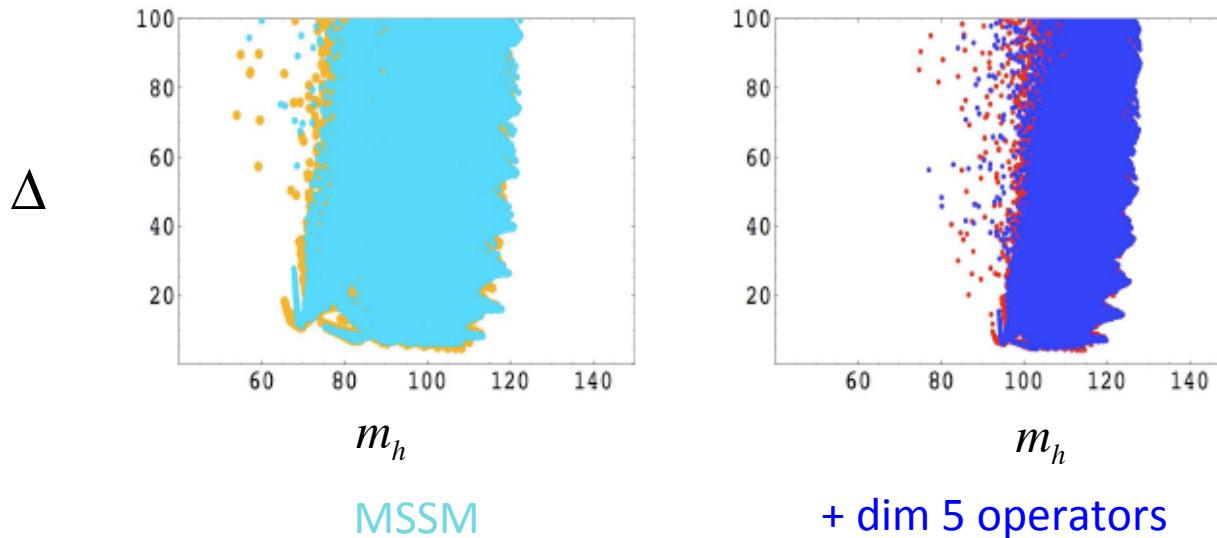
$$m_{h_u}^2 \sim 10^{-2} m_{3/2}^2$$

$$\Delta_{h_t} \sim 10^4, \quad M_0 \sim 10 \text{TeV}$$

I. Reduced fine tuning : New heavy states - higher dimension operators

$$\delta L = \int d^2\theta \frac{1}{M_*} (\mu_0 + c_0 S) (H_1 H_2)^2, \quad S = m_0 \theta \bar{\theta} \quad \text{Dimension 5}$$

$$\delta V = \zeta_1 (|h_1|^2 + |h_2|^2) h_1 h_2 + \zeta_2 (h_1 h_2)^2; \quad \zeta_1 = \frac{\mu_0}{M_*}, \quad \zeta_2 = \frac{c_0 m_0}{M_*}$$



Cassel, Ghilencea, GGR
 Casas, Espinosa, Hidalgo
 Dine, Seiberg, Thomas
 Batra, Delgado, Tait
 Kaplan,

Even for $M_* = 65$ μ_0 a significant shift of m_h for constant Δ

...effect mainly comes from ζ_1 term ... origin?

Reduced fine tuning : singlet extensions

$$W = W_{\text{Yukawa}} + (\mu + \lambda S) H_u H_d + \frac{\mu_S}{2} S^2 + \frac{\kappa}{3} S^3 + \xi S$$

c.f. $W = W_{\text{Yukawa}} + \lambda S H_u H_d + \frac{\kappa}{3} S^3$

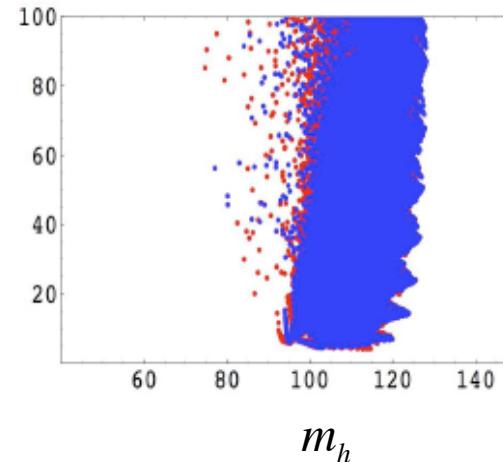
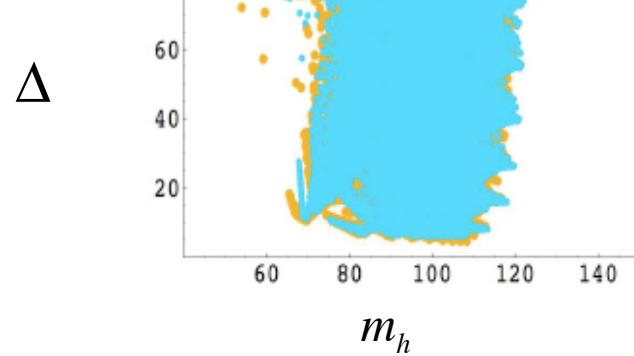
GNMSSM

$$\mu_S \gg m_{3/2}^\dagger$$

$$W_{\text{eff}}^{\text{GNMSSM}} = (H_u H_d)^2 / \mu_s + \mu H_u H_d$$

NMSSM

$$\frac{\mu}{\mu_s} (|H_u|^2 + |H_d|^2) H_u H_d \quad \text{v}^2 = -\frac{m^2}{\lambda}$$



Reduced fine tuning in GNMSSM (but not NMSSM)

SUSY extensions of the Standard Model

R-symmetry ensures Singlet extensions natural

SUSY extensions of the Standard Model

NMSSM spectrum

No perturbative μ term

Commutes with $SO(10)$

Anomaly cancellation

N	q_{10}	$q_{\bar{5}}$	q_{H_u}	q_{H_d}	q_S
4	1	1	0	0	2
8	1	5	0	4	6

Can be embedded in $SO(10)$

D=5 operators

up and down Yukawas allowed

$$3q_{10} + q_{\bar{5}} + q_{H_u} + q_{H_d} = 4 \pmod{N} \Rightarrow 3q_{10} + q_{\bar{5}} = 0 \pmod{N} \Rightarrow \frac{1}{M} Q \cancel{Q} L - \frac{1}{M} LL H_u H_u$$

Weinberg operator

R-symmetry ensures singlets light

SUSY breaking

$\langle W \rangle, \langle \lambda \lambda \rangle$ R=2 non-perturbative breaking

$Z_{4,8}^R \rightarrow Z_2^R$ R-parity

Domain walls and tadpoles safe

Abel

$\mu \sim m_{3/2}, O(\frac{m_{3/2}}{M^2} Q \cancel{Q} L)$

SUSY extensions of the Standard Model

NMSSM spectrum

No perturbative μ term

Commutes with $SO(10)$

Anomaly cancellation

N	q_{10}	$q_{\bar{5}}$	q_{H_u}	q_{H_d}	q_S
4	1	1	0	0	2
8	1	5	0	4	6



R-symmetry ensures singlets light

D=5 operators

up and down Yukawas allowed

$$3q_{10} + q_{\bar{5}} + q_{H_u} + q_{H_d} = 4 \pmod{N} \Rightarrow 3q_{10} + q_{\bar{5}} = 0 \pmod{N} \Rightarrow \frac{1}{M} Q \cancel{Q} Q L - \frac{1}{M} L L H_u H_u$$

Weinberg operator

SUSY breaking

$\langle W \rangle, \langle \lambda \lambda \rangle$ R=2 non-perturbative breaking

$Z_{4,8}^R \rightarrow Z_2^R$ R-parity

Domain walls and tadpoles safe

Abel

$\mu \sim m_{3/2}, O(\frac{m_{3/2}}{M^2} Q Q Q L)$

$$W = W_{MSSM} + \lambda S H_u H_d + \kappa S^3 + \Delta W$$

$$\Delta W_{Z_4^R} \sim m_{3/2} H_u H_d + m_{3/2}^2 S + m_{3/2} S^2$$

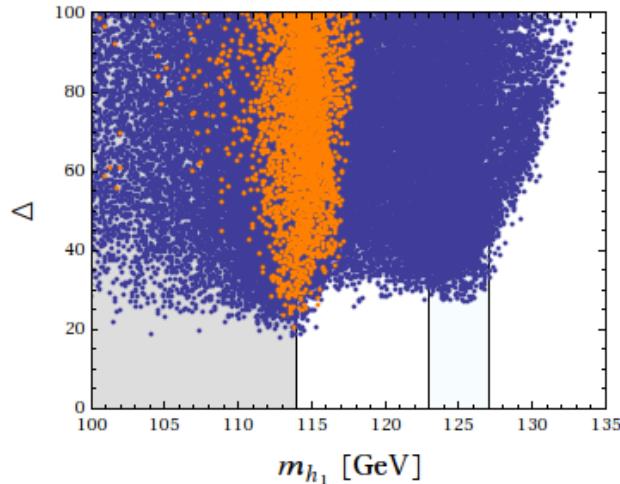
$$\Delta W_{Z_8^R} \sim m_{3/2}^2 S$$

← μ term and mass terms “natural”

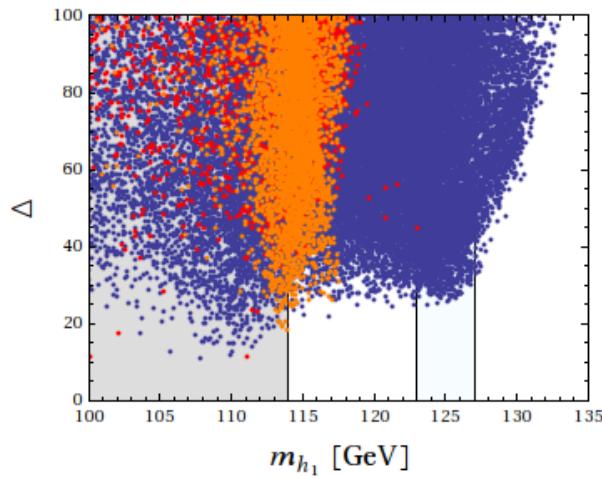
GNMSSM (c.f. NMSSM)

Fine tuning in the GNMSSM

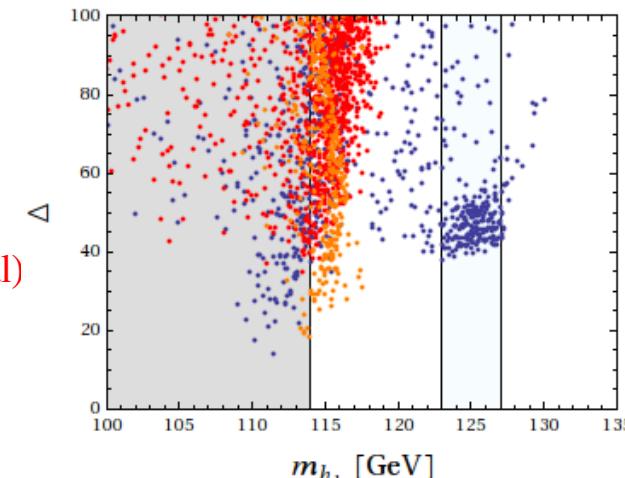
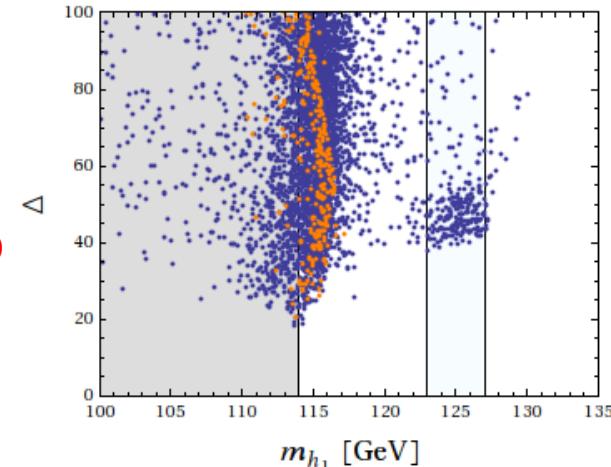
$(\lambda \leq 0.7^\dagger)$



- CMSSM
- CGNMSSM
(universal masses)



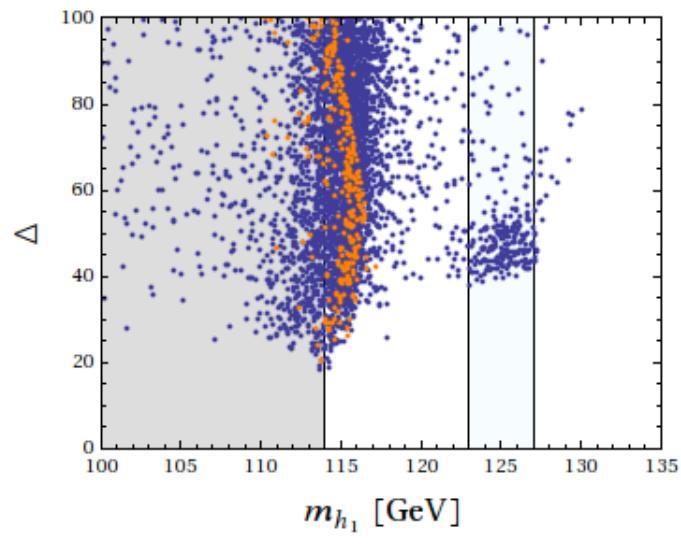
LHC constraints
applied



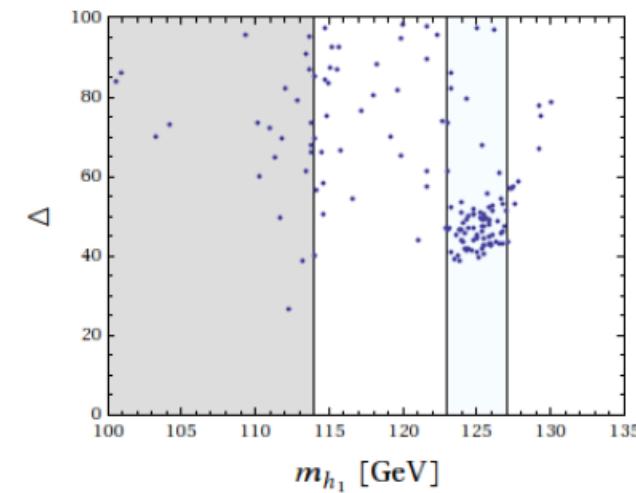
LHC + DM constraints
applied

GGR, Schmidt-Hoberg , Staub
† c.f. Hall, Pinner, Ruderman

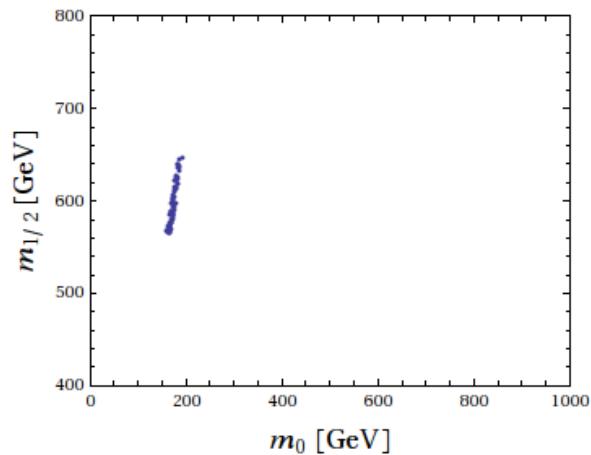
Dark Matter structure



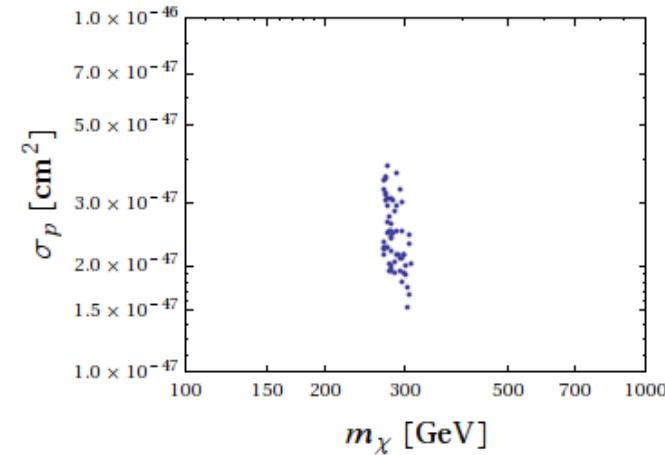
$$\rho_{LSP} \leq \rho_{DM}$$



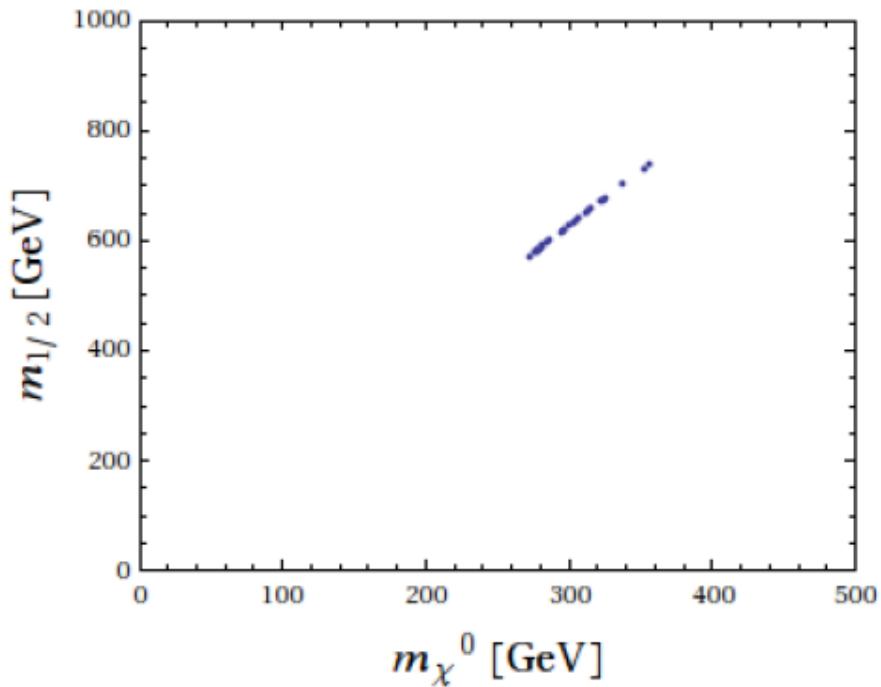
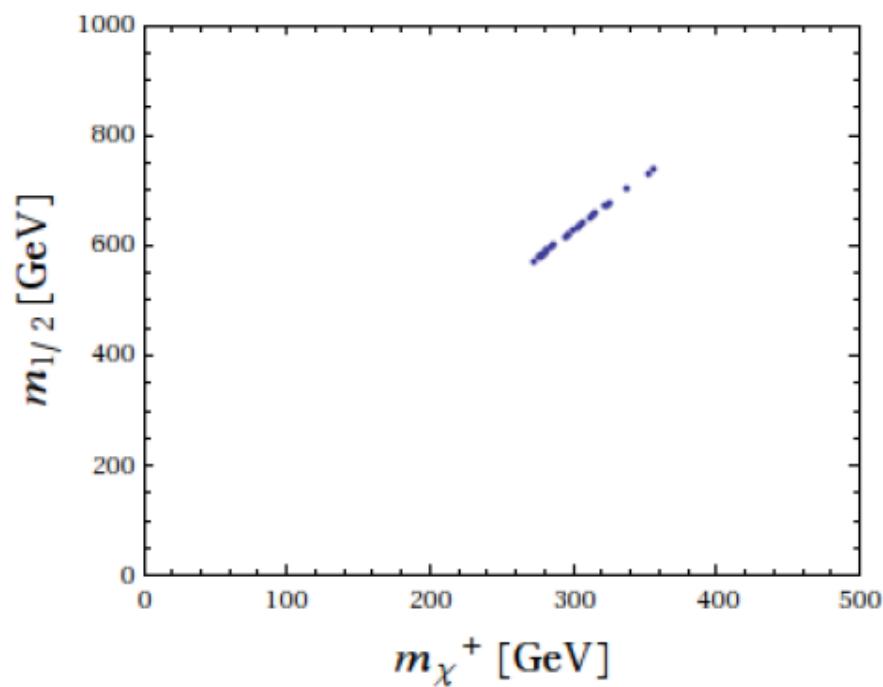
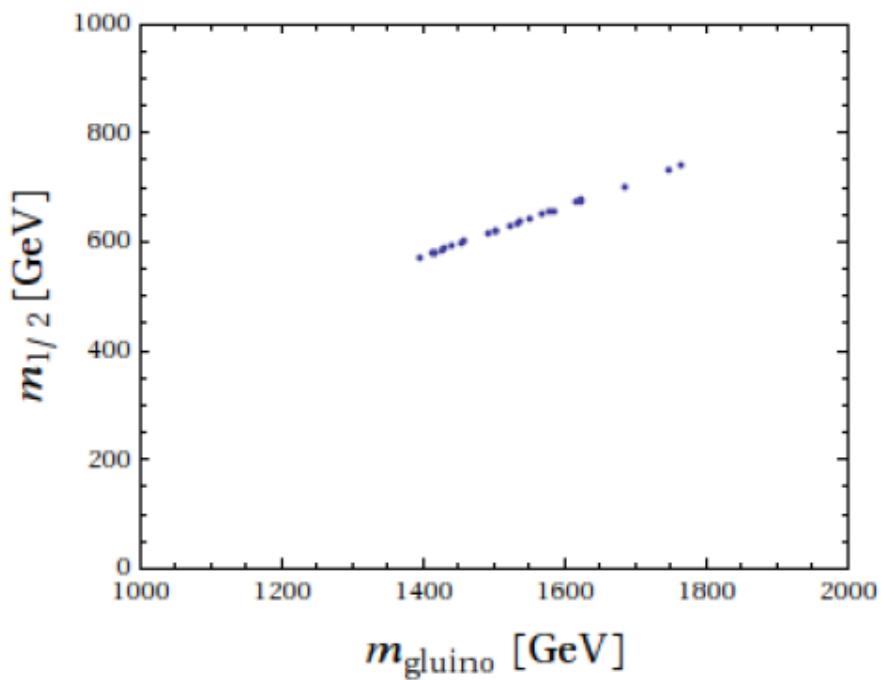
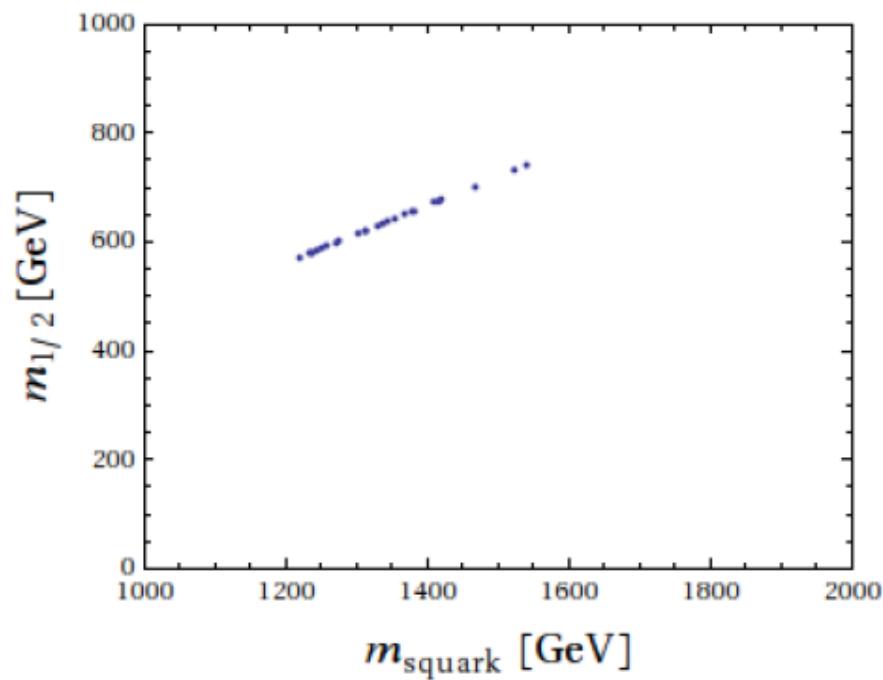
$$\rho_{LSP} \approx \rho_{DM}$$



Stau co-annihilation



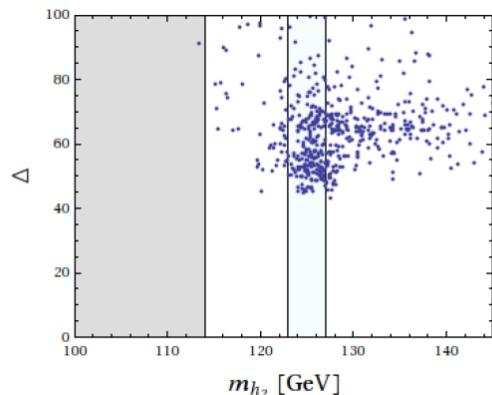
DM searches insensitive



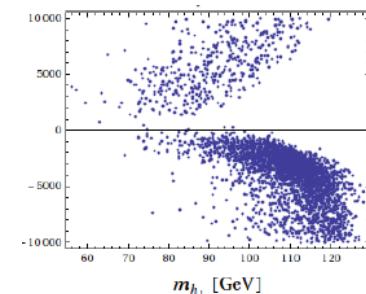
GENERAL-NMSSM PHENOMENOLOGY

Higgs structure (h_u, h_d, s)

- $\mu_s \gg \mu$ MSSM SUSY structure with heavy Higgs
- $\mu_s, m_s, b_s \sim \mu$ $h_1 \simeq H_{u,d} + \varepsilon S, \quad h_2 = S - \varepsilon H_{u,d}$
... h_2 may be lighter than LEP bound



m_{h_1} v/s Δ for the case $m_{h_2} < m_{h_1}$



Benchmark points

	BP1	BP2	BP3	BP4	BP5
m_0 [GeV]	746	163	957	573	752
$m_{1/2}$ [GeV]	476	568	557	482	472
$\tan \beta$	2.7	2.9	2.8	3.4	2.8
A_0 [GeV]	1433	1666	782	27	-198
λ	1.43	1.47	1.58	1.34	1.12
κ	-0.1	0.09	-0.005	1.52	1.03
A_λ [GeV]	A_0	A_0	A_0	400	192
A_κ [GeV]	A_0	A_0	A_0	-323	-326
v_S [GeV]	-841	-190	-929	390	281
μ_S [GeV]	-5931	-5354	-5799	131	-37
$m_{h_d}^2$ [GeV 2]	m_0^2	m_0^2	m_0^2	$9.1 \cdot 10^5$	$5.4 \cdot 10^5$
$m_{h_u}^2$ [GeV 2]	m_0^2	m_0^2	m_0^2	$2.3 \cdot 10^6$	$2.4 \cdot 10^6$
m_s^2 [GeV 2]	m_0^2	m_0^2	m_0^2	$2.8 \cdot 10^6$	$1.7 \cdot 10^6$
μ [GeV]	-750	-1136	-934	-33	10
$b\mu$ [GeV 2]	$-2.4 \cdot 10^6$	$-1.2 \cdot 10^6$	$-2.3 \cdot 10^6$	147	26
b_s [GeV 2]	$-1.9 \cdot 10^7$	$-5.4 \cdot 10^6$	$-1.4 \cdot 10^7$	326	144
ξ_s [GeV 3]	$2.2 \cdot 10^9$	$1.5 \cdot 10^9$	$3.0 \cdot 10^9$	22	-8
m_{squark} [GeV]	1256-1293	1207-1263	1507-1548	1211-1248	1280-1315
$m_{\tilde{g}}$ [GeV]	1219	1389	1416	1242	1235
m_{h_1} [GeV]	124	123.5	125	93.5	78
m_{h_2} [GeV]	1002	856	1257	125	124
h_1 singletfraction	$\mathcal{O}(10^{-4})$	$\mathcal{O}(10^{-6})$	$\mathcal{O}(10^{-4})$	0.8	0.85
$\text{Br}(h \rightarrow \gamma\gamma)$	$2.29 \cdot 10^{-3}$	$2.28 \cdot 10^{-3}$	$2.2 \cdot 10^{-3}$	$2.5 \cdot 10^{-3}$	$2.66 \cdot 10^{-3}$
$\text{Br}(b \rightarrow s\gamma)$	$3.1 \cdot 10^{-4}$	$3.1 \cdot 10^{-4}$	$3.1 \cdot 10^{-4}$	$3.1 \cdot 10^{-4}$	$3.3 \cdot 10^{-4}$
Δa_μ	$-7.8 \cdot 10^{-11}$	$-2.5 \cdot 10^{-10}$	$-5.4 \cdot 10^{-11}$	$1.7 \cdot 10^{-10}$	$8 \cdot 10^{-11}$
$\delta\rho$	$6.2 \cdot 10^{-5}$	$6.6 \cdot 10^{-5}$	$7.5 \cdot 10^{-5}$	$1.9 \cdot 10^{-4}$	$3.1 \cdot 10^{-4}$
$m_{\tilde{\chi}_1^0}$ [GeV]	229	270	168	99	70
$\tilde{\chi}_1^0$ singlinofraction	$\mathcal{O}(10^{-5})$	$\mathcal{O}(10^{-5})$	$\mathcal{O}(10^{-5})$	0.1	0.2
Ωh^2	7.5	0.10	7.4	0.017	0.11
σ_p [cm 2]	$2.8 \cdot 10^{-47}$	$2.2 \cdot 10^{-47}$	$6 \cdot 10^{-47}$	$1.2 \cdot 10^{-44}$	$1.3 \cdot 10^{-45}$
Δ (Fine-tuning)	34.9	51.0	51.8	44.9	52.7

Table 1: Benchmark scenarios for the GNMSSM for the universal (BP1-BP3) and the general (BP4-BP5) case. m_{squark} shows the range of squark masses of the first two generations. For the last two points the second lightest Higgs is mostly MSSM-like.

GENERAL-NMSSM PHENOMENOLOGY

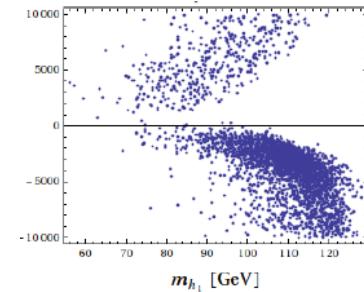
Higgs structure (h_u, h_d, s)

- $\mu_s \gg \mu$ MSSM SUSY structure with heavy Higgs

- $\mu_s, m_s, b_s \sim \mu$ $h_1 \simeq H_{u,d} + \epsilon S, \quad h_2 = S - \epsilon H_{u,d}$

... h_2 may be lighter than LEP bound

$h_1 \rightarrow h_2 h_2$ large



... h_1 may have enhanced $\gamma\gamma$ rate (later)

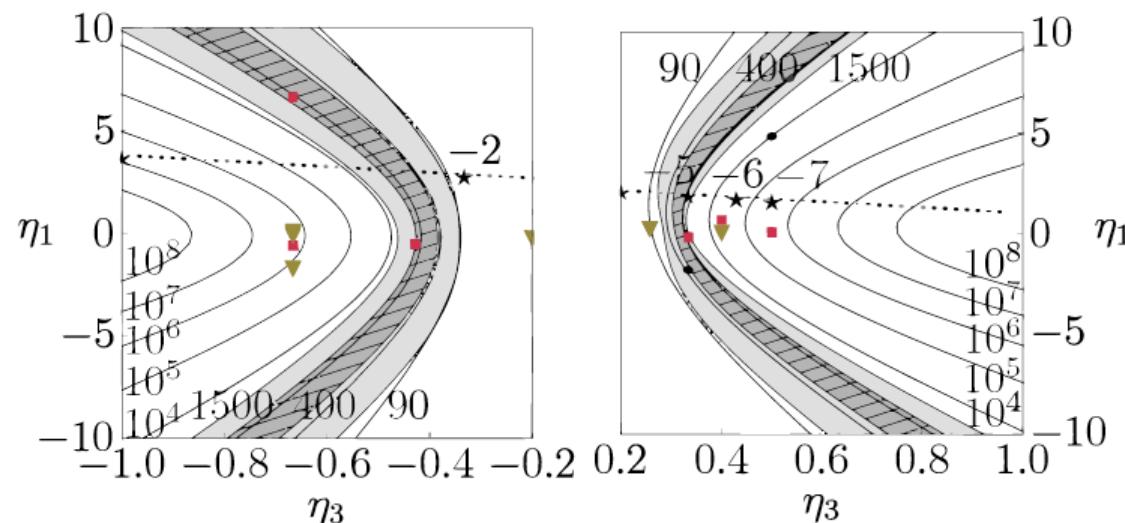
... DM and SUSY phenomenology modified

II. Reduced fine tuning : nonuniversal gaugino masses

Reduced fine tuning: nonuniversal gaugino masses

$$16\pi^2 \frac{d}{dt} m_{H_u}^2 = 3 \left(2 |y_t|^2 (m_{H_u}^2 + m_{Q_3}^2 + m_{u_3}^2) + 2 |a_t|^2 \right) - 6g_2^2 |M_2|^2 - \frac{6}{5} g_1^2 |M_1|^2$$

New focus point: cancellation between M_3 and M_2 contributions if $|M_2|^2 \simeq |M_3|^2$ at M_{SUSY}



$$M_3 : M_2 : M_1 = \eta_3 : 1 : \eta_1$$

Reduced fine tuning: nonuniversal gaugino masses

$$16\pi^2 \frac{d}{dt} m_{H_u}^2 = 3 \left(2 |y_t|^2 (m_{H_u}^2 + m_{Q_3}^2 + m_{u_3}^2) + 2 |a_t|^2 \right) - 6g_2^2 |M_2|^2 - \frac{6}{5} g_1^2 |M_1|^2$$



New focus point: cancellation between M_3 and M_2 contributions if $|M_2|^2 \simeq |M_3|^2$ at M_{SUSY}

Natural ratios? e.g.:

GUT: $SU(5)$: $\Phi^N \subset (24 \times 24)_{symm} = 1 + 24 + 75 + 200$; $SO(10)$: $(45 \times 45)_{symm} = 1 + 54 + 210 + 770$

Representation	$M_3 : M_2 : M_1$ at M_{GUT}	$M_3 : M_2 : M_1$ at M_{EWSB}
1	1:1:1	6:2:1
24	2:(-3):(-1)	12:(-6):(-1)
75	1:3:(-5)	6:6:(-5)
200	1:2:10	6:4:10

String: $(3 + \delta_{GS}) : (-1 + \delta_{GS}) : \left(-\frac{33}{5} + \delta_{GS} \right)$ (OII, also mixed moduli anomaly)

Phenomenology

- Gaugino mass ratios

$$\frac{M_i(Q)}{M_{1/2}} = \eta_i \frac{\alpha_i(Q)}{\alpha_i(M_X)} \Rightarrow \begin{aligned} \frac{M_1(Q)}{M_2(Q)} &\approx 0.5\eta_1 \\ M_2(Q) &\approx 0.8M_{1/2} \\ \frac{M_3(Q)}{M_2(Q)} &\approx 2.7\eta_3 \end{aligned}$$

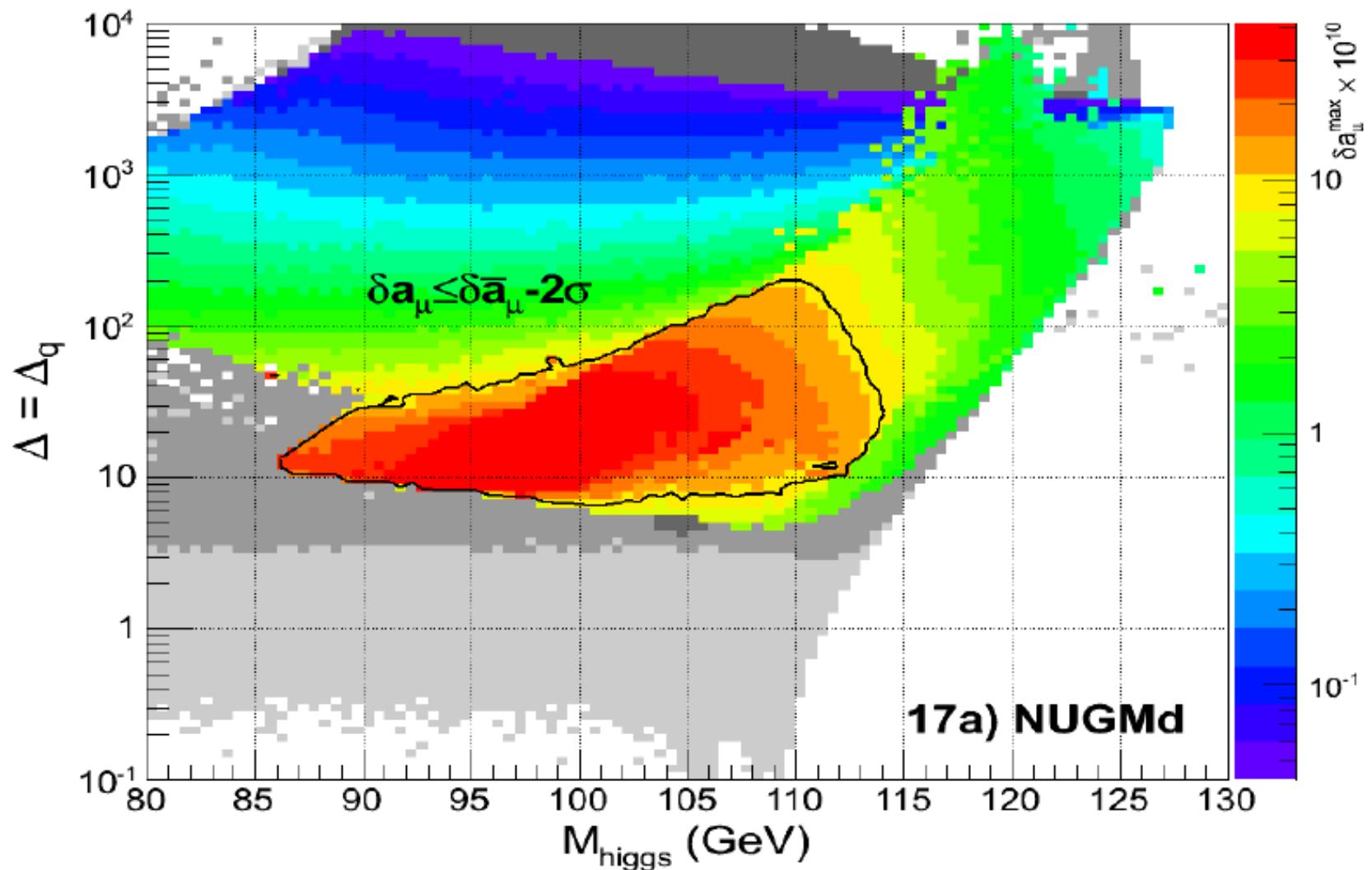
.... gauginos can be very heavy

- Light neutralino and 2 charginos nearly degenerate

$$m_{\chi_2^0} - m_{\chi_1^0} = M_Z^2 \left(\frac{s_W^2}{M_1} + \frac{c_W^2}{M_2} \right) + \mathcal{O}\left(\frac{M_Z^3}{M_2^2}\right)$$

$$m_{\chi_1^\pm} - m_{\chi_1^0} = \frac{1}{2} M_Z^2 \left(\frac{s_W^2}{M_1} + \frac{c_W^2}{M_2} \right) + \frac{1}{2} M_Z^2 \left(\frac{s_W^2}{M_1} - \frac{c_W^2}{M_2} \right) \epsilon \sin 2\beta + \mathcal{O}\left(\frac{M_Z^3}{M_2^2}\right)$$

+ for $|M_1| < \mu$, Bino or Higgsino LSP candidate



2-loop fine tuning in 75 case

Ghilencea, Lee, Park

Muon $g-2$

$$a_{\mu}^{theory} - a_{\mu}^{expt} = -(28.7 \pm 8.0) \times 10^{-10}$$

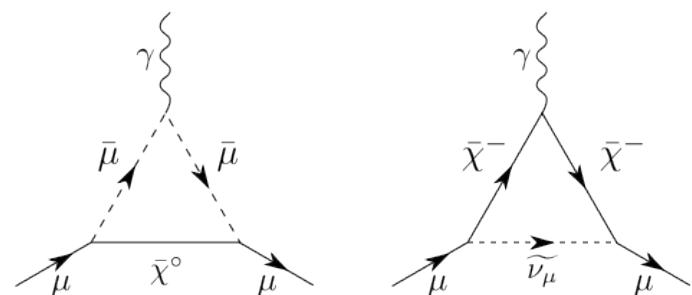
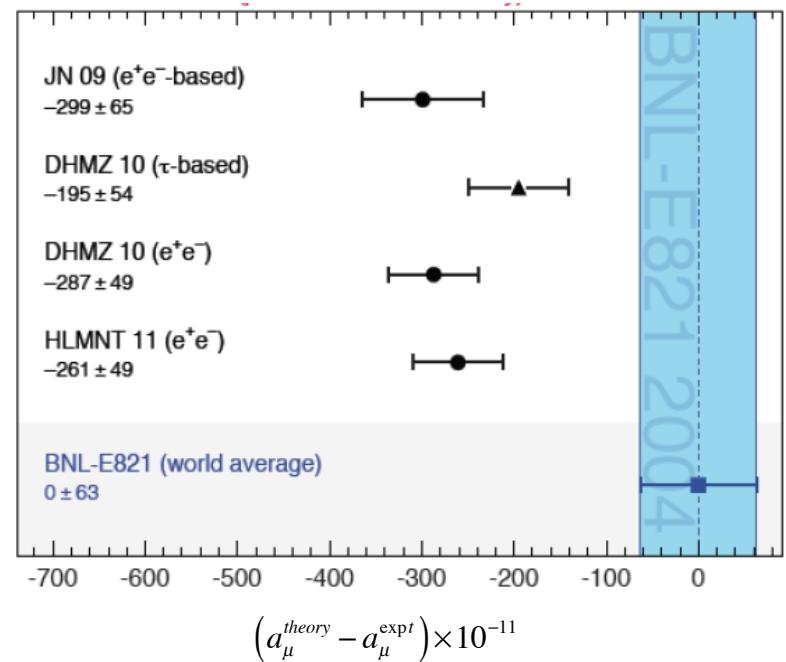
Theory error from hadronic contribution:

$$\delta a_{\mu}^{e^+e^-} = 3.6 \sigma$$

$$\delta a_{\mu}^{\tau} = 2.4 \sigma$$

SUSY

$$\delta a_{\mu}^{SUSY} = -13 \times 10^{-10} \left(\frac{100 \text{ GeV}}{M_{SUSY}} \right)^2 \tan \beta$$

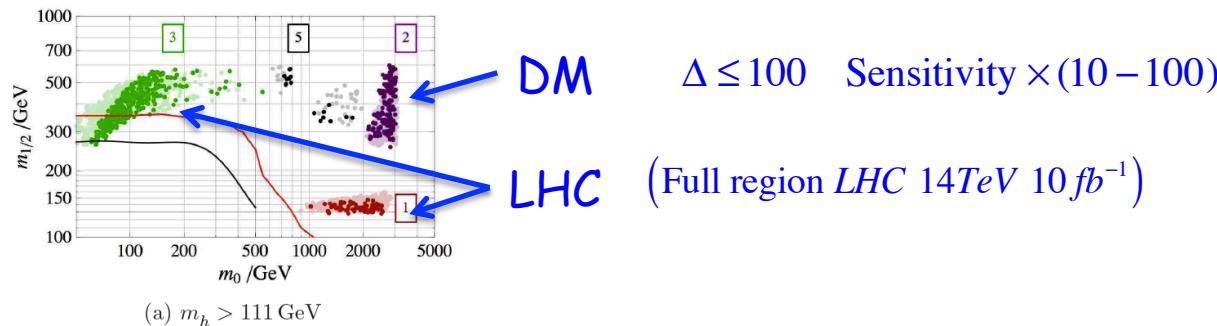


Needs light sleptons - anomaly/mirage spectrum?

Summary (before Higgs discovery)

- Hierarchy problem \Rightarrow SUSY breaking structure and/or further states
- CMSSM $m_i = M_0$ $\text{Max}[\Delta_{EW}, \Delta_\Omega] = 15(29)$, $m_h = 114(116) \pm 2\text{GeV}$

Complementary DM & LHC searches



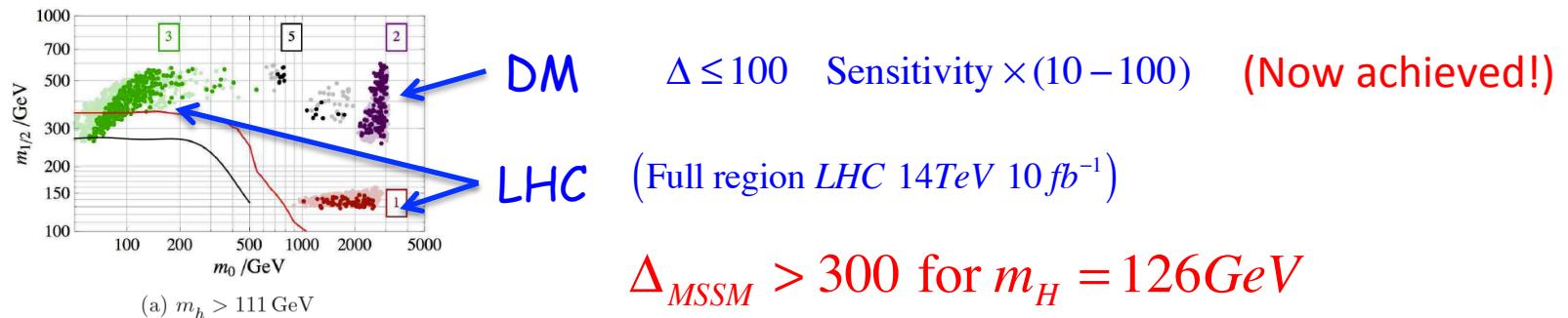
(Gauge mediation $\Delta \gg 100$)

- NMSSM Reduced $\Delta \Rightarrow$ GNMSSM $\Rightarrow Z_{4R}, Z_{8R}$
SUSY states can be (slightly) heavier
 $m_h \rightarrow 130\text{GeV}$
- Gaugino focus point $M_i = \eta_i M_{1/2}$ Characteristic η_i
Light $\chi^{0,\pm}$
 $\delta(b \rightarrow s\gamma)$ significant
 $\delta(g-2)$ Small(?)

Summary (after Higgs discovery)

- Hierarchy problem \Rightarrow SUSY breaking structure and/or further states
- CMSSM $m_i = M_0$ $\text{Max}[\Delta_{EW}, \Delta_\Omega] = 15(29)$, $m_h = 114(116) \pm 2\text{GeV}$

Complementary DM & LHC searches



- NMSSM Reduced $\Delta \Rightarrow$ GNMSSM $\Rightarrow Z_{4R}, Z_{8R}$
Further light/invisible Higgs a possibility

- Gaugino focus point
Mixed anomaly mediation, mirage mediation ...
Light sleptons...g-2!
- Natural SUSY, SPLIT SUSY...

