Miguel Cruz in colaboration with Rubén Cordero, Alberto Molgado and Efraín Rojas arXiv:1309.3031

Seventh Aegean Summer School, Paros, Greece



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Outline

The action of the model FRW-minisuperspace

Ostrogradski-Hamiltonian approach Gauge fixing and Canonical transformation

Quantum approach Nucleation rate

Concluding remarks



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The action of the model

We consider the following action

$$S[X] = \underbrace{\int d^4 \sqrt{-g} \left(-\Lambda + K + \frac{-\pi}{2}R\right)}_{\text{First brane Lovelock invariants}}$$
(1)

Important:

- This action is invariant under reparameterizations of the worldvolume.
- We consider the Minkowski spacetime as background.
- We have the codimension-1 case.





The induced geometry on the brane

Induced metric on *m* (First fundamental form):

$$g_{ab} = \ _{\mu\nu} X^{\mu}_a X^{\nu}_b; \qquad (2)$$

and the extrinsic curvature (Second fundamental form)

$$K^{i}_{ab} = -n^{i}_{\mu} \nabla_{a} X^{\mu}_{b} = -n^{i}_{\mu} \nabla_{a} \nabla_{b} X^{\mu}:$$
(3)

R. Capovilla y J. Guven, gr-qc/9411060



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By a variational procedure we obtain

$$G_{ab}K^{ab} - \mathcal{R} + \Lambda K = 0:$$
⁽⁴⁾

We have:

- ▶ *G_{ab}* is the brane Einstein tensor.
- ► The contracted Gauss-Codazzi integrability condition: $\mathcal{R} = K^2 - K_{ab}K^{ab}$.
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—The action of the model

FRW-minisuperspace

In this configuration, $S = 6^{-2} \int d L$, where the Lagrangian can be written as

$$L = \frac{a\dot{t}}{N^3}(a\ddot{a}\dot{t} - a\dot{a}\ddot{t} + N^2\dot{t}) - Na^3\bar{\Lambda}^2 + \frac{a^3}{N^2}(\dot{t}\ddot{a} - \dot{a}\ddot{t}) + 3a^2\dot{t}$$
(5)

We expect a conserved quantity. Certainly, L can be split as $L = L_b + L_d$ as follows:

$$L_b = \frac{d}{d} \left[\frac{a^2 \dot{a}}{N} + a^3 \operatorname{-arctanh} \left(\frac{\dot{a}}{\dot{t}} \right) \right]; \tag{6}$$

and

$$L_{d} = -\frac{a\dot{a}^{2}}{N} + aN(1 - a^{2}\bar{\Lambda}^{2}) + 3a^{2}\left[\dot{t} - \dot{a}\operatorname{arctanh}\left(\frac{\dot{a}}{\dot{t}}\right)\right]:$$
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The configuration space is given by $\{t(\); a(\); \dot{t}; \dot{a}\}$. We have defined the quantity $N^2 := \dot{t}^2 - \dot{a}^2$.



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$$\mathcal{F}_1 = P \cdot \dot{X} \approx 0; \tag{8}$$

$$\mathcal{F}_2 = \left(\frac{N\Theta}{a^2\Phi}\right)'_2 + H_0 \approx 0; \tag{9}$$

$$S_1 = N(P \cdot n) - \frac{a^2}{N}(\dot{t} + Na) = 2 \approx 0;$$
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$$S_2 = p_t \dot{a} + p_a \dot{t} = \prime_4 \approx 0; \qquad (11)$$



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Ostrogradski-Hamiltonian approach

Gauge fixing and Canonical transformation

We use the canonical transformation

$$N := \sqrt{\dot{t}^2 - \dot{a}^2};$$

$$\Pi_N := \frac{1}{N} (P \cdot \dot{X});$$

$$v := -\left[N(P \cdot n) - \frac{a^2}{N}(\dot{t} + Na)\right];$$

$$\Pi_v := \arctan\left(\frac{\dot{a}}{\dot{t}}\right):$$

Then, we can write

$$\begin{split} \mathcal{F}_1 &= & \mathsf{N}\Pi_{\mathcal{N}}, \\ \mathcal{F}_2 &= & -\frac{\mathsf{N}}{\mathsf{a}\left[(\gamma-1)\mathfrak{a}^2\bar{\lambda}^2+2+3\bar{\beta}\sqrt{\gamma}\mathfrak{a}^2\bar{\lambda}\right]} \, \left\{ \left(\mathbf{p}_{\mathfrak{a}}+3\bar{\beta}\mathfrak{a}^2\Pi_{}\right)^2 - \mathfrak{a}\left(\frac{p_t^2}{\mathfrak{a}\left[(\gamma-1)\mathfrak{a}^2\bar{\lambda}^2+3\bar{\beta}\sqrt{\gamma}\mathfrak{a}^2\bar{\lambda}\right]}\right. \\ & + & \left.\frac{1}{\mathfrak{a}^3}N^2\Pi_{\mathcal{N}}^2+2\mathfrak{a}(\gamma\mathfrak{a}^2\bar{\lambda}^2-1)-(\gamma-1)\mathfrak{a}^3\bar{\lambda}^2-3\bar{\beta}\mathfrak{a}^3\sqrt{\gamma}\bar{\lambda}\right)\left[(\gamma-1)\mathfrak{a}^2\bar{\lambda}^2+2+3\bar{\beta}\sqrt{\gamma}\mathfrak{a}^2\bar{\lambda}\right] \right\}, \end{split}$$



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The constraints are quadratic in the momenta



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We adopt the standard way to pass to the quantum mechanical scheme,

$$p_t \longrightarrow \widehat{p}_t = -i \frac{@}{@t};$$
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$$\mathbf{p}_a \longrightarrow \widehat{\mathbf{p}}_a = -i \frac{@}{@a};$$
 (14)

$$\Pi_N \longrightarrow \widehat{\Pi}_N = -i \frac{@}{@N}; \qquad (15)$$

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Wheeler-DeWitt equation

 We assume then that Ψ is represented in the usual manner as

$$\Psi(a; t) := (a)e^{-i\Omega t}.$$
$$\left[-\frac{\partial^2}{\partial a^2} + U(a)\right] \quad (a) = 0.$$

Where:



$$U(a) = a^{2} \begin{bmatrix} (-1)a^{2}\bar{\Lambda}^{2} + 2 + 3\sqrt{a^{2}\bar{\Lambda}} \end{bmatrix}^{2} (1 - a^{2}\bar{\Lambda}^{2}):$$

Turning points for the potential:

$$\begin{array}{rcl} a_{l} &\simeq & \Omega, \\ a_{r} &\simeq & \frac{1}{\bar{\Lambda}} \left[\left(\frac{3\bar{\beta}}{2\bar{\Lambda}} \right) + \sqrt{\left(\frac{3\bar{\beta}}{2\bar{\Lambda}} \right)^{2} + 1} \right] - \frac{(\Omega/2)}{1 + \left(\frac{3}{2\bar{\Lambda}} \right) \left[\left(\frac{3}{2\bar{\Lambda}} \right) + \sqrt{\left(\frac{3}{2\bar{\Lambda}} \right)^{2} + 1} \right]}. \end{array}$$
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The nucleation probability: WKB approximation

$$\mathcal{P} \sim \exp\left(-2\int_{a_l}^{a_r} |\sqrt{U(a)}| da\right);$$
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• When we set $\Lambda = 0$, we obtain $\mathcal{P} \sim e^{-\frac{4}{27\beta^2}}$.

We have defined,
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—Quantum approach

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The Wheeler-DeWitt potential





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- The nucleation probability dictates a higher probability for negative values of where it is possible to obtain an effective cosmological constant in terms of the parameter. For a non-zero brane cosmological constant, the parameter which plays akin role to the crossover scale r_c in the self-accelerated branch of the DGP model, is similar as a cosmological constant at a quantum level^a.

^aClass. Quantum. Grav. **29** (2012) 175010



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