

3D non-relativistic supergravity

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Motivation

- *Fermions* are ubiquitous in description of the 'real' world; conceptually they are nicely incorporated via *supersymmetry*
- Most physics (for all practical purposes) is best described *non-relativistically*, e.g. *Newtonian gravity*

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- we want to study non-relativistic supergravity/-particles

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A simplification:

- to make our task easier we start with lower-dimensional 3D toy models

Diagrammatic motivation

General relativity

non - rel \longrightarrow

Newtonian gravity

SUSY



Supergravity

non - rel \longrightarrow

SUSY



??

Outline

- Motivation
- Our setup:
 - Gauge the super-Bargmann algebra
 - Impose constraints to obtain
Newton–Cartan super-gravity [Andringa, Bergshoeff, Rosseel, Sezgin '13]
 - Fix Stückelberg symmetries to obtain supergravity for the Galilean observer
- Results obtained so far
- Future directions

The setup

- Gauge the $\mathcal{N} = 2$ super-Bargmann algebra.
- Impose constraints, e.g. to convert time- and space-translations into general coordinate transformations
- At this point we have *Newton–Cartan supergravity*
- The constraints turn some fields into pure Stückelberg fields that can be gauged away
- Doing so will lead to supergravity for the Galilean observer, thus to a supersymmetric version of Newtonian gravity

Gauging the $\mathcal{N} = 2$ super-Bargmann algebra

- Bargmann algebra in 3D:

$$\left(H, P_a, G_a, J_{ab}, Z, Q_\alpha^+, Q_\alpha^- \right)$$

$$\begin{aligned} \text{Why } \mathcal{N} = 2 : \quad \{Q_\alpha^+, Q_\beta^+\} &= 2\delta_{\alpha\beta}H, \\ \{Q_\alpha^+, Q_\beta^-\} &= -[\gamma^{a0}]_{\alpha\beta}P_a. \end{aligned}$$

- Gauged algebra: for each generator (e.g. Z) we have associated
 - gauge fields (m_μ),
 - gauge parameters ($\delta m_\mu = \partial_\mu \sigma + \lambda^a e_{\mu a} + \text{fermionic}$) and
 - curvatures ($\hat{R}_{\mu\nu}(Z) = 2\partial_{[\mu} m_{\nu]} - 2\omega_{[\mu}{}^a e_{\nu]a} + \text{fermionic}$).

From Bargmann to Newton–Cartan to Galileo

- Gauged algebra: gauge fields, gauge parameters and curvatures.
- Impose constraints: $\hat{R}_{\mu\nu}{}^a(P) = 0$, $\hat{R}_{\mu\nu}{}^{ab}(J) = 0$, $\hat{R}_{\mu\nu}(H) = 0$, $\hat{R}_{\mu\nu}(Z) = 0$ and $\hat{\psi}_{\mu\nu+} = 0$.
- This allows us to convert H - and P^a -transformations into general coordinate transformations.
- Setting some of the gauge fields to zero, while fixing constant values for others we can remove all residual gauge freedom.
- The only fields left are the Newton force $\Phi_i = \partial_i\Phi = \partial_i m_\emptyset$ and its supersymmetric partner $\Psi = \psi_{\emptyset-}$.
- It is not possible to write down a closed algebra in terms of Φ and Ψ !

Results obtained so far

- We know the supersymmetric partner of the Newton force Φ_i .
- We found a representation of the $\mathcal{N} = 2$ super-Bargmann algebra realized on the “Newton potentials” Φ_i and Ψ :

$$\delta\Phi_i = \dots - \lambda^m_n x^n \partial_m \Phi_i + \bar{\epsilon}_-(t) \gamma^0 \partial_i \Psi + \frac{1}{2} \bar{\epsilon}_+ \gamma_i \Psi$$

$$\delta\Psi = \dots - \lambda^i_j x^j \partial_i \Psi + \frac{1}{4} \lambda^{ab} \gamma_{ab} \Psi + \dot{\epsilon}(t) - \frac{1}{2} \Phi^k \gamma_{k0} \epsilon_+$$

- We can write the algebra in terms of Φ thereby introducing the dual potentials χ and Ξ : $\Psi = (1/2) \gamma^i \partial_i \chi$ and $\partial_i \Phi = \epsilon_{ij} \partial^j \Xi$.
- Ξ is a bosonic dual to Φ that is needed to close the algebra.

Future directions

- Repeat the analysis in 4D, i.e. find the supersymmetric partner(s) of Φ at the level of the algebra.
- Find other 'multiplets' in 3D.
- Find an action for a (3D) superparticle:

$$I = \frac{m}{2} \int dt \{ \dot{x}^i \dot{x}^i + 2\Phi + \text{SUSY extension} \}$$

Summary

GR

non - rel \longrightarrow

Newtonian gravity, Φ

SUSY



SUGRA

non - rel \longrightarrow

Φ, χ, Ξ (algebra) *particle?*

SUSY

