

Introduction to Chern-Simons forms in Physics - II

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Lecture I:

1. Topological invariants
2. Minimal coupling
3. Chern-Simons
4. Brane couplings
5. Quantization
6. Summary

Lecture II:

1. Spacetime geometry
2. Metric and affine structures: e^a , ω^{ab}
3. Building blocks
4. Gravity actions: L-L, T, CS
5. CS gravities, BHs, SUGRAs
6. Summary

The Euler characteristic belongs to a family of famous invariants:

- ◆ Sum of exterior angles of a polygon
- ◆ Residue theorem in complex analysis
- ◆ Winding number of a map
- ◆ Poincaré-Hopf theorem (“*one cannot comb a sphere*”)
- ◆ Atiyah-Singer index theorem
- ◆ Witten index
- ◆ Dirac’s monopole quantization
- ◆ Aharonov-Bohm effect
- ◆ Gauss’ law
- ◆ Bohr-Sommerfeld quantization
- ◆ Soliton/Instanton topologically conserved charges
- ◆ etc...

All of these examples involve topological invariants called the Chern characteristic classes and Chern-Simons forms.

Chern-Simons forms

$$C_{2n+1} = \left\langle A \wedge (dA)^n + \alpha_1 A^3 \wedge (dA)^{n-1} + \cdots + \alpha_n A^{2n+1} \right\rangle \equiv \left\langle \tilde{C}_{2n+1} \right\rangle$$

- ◆ *No dimensionful constants: $\alpha_k =$ fixed rational numbers*
- ◆ *No adjustable coefficients: α_k cannot get renormalized*
- ◆ *No metric needed; scale invariant* ←
- ◆ *Entirely determined by the Lie algebra and the dimension*
- ◆ *Only defined in odd dimensions*
- ◆ *Unique gauge quasi-invariants: $\delta C_{2p+1} = d\Omega_{2p}$* ←
- ◆ *Related to the Chern characteristic classes*

Is “CS gravity” an oxymoron?

What gauge symmetry is this?

How are the α 's found?

Chern-Simons recipe

- Lie algebra \mathbf{G} , generators J_a
- Invariant tensor $\tau_{a_1 \cdots a_n} = \langle J_{a_1} \cdots J_{a_n} \rangle$
- \mathbf{G} -valued connection 1-form A , $F = dA + A^2$
- Invariant polynomial (characteristic class)

$$P_{2n}[A] = \frac{1}{n} \langle (F \wedge)^n \rangle = \frac{1}{n} \langle F \wedge \cdots \wedge F \rangle$$

- Odd-dimensional manifold M^{2n-1}

CS action: $I[A] = \kappa \int_{\mathcal{M}^{2n-1}} C_{2n-1}(A)$

where $dC_{2n-1}(A) = P_{2n}[A]$

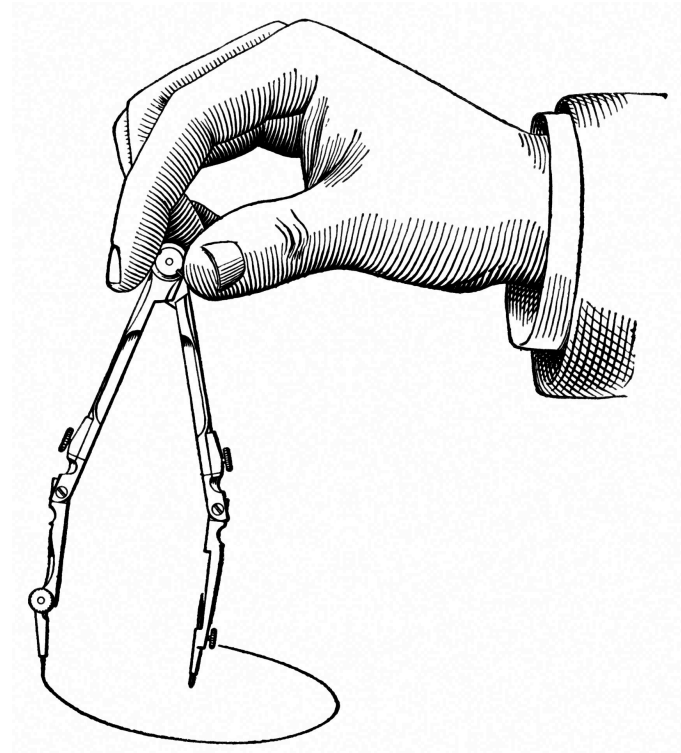
$$L(e, \omega) = (n+1) \int_0^1 dt \langle A \wedge F_t^n \rangle,$$

$$F_t = dA_t + A_t^2, \quad A_t \equiv tA$$

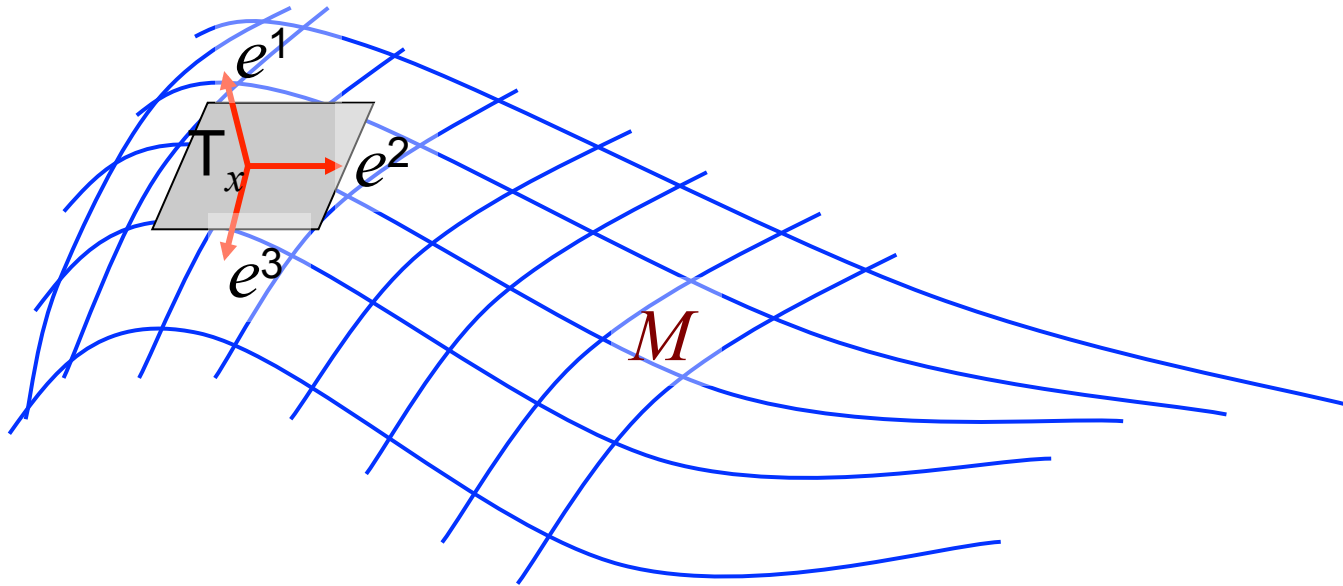
1. Spacetime geometry

Geometry has two ingredients:

- Metric structure (length/area/volume, scale) $\longrightarrow e^a$
- Affine structure (parallel transport, congruence) $\longrightarrow \omega^a_b$



- Spacetime M is a D -dimensional differentiable manifold (up to isolated singularities in sets of zero-measure).
- M admits a D -dimensional tangent space T_x at each point.



- An open set around any point x is diffeomorphic to an open set in the tangent.

2. Metric & affine structures

Metric Structure

- Each tangent space \mathbb{T}_x is a copy of Minkowski space.
- Since Minkowski space is endowed with the Lorentzian metric, the diffeomorphism induces a metric structure on M :

$$dz^a = e^a_{\mu}(x) dx^{\mu} \equiv e^a$$

(e^a : vielbein, soldering form, local orthonormal frame).

$$\begin{aligned} ds^2 &= \eta_{ab} dz^a dz^b \\ &= \eta_{ab} e^a_{\mu} e^b_{\nu} dx^{\mu} dx^{\nu} \equiv g_{\mu\nu} dx^{\mu} dx^{\nu} \end{aligned}$$

Metric on M :

$$g_{\mu\nu}(x) = \eta_{ab} e^a_{\mu}(x) e^b_{\nu}(x)$$

- The vielbein can be viewed as the Jacobian matrix that relates differential forms in \mathbb{T}_x and M .

$$dz^a = e_\mu^a(x) dx^\mu$$

- This is a mapping between differentials in two distinct spaces.
- It can also be viewed as the operator that relate tensors in each space, for example, the metric:

$$g_{\mu\nu}(x) = \eta_{ab} e_\mu^a(x) e_\nu^b(x)$$



- The metric is a *composite*
- e_μ^a is the square root of the metric, $e_\mu^a \sim \sqrt{g_{\mu\nu}}$.

Equivalence Principle

A sufficiently small vicinity of any point of M can be accurately approximated by T_x .

- In a sufficiently small region of spacetime a reference frame can always be found in which the laws of physics are those of special relativity (**free fall**).*
- The laws of physics are invariant under local Lorentz transformations.*
- General Relativity is a nonabelian gauge theory for the group $SO(3,1)$ in 4D. (**40 years before Yang & Mills!**)*

Affine Structure

- *The differential operations appropriate for the theory should be Lorentz covariant, e.g.,*

$$D_{\mu}u^a = \partial_{\mu}u^a + \omega^a{}_{b\mu}u^b$$

where u^a is a Lorentz vector and $\omega^a{}_{b\mu}$ is the connection for the Lorentz group.

- *This means that under a Lorentz transformations, u^a and $D_{\mu}u^a$ must transform in the same manner:*

$$u^a \rightarrow u'^a = \Lambda^a{}_b u^b, \Lambda^a{}_b \in SO(D-1,1)$$

$$Du^a \rightarrow (Du^a)' = \Lambda^a{}_b (Du^b)$$

- *Geometric interpretation*

$u^a(x)$: *vector field at x*

$u^a(x + dx)$: *vector field at $x + dx$*

$$\begin{aligned} u^a_{\parallel}(x + dx) &= [\delta^a_b + \omega^a_{b\mu} dx^\mu] u^b(x + dx) \\ &= u^a(x + dx) + \omega^a_{b\mu} dx^\mu u^b(x): \textit{u}^a \textit{ parallel-} \\ &\quad \textit{transported from } x + dx \textit{ to } x \end{aligned}$$

Hence, $u^a_{\parallel}(x + dx) - u^a(x)$

$$= u^a(x + dx) + \omega^a_{b\mu} dx^\mu u^b - u^a$$

$$= dx^\mu [\partial_\mu u^a + \omega^a_{b\mu} u^b] = dx^\mu D_\mu u^a$$

- *The covariant derivative and the curvature 2-form are defined uniquely by the Lorentz connection 1-form*

$$\omega^a = \omega^a_{b\mu} dx^\mu$$

- *Covariant derivative:*

$$Du^a = du^a + \omega^a_b u^b$$

- *Curvature:*

$$DDu^a = R^a_b u^b, \quad R^a_b = d\omega^a_b + \omega^a_c \omega^c_b$$

- *Bianchi identity:*

$$DR^a_b u^b = R^a_b Du^b$$

where $DR^a_b = dR^a_b + \omega^a_c R^c_b - \omega^a_c R^c_b \equiv 0$

3. Building Blocks

- *So far, we have defined the fundamental elements to construct a dynamical theory of the spacetime geometry:*

e^a	<i>vector, 1-form</i>
ω^a_b	<i>connection 1-form</i>
$R^a_b = d\omega^a_b + \omega^a_c \omega^c_b$	<i>curvature 2-form</i>
$T^a = de^a + \omega^a_c e^c$	<i>torsion 2-form</i>
$\varepsilon_{a_1 a_2 \dots a_D}, \eta_{ab}$	<i>invariant tensors 0-forms</i>

- *By taking successive derivatives no new building blocks are produced:*

$$DR^a_b \equiv 0, \quad DT^a = R^a_b e^b, \quad dd = 0.$$

- *The most general D -dimensional Lorentz invariant Lagrangian built with these ingredients is a polynomial in e^a , ω^a_b , and their exterior derivatives.*
- *We assume nothing about the invertibility of the vielbein, or the metric, and **do not include the *Hodge**.*
- *The use of exterior derivatives guarantees that all field equations will be first order. If the torsion is set to zero, the equations for the metric will be second-order at most.*
- *This recipe gives all the Lovelock series and a lot more.*

If no further ingredients or assumptions are added, there is a finite family of Lagrangians that can be constructed with these elements in each dimension.

4. Gravity actions

D - dimensional gravity

Local frames and parallelism are the essential features of geometry.

- They are logically independent.
- They should be dynamically independent.

Dynamical fields

$$\text{Vielbein: } e^a = e^a_\mu dx^\mu$$

$$\text{Lorentz (spin) connection: } \omega^a_b(x) = \omega^a_{b\mu}(x) dx^\mu.$$

Local symmetry:

Local Lorentz transformations,

$$e^a(x) \rightarrow e'^a(x) = \Lambda^a_b(x) e^b(x)$$

$$\omega^a_b(x) \rightarrow \omega'^a_b(x) = (\Lambda^{-1})^a_c \left[\omega^c_d(x) \Lambda^d_b + d\Lambda^c_b \right]$$

$$\Lambda^a_b(x) \in SO(D-1,1)$$

Manifestly Lorentz invariant theory

Coordinate independence is achieved naturally using coordinate-free one-form fields e^a , ω^a_b .

The Lagrangian for D -dimensional gravity is assumed to be:

- A D -form constructed out of e^a , ω^a_b , their exterior derivatives, and the Lorentz invariant tensors η_{ab} and $\varepsilon_{a_1 a_2 \dots a_D}$,

$$L = L(e, de, \omega^a_b, d\omega^a_b)$$

- (Quasi*-) invariant under local Lorentz transformations, in order to ensure Lorentz covariance of the field equations.

(* Up to total derivatives)

Using only exterior derivatives limits the number of options:
 Taking exterior derivatives of e and ω , two tensors are found

$$\begin{aligned}
 De^a &= de^a + \omega^a_b \wedge e^b = \boxed{T^a} \quad \text{Torsion} \\
 d\omega^a_b + \omega^a_c \wedge \omega^c_b &= \boxed{R^a_b} \quad \text{Curvature}
 \end{aligned}
 \left. \vphantom{\begin{aligned} De^a \\ d\omega^a_b \end{aligned}} \right\} \underline{\text{2-forms}}$$

Bianchi identities

$$DR^a_b \equiv 0$$

$$DT^a \equiv R^a_b \wedge e^b$$

This limits the number of ingredients in L to:

$$\underbrace{\eta_{ab}, \varepsilon_{a_1 a_2 \dots a_D}}_{\text{Zero-forms}}, \quad \underbrace{e^a, \omega^a_b}_{\text{one-forms}}, \quad \underbrace{T^a, R^a_b}_{\text{two-forms}}$$

Three series are found:

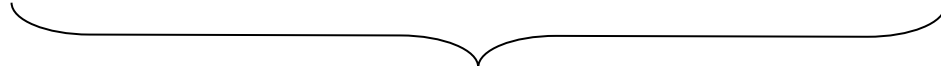
A. Lovelock

B. Torsion

C. Lorentz C-S



Euler classes



(Exotic gravities)



Pontryagin classes

A: Generalization of E-H for all D , parity even.

B: Some D only, vanish for $T^a=0$ or $\Lambda=0$, parity odd.

C: For $D=4k-1$ only, parity odd.

$$L(e, \omega) = L_{Lovelock} + L^T + L_{Lorentz}^{CS}$$

This series includes:

$$L_0 = \varepsilon_{a_1 \dots a_D} e^{a_1} \dots e^{a_D}$$

Cosmological constant term (volume)

$$L_1 = \varepsilon_{a_1 \dots a_D} R^{a_1 a_2} e^{a_3} \dots e^{a_D}$$

Einstein-Hilbert

$$L_2 = \varepsilon_{a_1 \dots a_D} R^{a_1 a_2} R^{a_3 a_4} e^{a_5} \dots e^{a_D}$$

Gauss-Bonnet

.

.

.

$$L_{D-1} = \varepsilon_{a_1 \dots a_D} R^{a_1 a_2} \dots R^{a_{D-2} a_{D-1}} e^{a_D}$$

Poincaré invariant term
(odd D) [see below]

$$L_D = \varepsilon_{a_1 \dots a_D} R^{a_1 a_2} \dots R^{a_{D-2} a_D} = \mathbf{E}_D$$

Euler density
(even D) [see above]

B. Torsion series

If $T^a \neq \mathbf{0}$, more invariant terms are allowed:

$$t_p = T_{a_1} R^{a_1}_{a_2} R^{a_2}_{a_3} \cdots R^{a_{2p-1}}_{a_{2p}} T^{a_{2p}}, \quad [p = \text{even}]$$

$$e_p = e_{a_1} R^{a_1}_{a_2} R^{a_2}_{a_3} \cdots R^{a_{2p-1}}_{a_{2p}} e^{a_{2p}}, \quad [p = \text{odd}]$$

$$k_p = T_{a_1} R^{a_1}_{a_2} R^{a_2}_{a_3} \cdots R^{a_{2p-1}}_{a_{2p}} e^{a_{2p}}$$

..and products thereof.

These traces are related to the Pontryagin form (Chern class),

$$P_{4k} = R^{a_1}_{a_2} R^{a_2}_{a_3} \cdots R^{a_{2k}}_{a_1} = \text{Tr}(R^{2k}),$$

whose integral is a topological invariant in $4k$ dimensions.

C. Lorentz Chern-Simon series

These are quasi-invariant terms involving ω^{ab} explicitly:

$$C_3 = d\omega^a_b \omega^b_a + \frac{2}{3} \omega^a_b \omega^b_c \omega^c_a,$$

Fixed numbers

and in

$$C_{4n-1}^{\text{general}} = \left(d\omega^a_b\right)^{2n-1} \omega^b_a + \cdots + \gamma_{4n-1} \left(\omega^b_a\right)^{4n-1}$$

- They are also related to the Pontryagin forms:

$$dC_{4n-1} = P_{4n}$$

- They do not vanish if $T^a \equiv \mathbf{0}$ or if $\Lambda = \mathbf{0}$
- Products of these and torsional terms are also acceptable Lagrangians, provided they are D -forms .

The Action

The most general lagrangian for gravity in D dimensions is

$$L_G[e, w] = \int_{M_D} L_{Lovelock}^D + L_{Exotic}^D ,$$

$$\text{where } L_{Lovelock}^D = \sum_{p=0}^{[D/2]} a_p L_p$$

$$\text{with } L_p = e_{a_1 \dots a_D} R^{a_1 a_2} \dots R^{a_{2p-1} a_{2p}} e^{a_{2p+1}} \dots e^{a_D} ,$$

Arbitrary
coefficients!

The exotic part contains torsional and Lorentz C-S terms,
with more arbitrary constant coefficients,

$$\beta_1, \beta_2, \dots, \beta_s .$$

Features

General action for a D - dimensional geometry.

- 1st order field equations for e, ω .

- Torsion-free sector :

$$T^a = de^a + \omega^a_b e^b = 0 \Rightarrow \omega^{ab} = \omega^{ab}(e, \partial e)$$



- Second order theory for $g_{\mu\nu}$
- Same degrees of freedom as ordinary gravity

- Diffeomorphic invariant by construction

- Invariant under local Lorentz transformations

- For $D=4$ Einstein-Hilbert is the only option

Puzzles / problems

- The action has a number of cosmological constants

$$\lambda_i = \lambda_i(\alpha_p, \beta_q), \quad i = 1, 2, \dots, [(D-1)/2]$$

whose values *are not fixed* (the cosmological constant problem raised to a large number)

- Violently different behavior for different choices of α_p, β_q .
- Hopeless quantum scenario: α_p, β_q dimensionful and unprotected from renormalization.
- How could this be so if gravity descends from a fundamental renormalizable or finite theory?
- **Dynamics:** Non-uniform phase space (??)

At the point in parameter space where all the cosmological constants coincide, a miracle occurs

- All the dimensionful coupling constants can be absorbed by a redefinition of the vielbein

$$e^a \rightarrow \frac{e^a}{l}$$

- The vielbein and the spin connection can be combined as components of a single connection 1-form,

$$A = l^{-1} e^a J_a + \frac{1}{2} \omega^{ab} J_{ab}$$

where $J_{ab} =$ Lorentz generator
 $J_a =$ generators of (A)dS boosts/translations

} (A)dS/Poincaré algebra

- In odd dimensions, the action describes a gauge theory for the corresponding group. *Symmetry enhancement and no dimensionful coupling constants!*



The new curvature is

$$F^{AB} = dW^{AB} + W^A{}_C W^{CB} = \begin{bmatrix} R^{ab} \oplus l^{-2} e^a e^b & l^{-1} T^a \\ -l^{-1} T^b & 0 \end{bmatrix} \begin{array}{l} + \text{AdS} \\ - \text{dS} \end{array}$$

The Euler (\mathbf{E}_{2n}) and Pontryagin (\mathbf{P}_{4k}) invariants can be constructed with F^{AB} . These invariants are closed:

$$d\mathbf{E}_{2n} = 0 \implies \mathbf{E}_{2n} = d\mathbf{C}^E_{2n-1}$$

$$d\mathbf{P}_{4k} = 0 \implies \mathbf{P}_{4k} = d\mathbf{C}^P_{4k-1}$$

- Hence, they can be (locally) written as the exterior derivative of “something else”.

\mathbf{C}^E_{2n-1} , \mathbf{C}^P_{4k-1} are functions of e , ω^{ab} and their derivatives.
 \mathbf{E}_{2n} , \mathbf{P}_{4k} invariant under (A)dS, \implies so are the C’s.

These are the combinations we’re looking for!!

4.a Special choices (*Lovelock series*)

For $D = 2n+1$, choosing the coefficients in the Lovelock series as

$$\alpha_p = \alpha_p^{AdS} = \frac{l^{2p-D}}{D-2p} \binom{\frac{D-1}{2}}{p},$$

produces the particular form of the Lovelock lagrangian

$$L_{AdS} = l^{-1} R^n e + \frac{l^{-3}}{3} n R^{n-1} e^3 + \cdots + \frac{l^{-(1+2n)}}{2n+1} e^{2n+1},$$

with an enhanced symmetry, Lorentz \longrightarrow AdS(*).

The secret of the enhancement:

$SO(D-1, 2)$ invariant

$$dL_{AdS} = \mathbf{E}_{2n+2}$$

(*) Alternate signs for dS.

In the limit $l \rightarrow \infty$, only one constant survives

$$\alpha_p = \begin{cases} 1 & \text{if } p = \frac{D-1}{2} \\ 0 & \text{otherwise} \end{cases}$$

The result is a Poincaré-invariant theory,

$$L_{Poincaré} = \varepsilon_{a_1 \dots a_D} R^{a_1 a_2} \dots R^{a_{D-2} a_{D-1}} e^{a_D}$$

with the enhancement $SO(D-1,1) \rightarrow ISO(D-1,1)$

N.B. The Einstein-Hilbert action,

$$L_{E-H} = \varepsilon_{a_1 \dots a_D} R^{a_1 a_2} e^{a_3} \dots e^{a_D},$$

is not Poincaré - invariant.

4b. Special choices (*torsion series*)

In $D=4k-1$, there is a particular combination of the torsion series invariant under (A)dS, related to the Pontryagin class.

$$dL_{AdS}^* = F^{A_1}_{A_2} F^{A_2}_{A_3} \cdots F^{A_{2k}}_{A_1}$$

$$= P_{4k} \text{ Pontryagin invariant of } 4k \text{ dimensions}$$

(The explicit form of L_{AdS}^* is not very interesting and it involves torsion explicitly.)

- In the limit $l \rightarrow \infty$, the only surviving term in L_{AdS}^* is the Lorentz Chern-Simons term

$$L^L_{4n-1} = \left(d\omega^a_b \right)^{2n-1} \omega^b_a + \cdots + c_{4n-1} \left(\omega^b_a \right)^{4n-1}$$

5. CS actions

AdS Theory as Chern-Simons

The AdS invariant theory can be expressed in more standard form in terms of its Lie algebra valued connection and curvature:

$$A = l^{-1} e^a J_a + \frac{1}{2} \omega^{ab} J_{ab}$$

$$F = dA + A^2$$

$$= l^{-1} T^a J_a + \frac{1}{2} (R^{ab} + l^{-2} e^a e^b) J_{ab}$$

Both the Euler and Pontryagin Chern-Simons can be written as

$$dL_{AdS}^{(*)} = \langle \mathbf{F}^n \rangle$$

Appropriate multilinear invariant product (trace)

Dynamics (*generic C-S case*)

For $D \leq 3$ all CS theories are topological: No propagating degrees of freedom.

For $D = 2n+1 \geq 5$, a CS theory for a gauge algebra with N generators can as many as $\Delta = Nn - N - n$ local d. of f. (Bañados, Garay, Henneaux)

Phase space is not homogeneous:

$$\Delta = \begin{cases} Nn - N - n & \text{generic configurations (most states, no symmetry)} \\ 0 & \text{maximally degenerate (measure zero, max. symmetric)} \end{cases}$$

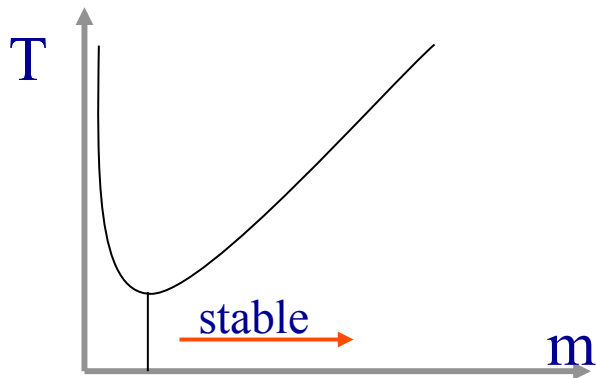
Degeneracy: The system may evolve from an initial state with Δ_1 d.o.f., reaching a state with $\Delta_2 < \Delta_1$ d.o.f. in a finite time. This is an irreversible process in which there is no record of the initial conditions in the final state.

Dynamics of C-S Gravity

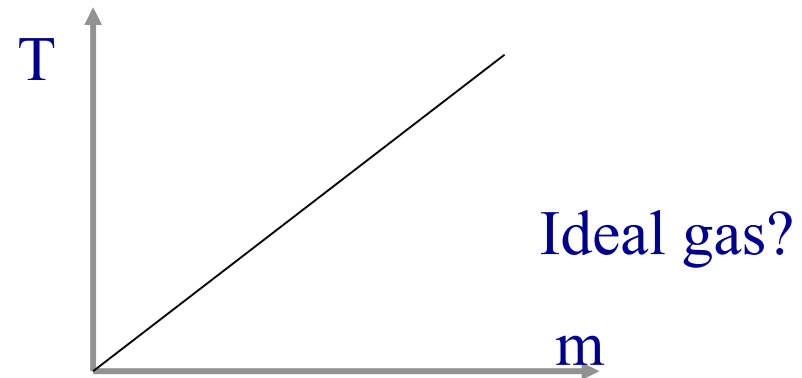
Around a flat background, $T^a = 0$, $R^{ab} = 0$, Chern-Simons theories behave as the standard Einstein-Hilbert theory. (same d.o. f., propagators, etc.)

- But, D -dimensional Minkowski space is not a classical solution!!
- The natural vacuum is, D -dimensional AdS space (max. symmetric).
- But around D -dimensional AdS space the theory has no d.o.f.!!
- Black holes are particularly sensitive to the theory:

Generic Lovelock theory



Chern-Simons theory



Supersymmetry

Supersymmetry is the only nontrivial extension of the spacetime symmetry [Lorentz, Poincaré, (A)dS].

New ingredient: Gravitino $\psi = \psi_\mu dx^\mu$

This is also a 1-form, which suggests a new connection,

$$\mathbf{A} = l^{-1} e^a \mathbf{J}_a + \frac{1}{2} \omega^{ab} \mathbf{J}_{ab} + \frac{1}{2} (\bar{\psi} \mathbf{Q} + \bar{\mathbf{Q}} \psi) + b \cdot \mathbf{X}$$

Necessary to close the algebra.
Depends on the dimension.

The generators $\{\mathbf{J}_a, \mathbf{J}_{ab}, \mathbf{Q}, \bar{\mathbf{Q}}, \mathbf{X}\}$ should form a (super) algebra.

Supersymmetry transformations

$$\delta e^a = \bar{\varepsilon} \Gamma^a \psi$$

$$\delta \omega^{ab} = \bar{\varepsilon} \Gamma^{ab} \psi$$

$$\delta \psi = \nabla \varepsilon = \left(d + e^a \mathbf{J}_a + \frac{1}{2} \omega^{ab} \mathbf{J}_{ab} \right) \varepsilon$$



$$\delta L(e, \omega) = \underbrace{\delta L_F(e, \omega, \psi, \bar{\psi})}_{\text{Tentative fermionic Lagrangian}}$$



Tentative fermionic Lagrangian

$$\delta L_F(e, \omega, \psi, \bar{\psi}) = \delta L_B^{New}(e, \omega, b^{[a_1 a_2 \dots]}) \dots \text{ etc.}$$

... this eventually yields a supersymmetric Lagrangian.

The construction is straightforward and the resulting theories are Chern-Simons systems . The field content and the superalgebras for the first cases are

D	Algebra	Field content	Standard Supergravity
5	$su(2,2 1)$	$e_\mu^a, \omega_\mu^{ab}, A_\mu, \psi_\mu^\alpha, \bar{\psi}_{\alpha\mu}$	$g_{\mu\nu}, A_\mu, \psi_\mu^\alpha, \bar{\psi}_{\alpha\mu}$
7	$osp(2 8)$	$e_\mu^a, \omega_\mu^{ab}, A_{j\mu}^i, \psi_\mu^{\alpha i} (i, j = 1,2)$	$g_{\mu\nu}, A_{[3]}, A_{j\mu}^i, \lambda^\alpha, \phi, \psi_\mu^{\alpha i}$
11	$osp(1 32)$	$e_\mu^a, \omega_\mu^{ab}, b_\mu^{abcde}, \psi_\mu^\alpha$	$g_{\mu\nu}, A_{[3]}, \psi_\mu^\alpha$

Local SUSY (off-shell algebra):

$$\delta e^a = \bar{\varepsilon} \Gamma^a \psi, \quad \delta \omega^{ab} = 0, \quad \delta \psi = D\varepsilon,$$

$$\delta b^{ab} = \bar{\varepsilon} \Gamma^{ab} \psi, \quad \delta b^{abcde} = \bar{\varepsilon} \Gamma^{abcde} \psi, \quad \text{etc.}$$

This is an exact local symmetry: no auxiliary fields, no on-shell conditions, no assumptions on the background ($T^a \equiv 0$, etc.)

Summary

- The equivalence principle leads to a geometric action where the vielbein and the spin connection are the dynamical fields.
- These theories are invariant under local Lorentz transformations and invariant (by construction) under diffeomorphisms.
- They give rise to 2nd order field eqs. for $g_{\mu\nu}$, and spin $S \leq 2$.
- Around a flat background ($R^{ab}=0=T^a$) these theories behave like ordinary GR (Einstein-Hilbert). However, this is just one of the many sectors of the theory.
- These theories possess a number of indeterminate dimensionful parameters $\{\alpha_p, \beta_q\}$.
- In odd dimensions and for a particular choice of these parameters, the theory has no dimensionful constants and its symmetry is enhanced to (A)dS. The resulting action is a Chern-Simons form for the (A)dS or Poincaré groups.

- The classical evolution can go from one sector to another with fewer degrees of freedom in an irreversible manner.
- The supersymmetric extensions only exist for special combinations of AdS or Poincaré invariant gravitational actions.
- The maximally (super)symmetric configuration is a vacuum state that has no propagating degrees of freedom around it. At that point the theory is topological.
- This shows that the standard supergravity of Cremmer, Julia and Sherk is not the only consistent supersymmetric field theory containing gravity in $11D$. There exist at least 3 more:
 - Super-AdS [Osp(32|1)]
 - Super-Poincaré
 - M-Algebra

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Ευχαριστω!