

# Chaotic Dynamics in the $SU(2)$ Gauge Matrix Model

Classical Ergodicity Breaking and Quantum Phases

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Workshop on Noncommutative and generalized geometry in  
string theory, gauge theory and related physical models

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Joint Work with Chaitanya Bhatt and Vijay Nenmeli  
**in progress**



# Introduction

Matrix Model for  $SU(2)$  Yang-Mills

The  $SU(2)$  Model in SVD Coordinates

A New Tetrahedral Symmetry

$SU(2)$  Matrix Model is the 3-d Henon-Heiles System

Stable Orbits and Their Classification

Perturbations About Periodic Orbits

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- ▶ Building blocks:
  - ▶  $2 \times 2$  hermitian matrices  $M_i(t)$ ,  $i = 1, 2, 3$ .
  - ▶ Equivalently, a single  $3 \times 3$  matrix  $M \in \mathcal{M}_3(\mathbb{R})$ .
- ▶ Physical rotations:  $M \rightarrow RM$ ,  $R \in SO(3)$ ,  
Gauge rotations:  $M \rightarrow MS^T$ ,  $S \in ad\ SU(2)$ .
- ▶ Gauge group  $ad\ SU(2)$  ( $= SO(3)$ ) is finite-dimensional.
- ▶ The configuration space  $\mathcal{C}_2 = \mathcal{M}_3(\mathbb{R})/SO(3)$ .
- ▶ Nontrivial topology:  $\mathcal{C}_2 \simeq \mathbb{R} \times (S^5 - \mathbb{R}P^2)$  (Narasimhan-Ramadas 1979).
- ▶ Curvature  $F_{ij} = -\epsilon_{ijk}M_k - i[M_i, M_j]$ .
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# Matrix Model Dynamics

- ▶ For dynamics, we need a gauge-invariant Lagrangian.
- ▶ The electric  $E_i \equiv D_t M_i$  and magnetic  $B_i \equiv \frac{1}{2} \epsilon_{ijk} F_{jk}$ :

$$E_i = \dot{M}_i - i[M_0, M_i], \quad B_i = -M_i - \frac{i}{2} \epsilon_{ijk} [M_j, M_k].$$

- ▶  $M_0$ : parallel transporter in the  $t$  direction. We set  $M_0 = 0$ .
- ▶ The matrix model Lagrangian is

$$L_{YM} = \frac{1}{2g^2} \text{Tr}(E_i E_i - B_i B_i) = \frac{1}{2g^2} \text{Tr}(D_t M_i D_t M_i) - V$$

(Balachandran-Queiroz-Vaidya 2015)

- ▶  $V(M)$  has upto quartic terms.
- ▶ The matrix model is just a multi-dimensional quartic oscillator.



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# SVD Coordinates

- ▶ We write  $M = RAS^T$ , with  $R \in O(3)$ ,  $S \in SO(3)$  (Iwai 2010).
- ▶  $A$  is diagonal with real entries.
- ▶  $R$  = physical rotations + parity;  $S$  = gauge rotations.
- ▶ Define  $\Omega = R^{-1}\dot{R}$  angular velocity,  $\Lambda = S^{-1}\dot{S}$  gauge angular velocity to get

$$L_{YM} = \frac{1}{2g^2} \text{Tr} \left( \dot{A}^2 - A^2(\Omega^2 + \Lambda^2) + 2\Omega A \Lambda \right) - V(A)$$

- ▶ The Hamiltonian can be computed easily.
- ▶ Gauge invariant dynamics: set  $\Lambda = 0$ .
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- ▶  $A$  is diagonal with real entries.
- ▶  $R$  = physical rotations + parity;  $S$  = gauge rotations.
- ▶ Define  $\Omega = R^{-1}\dot{R}$  angular velocity,  $\Lambda = S^{-1}\dot{S}$  gauge angular velocity to get

$$L_{YM} = \frac{1}{2g^2} \text{Tr} \left( \dot{A}^2 - A^2(\Omega^2 + \Lambda^2) + 2\Omega A \Lambda \right) - V(A)$$

- ▶ The Hamiltonian can be computed easily.
- ▶ Gauge invariant dynamics: set  $\Lambda = 0$ .
- ▶ We will focus on the gauge-invariant sector with  $\Omega = 0$ .



$$V(M) = \frac{1}{2g^2} \left( \text{Tr } M^T M - 6 \det M + \frac{1}{2} [(\text{Tr } M^T M)^2 - \text{Tr } M^T M M^T M] \right).$$

- ▶ With  $M = RAS^T$ , we have

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- ▶  $V$  possesses a *tetrahedral* symmetry  $\mathcal{T}_d$  in  $(a_1, a_2, a_3)$ .
- ▶ The full symmetry of  $H$  is  $\mathcal{T}_d \times \mathcal{T}$ ,  $\mathcal{T}$  is time-reversal.



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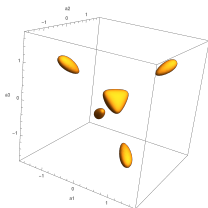
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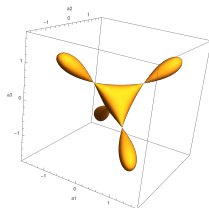


# Tetrahedral Symmetry

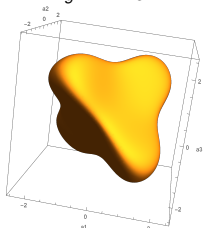
$$g^2 V = 0.05$$



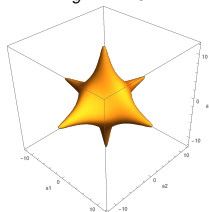
$$g^2 V = 0.09375$$



$$g^2 V = 1.5$$



$$g^2 V = 75$$



High energies: approximate octahedral symmetry.



# Extrema of $V(M)$

- ▶ The extrema are at  $a_1 = a_2 = a_3 = a$ , for  $a = 0, \frac{1}{2}, 1$ .
- ▶ Compute the Hessian  $\left[ \frac{\partial^2 V}{\partial a_i \partial a_j} \right]$ .
- ▶ It is positive definite at  $M = 0, \mathbf{1}$  (minima).
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# A 3-dimensional Henon-Heiles System

- ▶ In the color-0 spin-0 sector, the Hamiltonian is

$$H = \frac{g^2}{2} \left( p_{a_1}^2 + p_{a_2}^2 + p_{a_3}^2 \right) + \frac{1}{2g^2} \left( a_1^2 + a_2^2 + a_3^2 - 6a_1 a_2 a_3 + (a_1^2 a_2^2 + a_1^2 a_3^2 + a_2^2 a_3^2) \right)$$

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# Nonlinearity

- ▶ The constant energy surface is disconnected (5 pieces) for  $g^2 E < \frac{3}{32}$  and connected for  $g^2 E > \frac{3}{32}$ .
- ▶ The critical energy  $g^2 E_c = \frac{3}{32}$  is the saddle point  
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- ▶ The parameter  $\epsilon$  measures nonlinearity.
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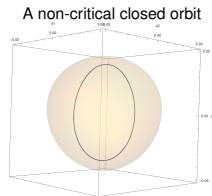
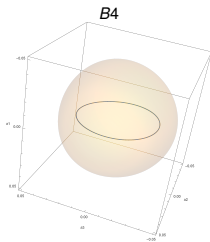
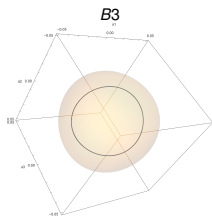
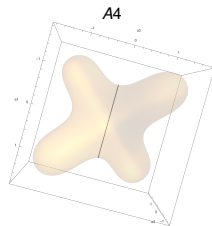
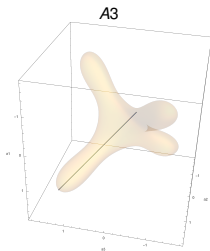
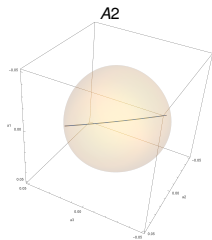
- ▶ Nonlinear normal modes are characterized by the stabilizer  $G \subset \mathcal{T}_d \times \mathcal{T}$  (Montaldi-Roberts-Stewart 1990).

Conjugacy class of Stabilizer	No. of orbits
$D_{2d} \times \mathcal{T} \quad (\equiv A_4)$	3
$C_{3v} \times \mathcal{T} \quad (\equiv A_3)$	4
$C_{2v} \times \mathcal{T} \quad (\equiv A_2)$	6
$S_4 \wedge \mathcal{T}_2 \quad (\equiv B_4)$	6
$C_3 \wedge \mathcal{T}_s \quad (\equiv B_3)$	8

Table: Non-linear normal modes - Classification



# Orbit Pictures



- ▶ **We will study perturbations about each type of orbit.**
- ▶ Generically, the orbits destabilize for large enough perturbations.
- ▶ Typically, the orbits destabilize by becoming weakly ergodic (Lyapunov = 0) and then (strongly) ergodic (Lyapunov  $\neq 0$ ).
- ▶ Orbits of  $A_4$  and  $A_3$  destabilize in a most interesting manner.
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# Monodromy Matrix I

- ▶ Let  $X(t)$  be a periodic solution of period  $T$ , and  $x(t)$  a perturbation around  $X$ .
- ▶ Then  $\frac{dx}{dt} = F[X(t)]x(t)$ .
- ▶ Equivalently, there is a time-evolution matrix  $U(t) : x(0) \rightarrow x(t)$ .
- ▶ The monodromy matrix is  $U(T)$ .
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- ▶ For Hamiltonian systems:
  - ▶ Eigenvalues come in reciprocal pairs
  - ▶ Always have at least two unit eigenvalues.
- ▶  $U$  is real so eigenvalues come in conjugate pairs.
- ▶ A periodic orbit is unstable iff at least one of its eigenvalues is outside the unit circle  $|z| = 1$ .



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# Stability of A4 Orbits

- ▶ Initial condition:  $(a_1, a_2, a_3, p_{a_1}, p_{a_2}, p_{a_3}) = (A, 0, 0, 0, 0, 0)$ .
- ▶ Fluctuation:  $(x_1, x_2, x_3, p_1, p_2, p_3)$ .
- ▶ Fluctuation equations:

$$\dot{x}_1 = p_1, \quad \dot{p}_1 = -x_1$$

$$\dot{x}_2 = p_2, \quad \dot{p}_2 = -x_2(1 + A^2 \cos^2 t) + (3A \cos t)x_3$$

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$$\ddot{b}_-(s) + (\eta + 2\alpha \cos 2s + 2\beta \cos 4s)b_-(s) = 0.$$

(Here  $s = \frac{t}{2}$ ,  $\alpha = -6A$ ,  $\beta = A^2$ ,  $\eta = 4 + 2A^2$ ).

- ▶ This is the Schrödinger equation for a particle in the periodic potential  $V(s) = -(\alpha \cos 2s + \beta \cos 4s)$  and energy  $\eta/2$ .
- ▶ The allowed energies form *bands*, and between them are *band gaps*.
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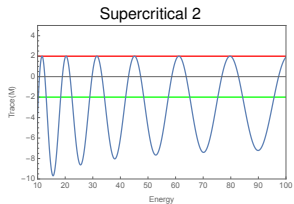
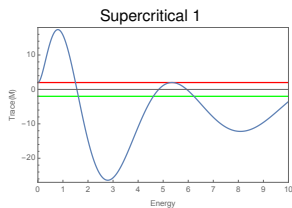
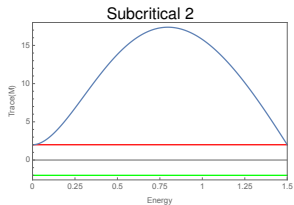
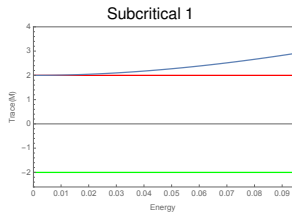


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- ▶ The time period can be computed exactly in terms of (in)complete elliptic functions of the first kind.
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Specifically, we find numerically that

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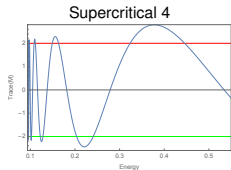
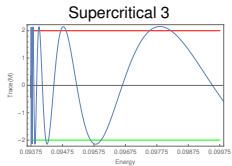
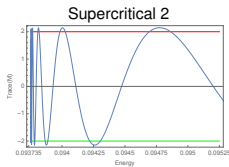
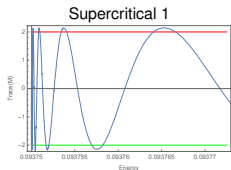
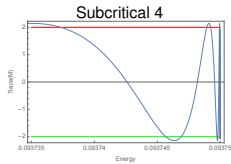
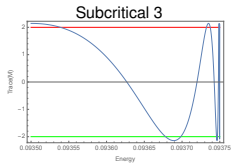
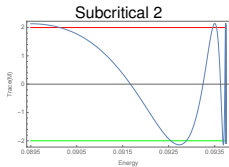
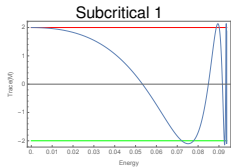
# Stability A3 Orbits

Specifically, we find numerically that

- ▶ For  $E < E_c$ ,  $\text{Tr } M$  oscillates between  $\pm 2$  with ever increasing frequency.
- ▶ Peak spacings vary geometrically as we approach  $E_c$  from below.
- ▶ For  $E > E_c$ ,  $\text{Tr } M$  oscillates between  $\pm 2$  with ever decreasing frequency.
- ▶ Peak spacings vary geometrically as we approach  $E_c$  from above.



# Stability of A3 Orbits



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- ▶ The bands/gaps have a *self-similar* structure as a function of energy.
- ▶ We can compute the *Feigenbaum constant* on either side of  $E_c$ .
- ▶ Define  $\nu = |1 - E_c/E|$ , look at ratios of bifurcation energies  $\nu_n/\nu_{n+2}$ .
  - ▶  $\nu_n/\nu_{n+2} \rightarrow \delta_1 = e^{-\pi\sqrt{\frac{2}{5}}}$  for  $E < E_c$ .
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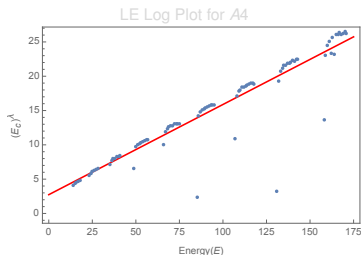
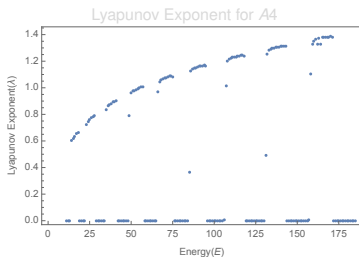
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- ▶ Non-zero Lyapunov exponent means chaotic dynamics.
- ▶ The Lyapunov exponents reflect the band structure.

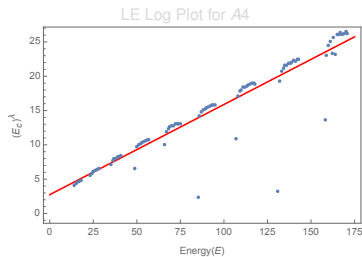
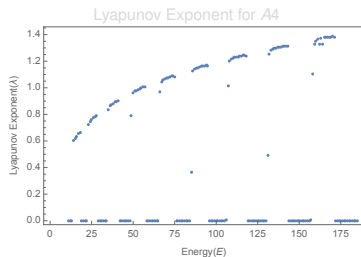


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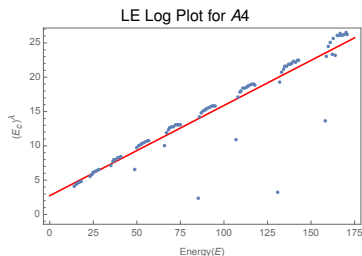
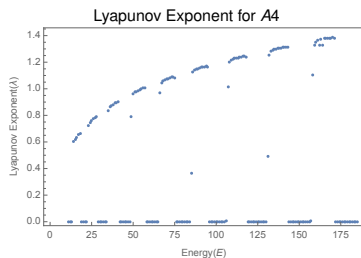


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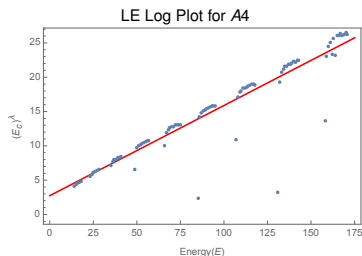
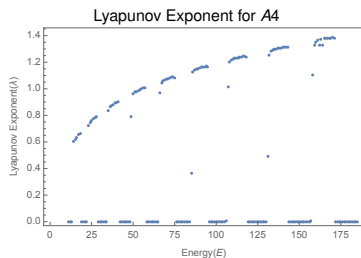


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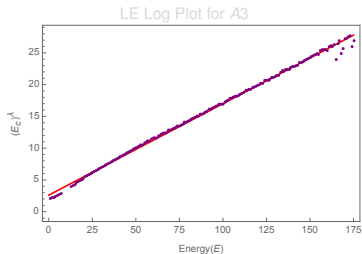
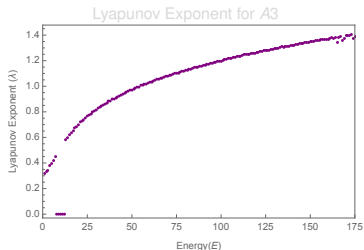
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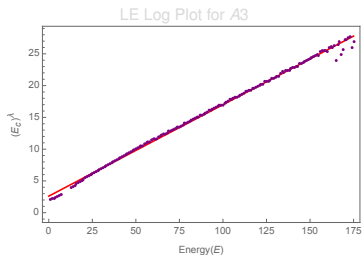
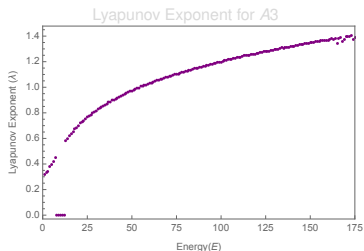


- ▶ Bands are wider, spaced apart by geometric progression.



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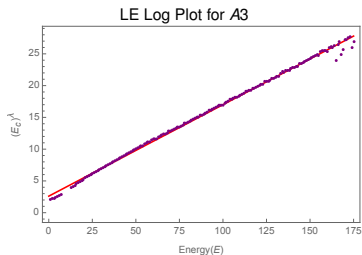
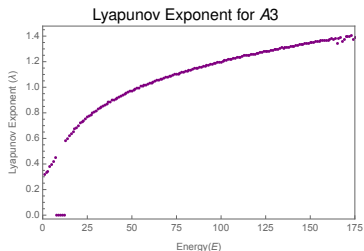


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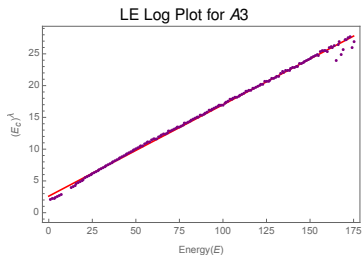
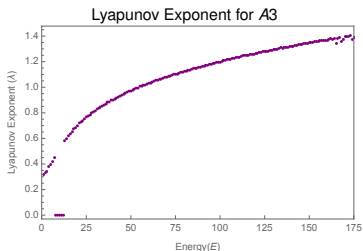


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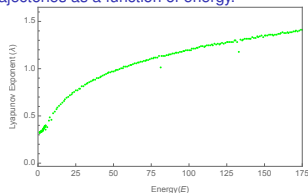


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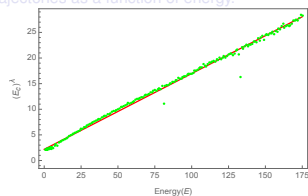
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Lyapunov Exponents for arbitrary trajectories as a function of energy.



- ▶ We can compute the Lyapunovs for trajectories with completely random initial conditions.

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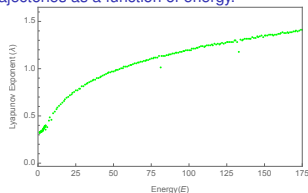


- ▶ The log plot is linear ( $y = 0.131465x + 2.27547$ ).



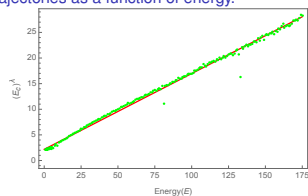
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- ▶ For arbitrary initial conditions, the trajectories are strongly chaotic ( $LE \neq 0$ ).
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- ▶ The  $A3$  and  $A4$  trajectories are special.
- ▶ Under a generic perturbation, even in the chaotic domain, the trajectories **do not** explore all available phase space.
- ▶ Rather, they are confined to particular corners.
- ▶ These trajectories *break ergodicity!*
- ▶ Numerical evidence: equipartition does **not** hold.
- ▶ Ergodicity breaking: different phases or glasses.
- ▶ There are at least two *corners* ( $A3$  and  $A4$ ) of phase space from which the dynamics cannot break out.



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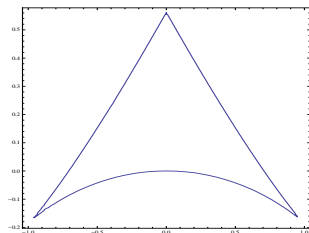


# Quantum Image of Ergodicity Breaking

- ▶ Remarkably, the quantum  $SU(2)$  model + quarks has quantum phases because of superselection sectors.

(Pandey-Vaidya 2017)

- ▶ Quarks get trapped in special corners of the gauge field space.

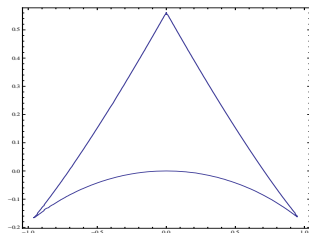


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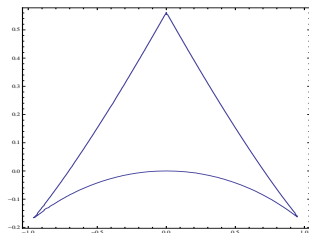


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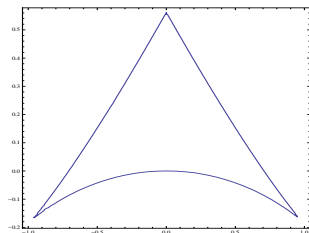


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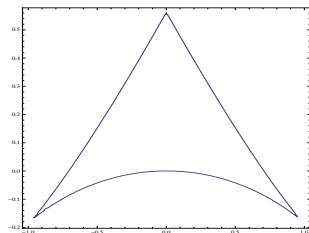


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# Summary

- ▶ The matrix model corresponds to the low energy sector of the Yang-Mills gauge theory.
- ▶ The  $SU(2)$  model has an unexpected tetrahedral symmetry.
- ▶ Its color-0 spin-0 sector is exactly the 3d Henon-Heiles system .
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## References

- ▶ M. S. Narasimhan and T. R. Ramadas: Comm. Math. Phys. 67, 121 136 (1979).
- ▶ A. P. Balachandran, A. R. de Queiroz and S. Vaidya: Int. J. Mod. Phys. A30, No. 9, 1550064 (2015).
- ▶ T. Iwai: J. Phys. A 43 415204 (2010), J. Phys. A 43 095206 (2010).
- ▶ K. Efstathiou and D. A. Sadovskii: Nonlinearity 17 415 (2004).
- ▶ V. L. Berdichevskii: Prikl. Matem. Mekhan. 52, 6, 947 (1988).
- ▶ M. Pandey and S. Vaidya: J. Math. Phys. 58, 022103 (2017).

