# Shift symmetries for the Higgs and the inflaton Arthur Hebecker (Heidelberg)

#### Outline

- The 125-GeV-Higgs (without SUSY) from a String-Pheno Perspective
  - $\rightarrow$  1204.2551 with **A. Knochel and T. Weigand**
- Main idea:  $\lambda = 0$  at some high scale (SUSY-breaking scale) due to shift symmetry in the Higgs sector
- Stringy origin of this shift symmetry
- <u>Closely related</u>: The very same symmetry may be reponsible for a flat potential in <u>fluxbrane inflation</u>

#### Motivation

We have a Higgs at 125 GeV and nothing else (yet?)

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Of course: low-scale SUSY is still OK Also: Muon-(g-2); h \to \gamma \gamma excess; 130-GeV \gamma-ray line...
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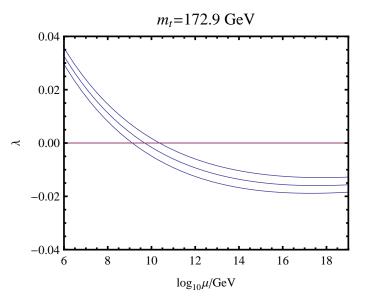
- <u>Nevertheless:</u> What if we just had to accept the fine-tuned non-SUSY SM for a large energy range?
- Well-known: for low  $m_h$ ,  $\lambda$  runs to zero at some scale  $< M_P$  (vacuum stability bound)

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Lindner, Sher, Zaglauer '89
Gogoladze, Okada, Shafi '07
...
Shaposhnikov, Wetterich 09'
Giudice, Isidori, Strumia, Riotto, ...
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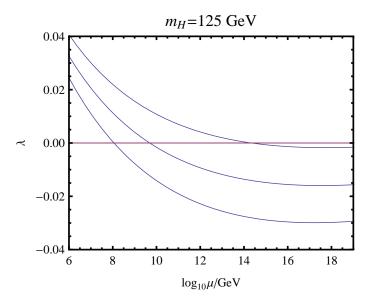
• It has been attempted to turn this into an  $m_h$  prediction



## Running of $\lambda$ (for a $\pm 1$ GeV variation of $m_{\rm Higgs}$ )



## Running of $\lambda$ (for a 2- $\sigma$ variation of $m_{\text{top}}$ )



### String-phenomenologist's perspective

- Insist on stringy UV completion (for conceptual reasons)
- Expect SUSY at string/compactification scale (stability!)
- Natural guess: The special scale  $\mu(\lambda=0)$  is the SUSY-breaking scale
- Crucial formula:

$$\lambda(m_s) = \frac{g^2(m_s) + g'^2(m_s)}{8} \cos^2(2\beta)$$

Reminder:

$$M_{H}^{2} = \left( \begin{array}{cc} |\mu|^{2} + m_{H_{d}}^{2} & b \\ b & |\mu|^{2} + m_{H_{u}}^{2} \end{array} \right) = \left( \begin{array}{cc} m_{1}^{2} & b \\ b & m_{2}^{2} \end{array} \right)$$

$$\sin(2\beta) = \frac{2m_3^2}{m_1^2 + m_2^2}$$
 Need this to be 1!

Of course, high-scale SUSY has been considered before

Arkani-Hamed, Dimopoulos '04 Giudice, Romanino '04

- Also, relations  $\tan \beta \leftrightarrow \lambda(m_s) \leftrightarrow m_h$  have been discussed cf. the 140-GeV-Higgs-mass-prediction of Hall/Nomura, '09
- Our goal: Identify as special structure/symmetry leading to  $\tan \beta = 1$  (i.e. to  $\lambda = 0$ )
- Indeed, such a structure is known in heterotic orbifolds:

Shift symmetry: 
$$K_H \sim |H_u + \overline{H}_d|^2$$

Lopes-Cardoso, Lüst, Mohaupt '94 Antoniadis, Gava, Narain, Taylor '94 Brignole, Ibanez, Munoz, Scheich, '95...'97

In more detail: 
$$K_H = f(S, \overline{S})|H_u + \overline{H}_d|^2$$

Assuming  $F_S \neq 0$  and  $m_{3/2} \neq 0$  this gives

$$m_1^2 = m_2^2 = m_3^2 = \left| m_{3/2} - \overline{F}^S f_{\overline{S}} \right|^2 + m_{3/2}^2 - F^S \overline{F}^S (\ln f)_{S\overline{S}}$$

 This shift-symmetric Higgs-Kähler potential has also been rediscovered/reused in orbifold GUTs

• In this language, it is easy to see the physical origin:

5d SU(6) 
$$\rightarrow$$
 SU(5)×U(1); 35 = 24+5+ $\overline{5}$ +1; Higgs=  $\Sigma + iA_5$  cf. Gogoladze, Okada, Shafi '07



#### Comments

 This simple understanding of the shift-symmetry lets us hope that it is more generic

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heterotic WLs \leftrightarrow type IIA / D6-WLs \leftrightarrow type IIB / D7-WLs or positions
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- These and other origins of the Higgs-shift-symmetry and of  $\tan \beta = 1$  have recently also been explored in Ibanez, Marchesano, Regalado, Valenzuela '1206...
- Clearly, we eventually need more phenomenological implications of 'stringy high-scale SUSY' (e.g. in cosmology)

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Chatzistavrakidis, Erfani, Nilles, Zavala '1206...
Higaki, Hamada, Takahashi '1206...
Anchordoqui, Goldberg, Huang, Lüst, Taylor, Vlcek '1208...
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#### Corrections? Precision?

- The superpotential (e.g. top Yukawa) breaks the shift symmetry
- The crucial point is compactification

Shift symmetry is exact (gauge symmetry!) in 10d. The shift corresponds to switching on a WL. This is not a symmetry in 4d (4d-zero modes 'feel' the WL). 4d-loops destroy the shift symmetry of Kähler potential.

 Optimistic approach to estimating the 'goodness' of our symmetry:

Symmetry-violating running between  $m_c$  and  $m_s$ 

 $\Rightarrow$  Correction  $\delta \sim \ln(m_c/m_s)$ 



#### More explicitly:

$$\begin{array}{lll} \mathit{M}_{H}^{2} & = & (|\mu|^{2} + \mathit{m}_{H}^{2}) \left(\begin{array}{cc} 1 & 1 \\ 1 & 1 \end{array}\right) + \left(\begin{array}{cc} \delta |\mu|^{2} + \delta \mathit{m}_{H_{d}}^{2} & \delta \mathit{b} \\ \delta \mathit{b} & \delta |\mu|^{2} + \delta \mathit{m}_{H_{u}}^{2} \end{array}\right) \\ & = & \mathsf{symmetric} & + & \mathsf{loop\ violation} \end{array}$$

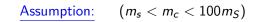
Leading effects: y<sub>t</sub> and gauge

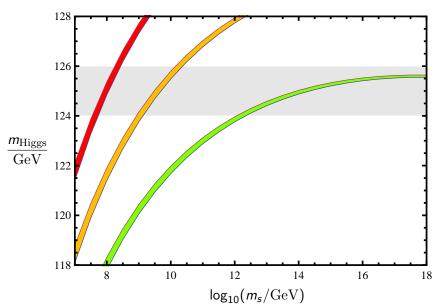
$$\delta M_H^2 = f(\epsilon_y, \epsilon_g, m_{\rm soft})$$
 ;  $\epsilon_y = \int_{\ln m_s}^{m_c} dt \, \frac{6|y_t|^2}{16\pi^2}$ 

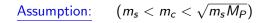
• Enforce det  $M_H^2=0$  after corrections  $\Rightarrow \epsilon_y, \epsilon_g, m_{\rm soft}$  are related

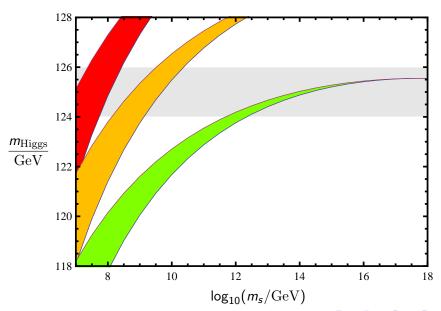
$$\cos 2\beta = \epsilon_{V} \times \{\text{calculable } \mathcal{O}(1) \text{ factor}\}\$$







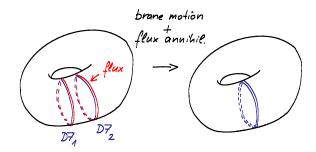




## A different application of the same shift symmetry

AH, Kraus, Lüst, Steinfurt, Weigand, 1104.5016 ..., Küntzler, 1207.2766 ..., Arends, Heimpel, Mayrhofer, Schick, 12...

 Fluxbrane inflation with flat direction protected by shift symmetry for D7-brane motion



Related to WLs by mirror symmetry / T-duality



#### Fluxbrane inflation

• Crucial fact: At large volume (i.e. weak flux *F*), the potential is much more flat than in brane-antibrane inflation:

$$V \sim 1 - rac{g_s}{r^{d_\perp - 2}} \quad o \quad V \sim extbf{F}^2 - extbf{F}^4 rac{g_s}{r^{d_\perp - 2}}$$

Hence:  $\eta \sim F^2 \ll 1$ 

• Note: This is conceptually similar to D3/D7 inflation

Dasgupta, Herdeiro, Hirano, Kallosh, '02

and T-dual to inflation from branes at angles and Wilson lines

Garcia-Bellido, Rabadan, Zamora, '01 Avgoustidis, Cremades, Quevedo, '06



## Flat direction / shift symmetry

- Chose brane/bulk fluxes such that  $W_0$  does not depend on  $\varphi$ .
- Of course, since  $W_0 \not\equiv 0$ , the usual ' $\eta$ -problem of supergravity' is still present:

$$\mathcal{K} = -\ln(S + \overline{S} + \kappa(\varphi, \overline{\varphi})) + \cdots \implies \eta \simeq 1 \text{ from } V_{\mathcal{F}}$$

[Here  $\kappa$  is the Kähler potential on the D7-brane moduli space; similar to situation in KKLMMT.]

- Fact: F-theory on K3×K3 has  $\kappa = \kappa(\varphi + \overline{\varphi})$
- We expect this shift-symmetric structure to arise more generally in the large complex structure limit.

Grimm, Ha, Klemm, Klevers, ... '09-'11 Alim, Hecht, Jockers, Mayr, Mertens, ...



## Conclusions / Summary

- In the absence of new electroweak physics at a TeV, the 'vacuum stability scale'  $(\lambda(\mu)=0)$  may be a crucial hint at new physics
- Well-motivated guess: SUSY broken with  $an \beta = 1$  at this scale
- Possible structural reason: shift symmetry in Higgs sector
   (Predictivity, i.e. m<sub>h</sub> + m<sub>t</sub> + α<sub>s</sub> ⇒ m<sub>s</sub> remains strong, even if shift symmetry is only approximate)
- The very same stringy symmetry (but in a different sector) may be crucial to maintain flatness in Fluxbrane inflation