



Geons in quadratic Palatini gravity

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Motivations

- **Quadratic gravity** has been thoroughly studied in the literature:
 - ◆ $R \Rightarrow R + l_P^2 (aR^2 + bR_{\mu\nu}R^{\mu\nu})$
 - ◆ Renormalizability of **QFT in curved spaces** requires such terms.
 - ◆ **String theories** predict similar higher-order curvature corrections.

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Field Equations

Spherically symmetric charged
solutions

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- It has been established that:

- ◆ The theory satisfies **fourth-order equations**.
- ◆ Massive spin-2 gravitons and a massive spin-0 **ghost** propagate in vacuum.
- ◆ **Perturbative methods** are generally required to explore deviations from GR.

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 - ◆ **Perturbative methods** are generally required to explore deviations from GR.
- **HOWEVER**, the **Palatini** version of the theory is **completely different**.
 - ◆ **Second-order** field equations govern the dynamics \Rightarrow **exact solutions**.
 - ◆ Only massless spin-2 gravitons propagate in vacuum \Rightarrow **GR+ Λ** recovered.
 - ◆ **New dynamics without new dynamical d.o.f.**

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 - ◆ Only massless spin-2 gravitons propagate in vacuum \Rightarrow **GR+ Λ** recovered.
 - ◆ **New dynamics without new dynamical d.o.f.**
- We will see that in the **quadratic Palatini theory**:
 - ◆ The **central singularity** of charged BHs is **replaced by a Wormhole**.
 - ◆ Reissner-Nordstrom solutions turn into **geons**.
 - ◆ **Stable remnants** arise in the lowest part of the mass and charge spectrum.

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- Palatini approach
- Dynamics of $f(R, Q)$
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Field Equations in Palatini Theories



Palatini -Vs- Metric: generalities

- In the Palatini formalism, one assumes that $g_{\mu\nu}$ and $\Gamma_{\beta\gamma}^{\alpha}$ are independent entities: $S = \int d^n x \sqrt{-g} L[g_{\mu\nu}, \Gamma_{\beta\gamma}^{\alpha}] + S_{matter}[g_{\mu\nu}, \Psi_m]$

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- The field equations follow from variation of the action:

- ◆ **Palatini approach:**

$$\delta S = \int d^n x \left[\sqrt{-g} \left(\frac{\delta L}{\delta g^{\mu\nu}} - \frac{L}{2} g_{\mu\nu} \right) \delta g^{\mu\nu} + \sqrt{-g} \frac{\delta L}{\delta \Gamma_{\beta\gamma}^{\alpha}} \delta \Gamma_{\beta\gamma}^{\alpha} \right] + \delta S_{matter}$$

$$\delta g^{\mu\nu} \Rightarrow \frac{\delta L}{\delta g^{\mu\nu}} - \frac{L}{2} g_{\mu\nu} = 8\pi G T_{\mu\nu}$$

$$\delta \Gamma_{\beta\gamma}^{\alpha} \Rightarrow \frac{\delta L}{\delta \Gamma_{\beta\gamma}^{\alpha}} = 0$$

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- ◆ **Metric approach:**

The relation $\delta \Gamma_{\beta\gamma}^{\alpha} = \frac{g^{\alpha\rho}}{2} [\nabla_{\beta} \delta g_{\rho\gamma} + \nabla_{\gamma} \delta g_{\rho\beta} - \nabla_{\rho} \delta g_{\beta\gamma}]$ implies

$$\frac{\delta L}{\delta \Gamma_{\beta\gamma}^{\alpha}} \delta \Gamma_{\beta\gamma}^{\alpha} = \left\{ g^{\alpha\mu} \frac{\delta L}{\delta \Gamma_{\lambda\nu}^{\alpha}} - \frac{g^{\alpha\lambda}}{2} \frac{\delta L}{\delta \Gamma_{\mu\nu}^{\alpha}} \right\} \nabla_{\lambda} \delta g_{\mu\nu} \text{ and leads to}$$

$$\delta g^{\mu\nu} \Rightarrow \left(\frac{\delta L}{\delta g^{\mu\nu}} - \frac{L}{2} g_{\mu\nu} \right) + \nabla_{\lambda} \left[g_{\gamma\nu} \frac{\delta L}{\delta \Gamma_{\lambda\gamma}^{\mu}} - g_{\beta\mu} g_{\gamma\nu} g^{\alpha\lambda} \frac{\delta L}{\delta \Gamma_{\beta\gamma}^{\alpha}} \right] = 8\pi G T_{\mu\nu}$$

- Metric and **Palatini** variations generally lead to **different field equations**.

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Dynamics of Palatini $f(R, Q)$ gravity

■ Consider the **Palatini** theory $S[g, \Gamma, \Psi_m] = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} f(R, Q) + S_m[g, \Psi_m]$

with $Q \equiv R_{\mu\nu} R^{\mu\nu}$, $R_{\mu\nu} = R^\rho{}_{\mu\rho\nu}$, and $R^\alpha{}_{\beta\mu\nu} = \partial_\mu \Gamma^\alpha_{\nu\beta} - \partial_\nu \Gamma^\alpha_{\mu\beta} + \Gamma^\alpha_{\mu\lambda} \Gamma^\lambda_{\nu\beta} - \Gamma^\alpha_{\nu\lambda} \Gamma^\lambda_{\mu\beta}$

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- Field equations (for vanishing torsion): $f_X \equiv \partial_X f$

- ◆ $g_{\mu\nu} \Rightarrow f_R R_{\mu\nu} - \frac{f}{2} g_{\mu\nu} + 2f_Q R_{\mu\alpha} R^\alpha{}_\nu = \kappa^2 T_{\mu\nu}$

- ◆ $\Gamma^\alpha{}_{\mu\nu} \Rightarrow \nabla_\beta [\sqrt{-g} (f_R g^{\mu\nu} + 2f_Q R^{\mu\nu})] = 0$

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- ◆ The connection seems to satisfy second-order differential equations.
- ◆ Only first-order derivatives of the metric.

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- ◆ Only first-order derivatives of the metric.

- In the **Palatini** version of **GR**:

- ◆ The connection can be **solved by algebraic means**.
- ◆ The dynamics boils down to **second-order equations** for the metric.

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- In these equations:

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 - Only first-order derivatives of the metric.

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- In general $f(R, Q)$ **Palatini** theories we find that:

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Solving for $\Gamma_{\mu\nu}^{\alpha}$

- Recall the field equations: $f_X \equiv \partial_X f$

- ◆ $g_{\mu\nu} \Rightarrow f_R R_{\mu\nu} - \frac{f}{2} g_{\mu\nu} + 2f_Q R_{\mu\alpha} R^{\alpha}_{\nu} = \kappa^2 T_{\mu\nu} \quad (1)$

- ◆ $\Gamma_{\mu\nu}^{\alpha} \Rightarrow \nabla_{\beta} [\sqrt{-g} (f_R g^{\mu\nu} + 2f_Q R^{\mu\nu})] = 0 \quad (2)$

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- ◆ $\Gamma_{\mu\nu}^\alpha \Rightarrow \nabla_\beta [\sqrt{-g} (f_R g^{\mu\nu} + 2f_Q R^{\mu\nu})] = 0$ (2)

- Step 1: define $P_\mu{}^\nu \equiv R_{\mu\alpha} g^{\alpha\nu}$, with $R = \text{Tr}[\hat{P}]$ and $Q = \text{Tr}[\hat{P}^2]$, and write (1) as

$$f_R P_\mu{}^\nu - \frac{f}{2} \delta_\mu{}^\nu + 2f_Q P_\mu{}^\alpha P_\alpha{}^\nu = \kappa^2 T_\mu{}^\nu \Leftrightarrow 2f_Q \left(\hat{P} + \frac{f_R}{4f_Q} \hat{I} \right)^2 = \kappa^2 \hat{T} + \frac{1}{2} \left(f + \frac{f_R^2}{4f_Q} \right) \hat{I}$$

- ◆ This establishes an **algebraic relation** between $P_\mu{}^\nu$ and $T_\mu{}^\nu$: $\hat{P} = \hat{P}(\hat{T})$

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- ◆ This establishes an **algebraic relation** between P_{μ}^{ν} and T_{μ}^{ν} : $\hat{P} = \hat{P}(\hat{T})$

- Step 2: since $\hat{P} = \hat{P}(\hat{T})$, with $R = \text{Tr}[\hat{P}] = R(\hat{T})$ and $Q = \text{Tr}[\hat{P}^2] = Q(\hat{T})$, then:

$$(f_R g^{\mu\nu} + 2f_Q R^{\mu\nu}) \Rightarrow g^{\mu\alpha} (f_R \delta_{\alpha}^{\nu} + 2f_Q P_{\alpha}^{\nu}) \equiv g^{\mu\alpha} \Sigma_{\alpha}^{\nu}(\hat{T}) .$$

- ◆ One finds that (2) $\rightarrow \nabla_{\beta} [\sqrt{-g} g^{\mu\alpha} \Sigma_{\alpha}^{\nu}] = 0 \Leftrightarrow \nabla_{\beta} [\sqrt{-h} h^{\mu\nu}] = 0$

- ◆ with $h^{\mu\nu} = \frac{g^{\mu\alpha} \Sigma_{\alpha}^{\nu}}{\sqrt{\det \Sigma}}$, $h_{\mu\nu} = \left(\sqrt{\det \Sigma} \right) [\Sigma^{-1}]_{\mu}^{\alpha} g_{\alpha\nu}$

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- $\Gamma_{\mu\nu}^{\alpha}$ turns out to be the **Levi-Civita connection** of $h_{\mu\nu}$.

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Metric field equations

■ From the definition of $P_\mu{}^\nu \equiv R_{\mu\alpha}g^{\alpha\nu}$ and $\Sigma_\alpha{}^\nu = (f_R\delta_\alpha{}^\nu + 2f_Q P_\alpha{}^\nu)$, we have:

◆ Step 1: Metric variation $\Rightarrow f_R P_\mu{}^\nu - \frac{f}{2}\delta_\mu{}^\nu + 2f_Q P_\mu{}^\alpha P_\alpha{}^\nu = \kappa^2 T_\mu{}^\nu$

◆ Step 2: $P_\mu{}^\alpha \Sigma_\alpha{}^\nu = \kappa^2 T_\mu{}^\nu + \frac{f}{2}\delta_\mu{}^\nu$

◆ Step 3: rewriting $P_\mu{}^\alpha \Sigma_\alpha{}^\nu$ as $R_{\mu\alpha} h^{\alpha\nu} \sqrt{\det \hat{\Sigma}}$, we get

$$R_\mu{}^\nu(h) = \frac{1}{\sqrt{\det \hat{\Sigma}}} \left(\frac{f}{2}\delta_\mu{}^\nu + \kappa^2 T_\mu{}^\nu \right)$$

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■ $h_{\mu\nu}$ satisfies an Einstein-like set of pde's.

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■ $h_{\mu\nu}$ satisfies an Einstein-like set of pde's.

■ From Step 1 $\Rightarrow 2f_Q \left(\hat{P} + \frac{f_R}{4f_Q} \hat{I} \right)^2 = \kappa^2 \hat{T} + \frac{1}{2} \left(f + \frac{f_R^2}{4f_Q} \right) \hat{I}$.

◆ The square root of this equation yields $\hat{P} = \hat{P}(\hat{T})$.

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- ◆ Step 2: $P_{\mu}^{\alpha} \Sigma_{\alpha}^{\nu} = \kappa^2 T_{\mu}^{\nu} + \frac{f}{2} \delta_{\mu}^{\nu}$

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$$R_{\mu}^{\nu}(h) = \frac{1}{\sqrt{\det \hat{\Sigma}}} \left(\frac{f}{2} \delta_{\mu}^{\nu} + \kappa^2 T_{\mu}^{\nu} \right)$$

- $h_{\mu\nu}$ satisfies an Einstein-like set of pde's.

- From Step 1 $\Rightarrow 2f_Q \left(\hat{P} + \frac{f_R}{4f_Q} \hat{I} \right)^2 = \kappa^2 \hat{T} + \frac{1}{2} \left(f + \frac{f_R^2}{4f_Q} \right) \hat{I}$.

- ◆ The square root of this equation yields $\hat{P} = \hat{P}(\hat{T})$.

- In vacuum: $\hat{P}_V = A(R_V, Q_V) \hat{I} \Rightarrow R_V = 4A(R_V, Q_V)$, $Q_V = 4A^2(R_V, Q_V)$.

- ◆ Solutions: R_V and Q_V are constant \Rightarrow de Sitter space-times: $Q_V = R_V^2/4$



Metric field equations

● Motivations

Field Equations

- Palatini approach
- Dynamics of $f(R, Q)$
- Solving for $\Gamma_{\mu\nu}^\alpha$
- Metric field equations

Spherically symmetric charged solutions

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- From the definition of $P_\mu{}^\nu \equiv R_{\mu\alpha} g^{\alpha\nu}$ and $\Sigma_\alpha{}^\nu = (f_R \delta_\alpha{}^\nu + 2f_Q P_\alpha{}^\nu)$, we have:

- ◆ Step 1: Metric variation $\Rightarrow f_R P_\mu{}^\nu - \frac{f}{2} \delta_\mu{}^\nu + 2f_Q P_\mu{}^\alpha P_\alpha{}^\nu = \kappa^2 T_\mu{}^\nu$

- ◆ Step 2: $P_\mu{}^\alpha \Sigma_\alpha{}^\nu = \kappa^2 T_\mu{}^\nu + \frac{f}{2} \delta_\mu{}^\nu$

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- ◆ Solutions: R_V and Q_V are constant \Rightarrow de Sitter space-times: $Q_V = R_V^2/4$
- ◆ In vacuum, the theory boils down to GR+ Λ , with $g_{\mu\nu} = \text{constant} \times h_{\mu\nu}$.
- ◆ No massive spin-2 gravitons. No ghost-like instabilities.
- ◆ Matter-induced nonlinearities instead of new d.o.f.



- Motivations

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- Electrovacuum geometry
- Smooth WH geometries
- Palatini Geons
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Spherically symmetric charged solutions



Electrovacuum geometry

- In GR, the RN solution is characterized by: ($r_S = 2M_0$ and $r_q^2 = 2Gq^2$)

$$R_{GR} = 0, \quad Q_{GR} \equiv R_{\mu\nu}R^{\mu\nu} = \frac{r_q^4}{r^8}, \quad K_{GR} \equiv R^\alpha{}_{\beta\mu\nu}R^\beta{}_{\alpha}{}^{\mu\nu} = \frac{12r_S^2}{r^6} - \frac{24r_S r_q^2}{r^7} + \frac{14r_q^4}{r^8}$$

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- When $r \gg r_c$:

$$R(g) \approx -\frac{48r_c^8}{r^{10}} + O\left(\frac{r_c^9}{r^{11}}\right), \quad Q(g) \approx \frac{r_q^4}{r^8} \left(1 - \frac{16l_P^2}{r^2} + \dots\right), \quad K(g) \approx K_{GR} + \frac{144r_S r_q^2 l_P^2}{r^9} + \dots$$

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- But when $z \equiv r/r_c \rightarrow 1^+$: $\left[\delta_1 = \frac{1}{2} \sqrt{\frac{r_q^3}{r_S^2 l_P}}, \delta_2 = \frac{r_c}{r_S}, \text{ and } \delta_1^* \approx 0.572 \right]$

$$r_c^2 R(g) \approx \left(-4 + \frac{16\delta_1^*}{3\delta_2}\right) + O(z-1) + \dots - \frac{1}{2\delta_2} \left(1 - \frac{\delta_1^*}{\delta_1}\right) \left[\frac{1}{(z-1)^{3/2}} - O\left(\frac{1}{\sqrt{z-1}}\right)\right]$$

$$r_c^4 Q(g) \approx \left(10 + \frac{86\delta_1^2}{9\delta_2^2} - \frac{52\delta_1}{3\delta_2}\right) + \dots + \left(1 - \frac{\delta_1^*}{\delta_1}\right) \left[\frac{6\delta_2 - 5\delta_1}{3\delta_2^2(z-1)^{3/2}} + \dots\right] + \left(1 - \frac{\delta_1^*}{\delta_1}\right)^2 \left[\frac{1}{8\delta_2^2(z-1)^3} - \dots\right]$$

$$r_c^4 K(g) \approx \left(16 + \frac{88\delta_1^2}{9\delta_2^2} - \frac{64\delta_1}{3\delta_2}\right) + \dots + \left(1 - \frac{\delta_1^*}{\delta_1}\right) \left[\frac{2(2\delta_1 - 3\delta_2)}{3\delta_2^2(z-1)^{3/2}} + \dots\right] + \left(1 - \frac{\delta_1^*}{\delta_1}\right)^2 \left[\frac{1}{4\delta_2^2(z-1)^3} + \dots\right]$$

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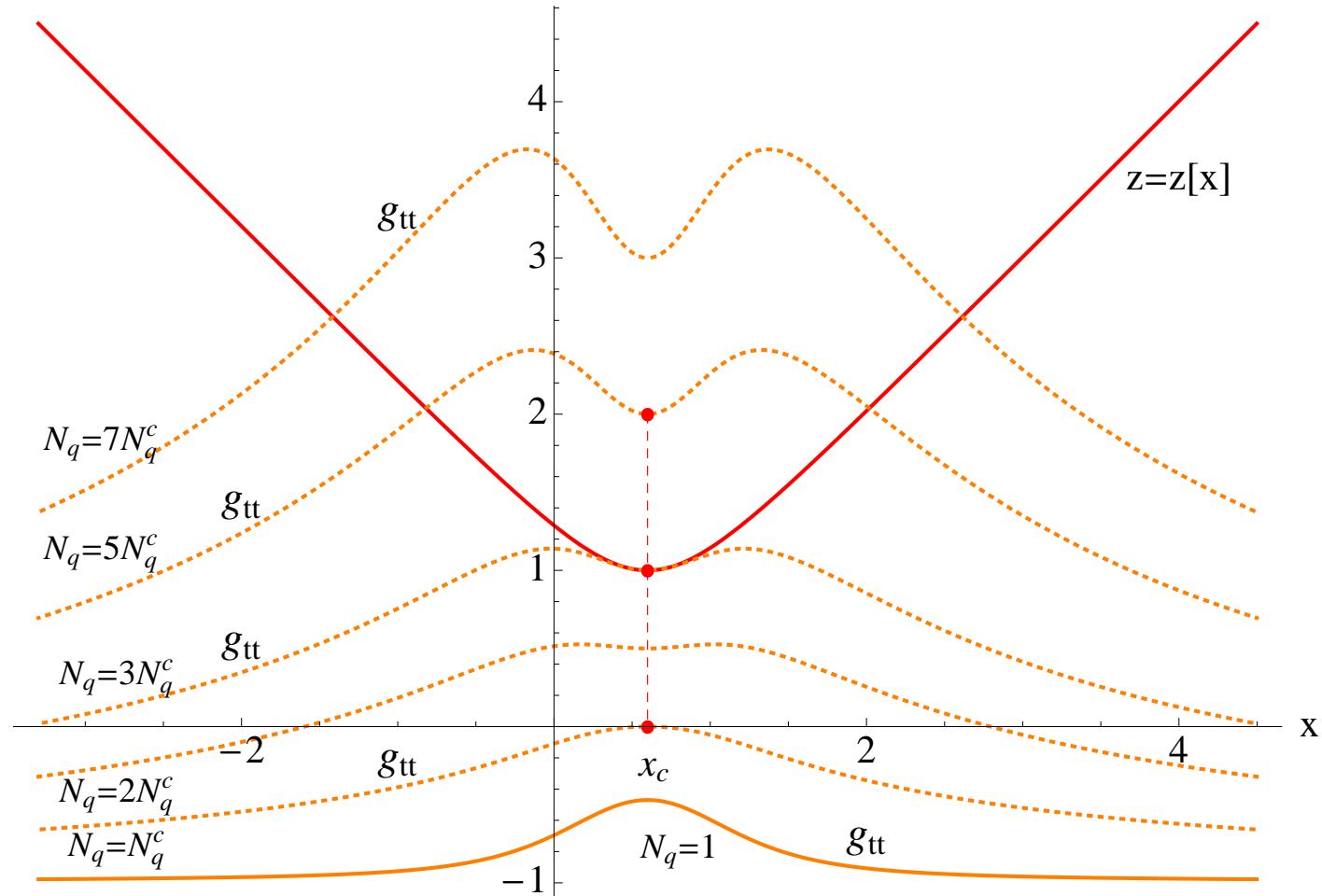
● Summary and Conclusions

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Smooth wormhole geometries

- When $\delta_1 = \delta_1^*$ a wormhole arises at $r = r_c \Leftrightarrow z = 1$:



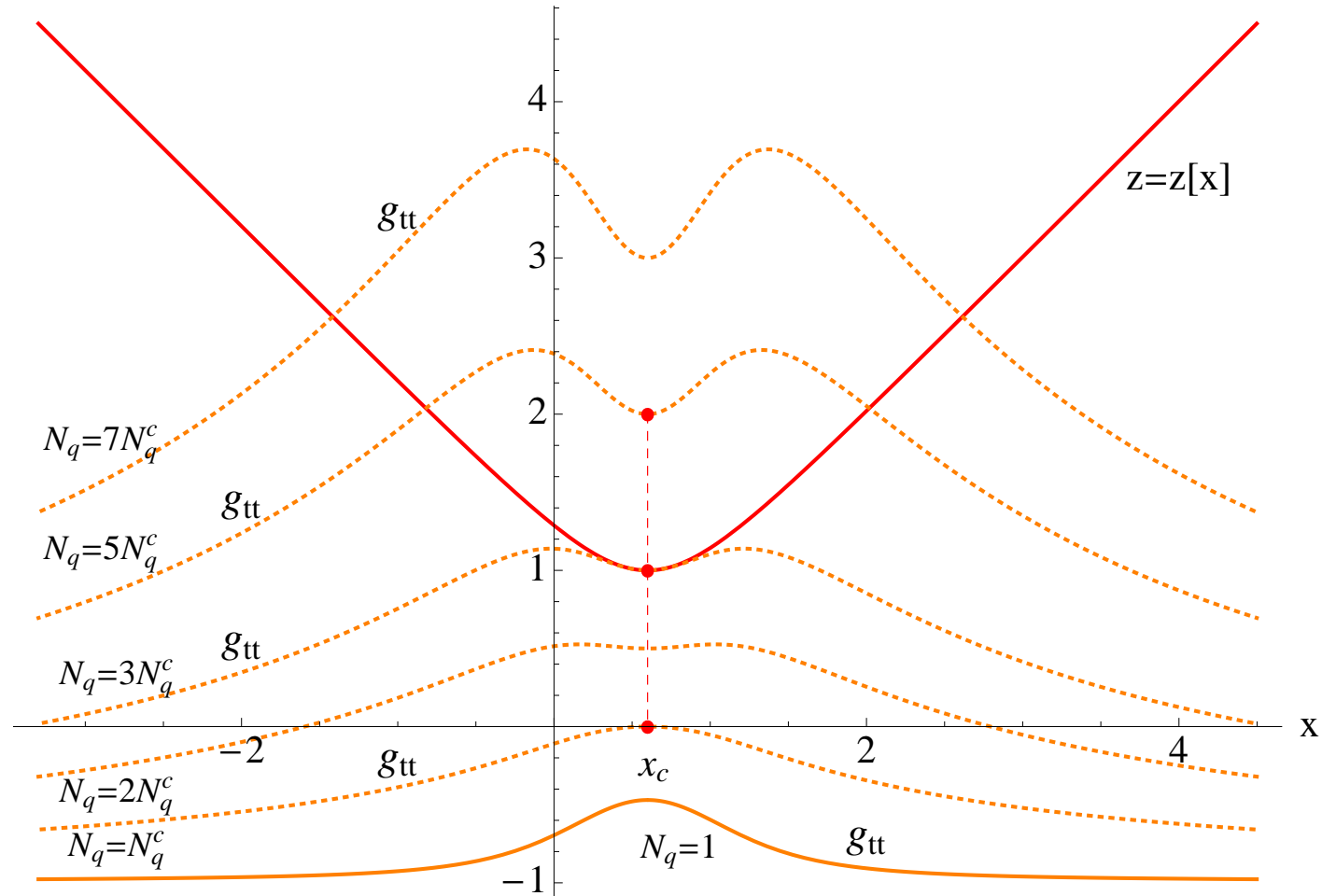
- $$ds^2 = g_{tt} dt^2 - \frac{1}{g_{tt}} dx^2 + r^2(x) d\Omega^2 .$$

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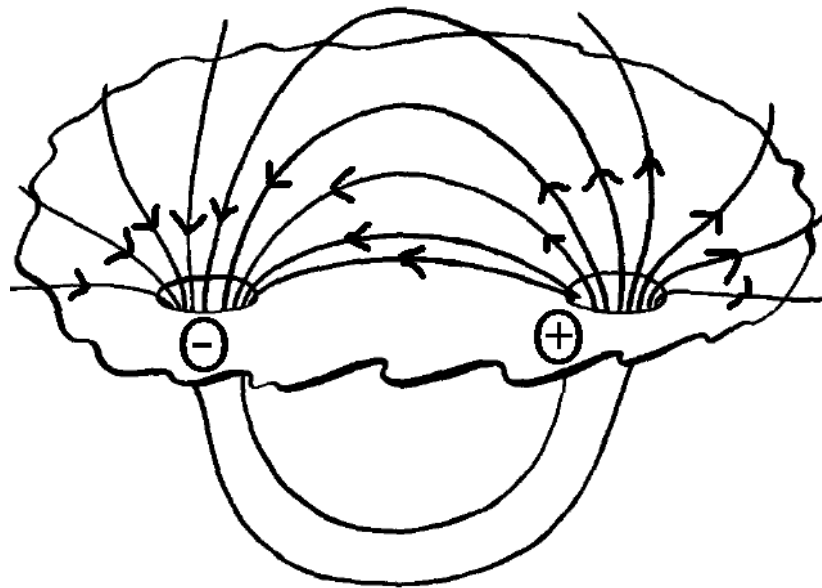
- Note that for $N_q \leq N_q^c \approx 16.55$ charges there is **no event horizon**.

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Palatini Geons

- Electric charge of **geons** (\equiv self-gravitating electromagnetic entities):
 - ◆ The lines of force of the electric field enter through one of the wormhole mouths and exit through the other creating the **illusion of a negatively charged object** on one side and a positively charged object on the other.



- ◆ The locally measured **electric charge is defined by the flux**
 $\Phi \equiv \int_S *F = 4\pi q$ through any 2-surface S enclosing a wormhole mouth.

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- Evaluating the action on the solutions we find:

$$S_T = S_{Quad.Grav.} + S_{e.m.} = 2M_0 c^2 \frac{\delta_1}{\delta_1^*} \int dt . ([S] = [\text{Energy}] \times [\text{time}])$$

- ◆ For $\delta_1 = \delta_1^* \Rightarrow M_0 c^2 = \text{e.m.} + \text{grav. binding energy}!!! .$
- ◆ Coincides with the action of a point-like particle at rest !!!???
(**geonic soliton**): $S_{p.p.} = mc^2 \int dt \sqrt{1 - v^2/c^2}$



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(**geonic soliton**): $S_{p.p.} = mc^2 \int dt \sqrt{1 - v^2/c^2}$

- Density of lines of force crossing the wormhole throat:

$$\mathcal{E} = \frac{\Phi}{4\pi r_c^2} = \frac{q}{r_c^2} = \sqrt{\frac{c^7}{2(\hbar G)^2}} \Rightarrow \text{Universal quantity.}$$

- ◆ Since \mathcal{E} is **independent of q and M** \Rightarrow **geon structure even when $\delta_1 \neq \delta_1^*$** .
- ◆ WH (**topological**) structure even if there are (**local**) curvature divergences.

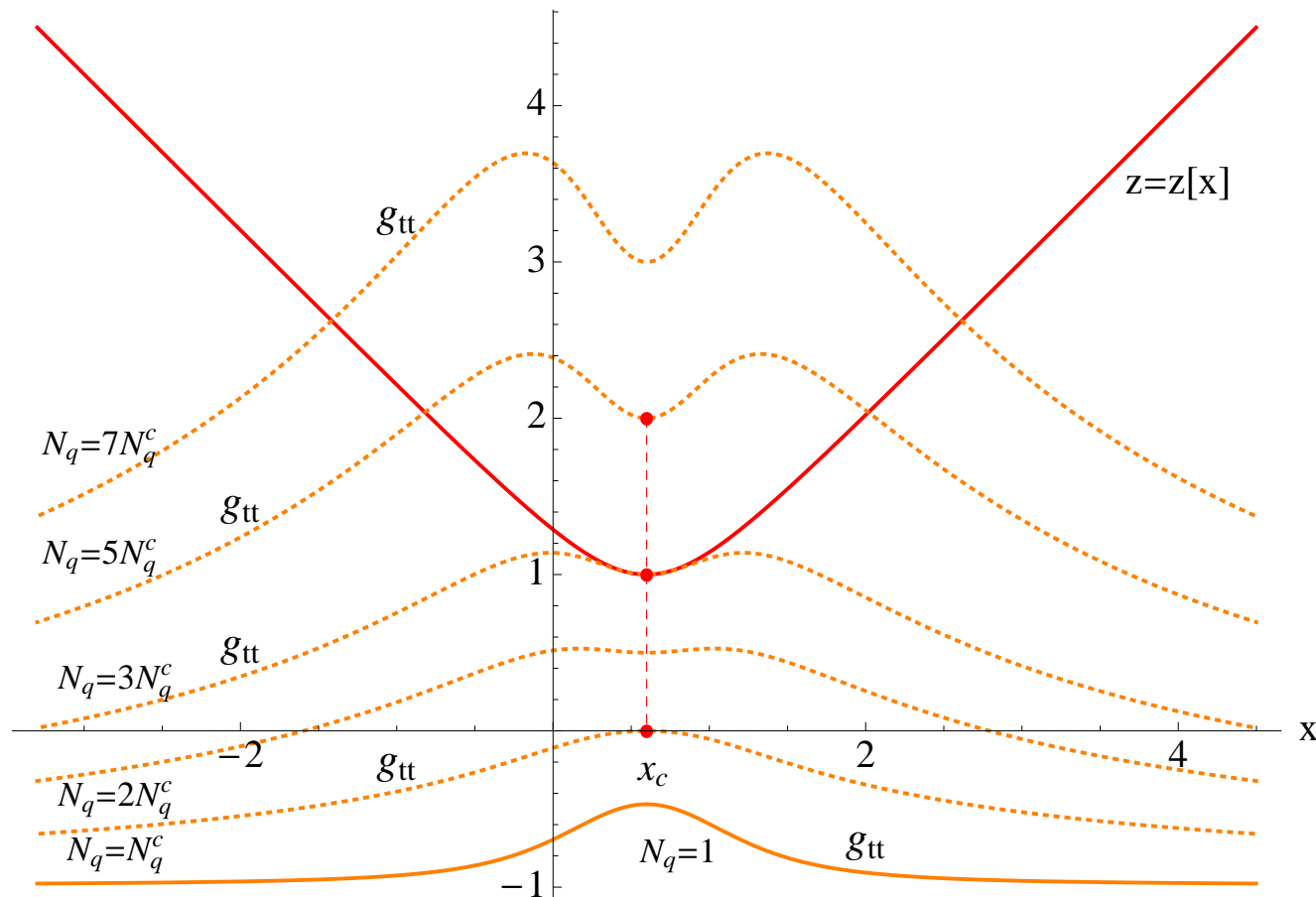


Stability and quantum properties

- The **event horizon** is expected to **force the decay** into $\delta_1 \neq \delta_1^*$ states:

- ◆ $g_{tt} = \frac{(1-\delta_1/\delta_1^*)}{4\delta_2\sqrt{z-1}} - \frac{1}{2} \left(1 - \frac{\delta_1}{\delta_2}\right) + O(\sqrt{z-1})$ Recall: $\delta_1 = \frac{1}{2r_S} \sqrt{\frac{r_q^3}{l_P}}$

- ◆ The sign of $(1 - \delta_1^*/\delta_2)$ determines if there is an event horizon.



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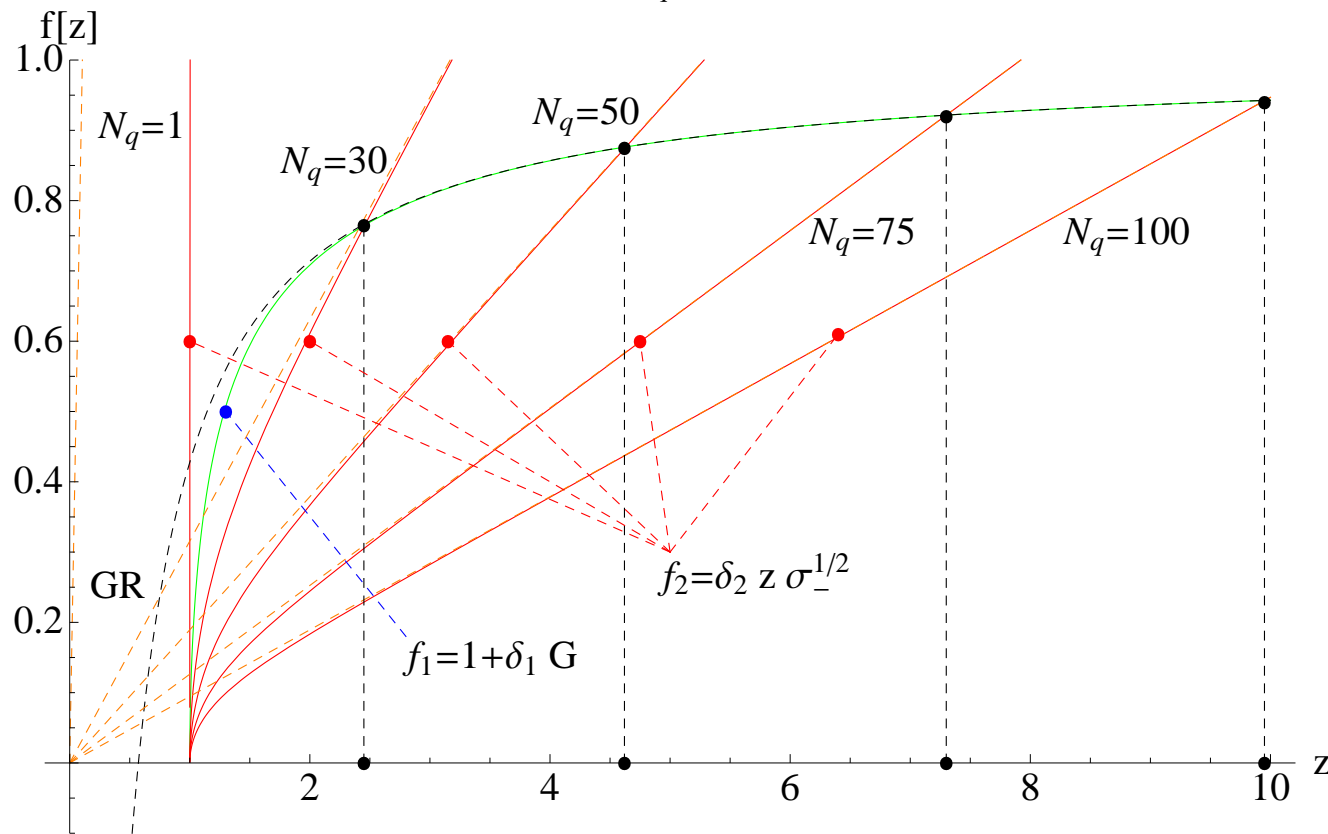
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Horizon location as N_q changes when $\delta_1 = \delta_1^*$



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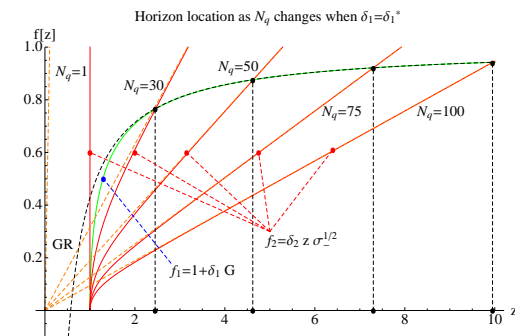
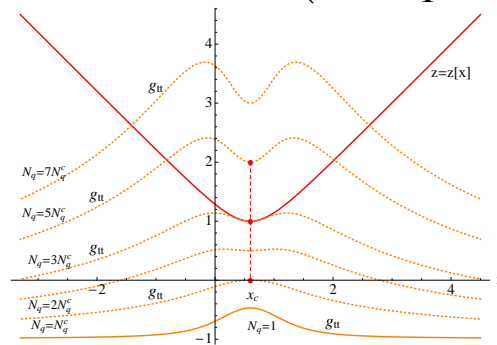


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- ◆ The sign of $(1 - \delta_1^*/\delta_2)$ determines if there is an event horizon.



- For $\delta_1 = \delta_1^*$ we find:

- ◆ Mass spectrum:
$$M \approx 1.23605 \left(\frac{N_q}{N_q^c} \right)^{3/2} m_P$$
, where $N_q^c \equiv \sqrt{2/\alpha_{em}} \approx 16.55$

- ◆ The absence of an event horizon for $N_q < 16.55$ yields **quantum mechanically stable objects: NO Hawking decay.**
- ◆ The **topological nature of their charge** makes them stable against arbitrary classical perturbations that preserve the topology.

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Hawking estimates and BH remnants

Mon. Not. R. astr. Soc. (1971) **152**, 75–78.

GRAVITATIONALLY COLLAPSED OBJECTS OF VERY LOW MASS

Stephen Hawking

(Communicated by M. J. Rees)

(Received 1970 November 9)

SUMMARY

It is suggested that there may be a large number of gravitationally collapsed objects of mass 10^{-5} g upwards which were formed as a result of fluctuations in the early Universe. They could carry an electric charge of up to ± 30 electron units. Such objects would produce distinctive tracks in bubble chambers and could form atoms with orbiting electrons or protons. A mass of 10^{17} g of such objects could have accumulated at the centre of a star like the Sun. If such a star later became a neutron star there would be a steady accretion of matter by a central collapsed object which could eventually swallow up the whole star in about ten million years.

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Hawking estimates and BH remnants

- **Hawking's estimates**, based entirely on the process of classical collapse, are in **excellent quantitative agreement with our results**:
 - ◆ According to Hawking, a large number of objects with $M \sim m_P$ and $N_q \lesssim 30$ could have been formed in the early universe.
 - ◆ A fraction of them could reach the stability conditions found here.
 - ◆ They may also arise from the evaporation of more massive objects.

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■ The existence of **stable solutions** in the lowest part of the mass and charge spectrum which can be **continuously connected with black hole states**, supports the view that these objects can be naturally identified as **black hole remnants**.

- ◆ Stable remnants implies a maximum temperature in the evaporation process.
- ◆ The lack of observations supporting black hole explosions is compatible with this result.

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Summary and Conclusions

- For a given $f(R, Q)$ Lagrangian, the field eqs in metric and **Palatini** differ:
 - ◆ **Metric formalism**: higher-order derivatives due to integration by parts.
 - ◆ **Palatini formalism**: second-order equations and algebraic relations.
 - ◆ **Lovelock theories** are an exception: metric and Palatini coincide !!!

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- **Geons in Quadratic Palatini gravity**:
 - ◆ The existence of **completely regular (non-perturbative) solutions with WH structure** put forward the **geonic nature** of such solutions.
 - ◆ **Universal properties** of the electric flux at $r = r_c$:
$$\frac{\Phi}{4\pi r_c^2} = \sqrt{\frac{c^7}{2(\hbar G)^2}}$$
 \Rightarrow the geonic structure persists even when curvature divergences exist.

All the spherically symmetric electrovacuum solutions are **geons**

● Motivations

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All the spherically symmetric electrovacuum solutions are **geons**
- The existence of **stable solutions** in the lowest part of the mass and charge spectrum which can be **continuously connected with black hole states**, supports the view that these objects can be naturally identified as **black hole remnants**.
- Nontrivial implications for **dark matter** and the **information loss problem**.



● Motivations

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solutions

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Thanks !!!