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Progress in the AdS_3 Landscape

arXiv: 1910.06326, 2011.00008 with Daniël Prins

arXiv: 2107.13562, 2203.09532 with Christopher Couzens & Niall Macpherson

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Anti-de Sitter

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- | By studying the space and properties of anti-de Sitter solutions we can gain insight into the space and properties of conformal field theories.
- | Vice versa, we can hope to shed light to characteristics of quantum gravity.

AdS₃/CFT₂ correspondence

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- | Conformal field theories in two dimensions feature a highly-constraining infinite-dimensional algebra of conformal transformations that often allows for their exact solution.
- | Gravity in three-dimensional asymptotically anti-de Sitter spacetime provides a toy model for quantum gravity.

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- | A way forward is to impose a symmetry on the background, at the expense of the size of the subspace of backgrounds one can access, depending on the degree of the symmetry.
- | We imposed supersymmetry, as (i) a technically simplifying assumption, (ii) a computational tool and (iii) a way out of swampland.

Layout

- I. Classification of minimally supersymmetric solutions.
- II.
- III.

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- II. Classification and construction of $\mathcal{N} = (2, 0)$ supersymmetric solutions.
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- I. Classification of minimally supersymmetric solutions.
- II. Classification and construction of $\mathcal{N} = (2, 0)$ supersymmetric solutions.
- III. Construction of $\mathcal{N} = (2, 2)$ supersymmetric solutions from D3-branes on Riemann surfaces.

Part I

Background

$$ds_{10}^2 = e^{2A} ds^2(\text{AdS}_3) + ds^2(M_7)$$

&

Φ, H, F_p

preserving the symmetries of AdS_3

Supersymmetry

$$\exists \epsilon_{1,2} : \delta_{\epsilon_{1,2}} \psi = 0 = \delta_{\epsilon_{1,2}} \lambda$$

Supersymmetry

$$\epsilon_1 = \zeta \otimes \chi_1 \otimes \begin{pmatrix} 1 \\ -i \end{pmatrix}, \quad \epsilon_2 = \zeta \otimes \chi_2 \otimes \begin{pmatrix} 1 \\ \pm i \end{pmatrix} ; \quad \nabla_\mu \zeta = \frac{1}{2} m \gamma_\mu \zeta$$

G-structure

$SO(7)$



G

stabilizer

G-structure

M_7 acquires a G-structure characterized by a set of tensors constructed as bilinears of $\{\chi_1, \chi_2\}$

$$G = \begin{cases} \text{SU}(3) : \{v, J, \Omega : v \lrcorner J = v \lrcorner \Omega = 0, \Omega \wedge J = 0\} \\ G_2 : \varphi \end{cases}$$

G-structure

The exterior derivatives of the G-structure tensors determine its intrinsic torsion, which is parametrized by torsion classes

$$dv = RJ + T_1 + \operatorname{Re}(\overline{V}_1 \lrcorner \Omega) + v \wedge W_0$$

$$dJ = \frac{3}{2} \operatorname{Im}(\overline{W}_1 \Omega) + W_3 + W_4 \wedge J + v \wedge \left(\frac{2}{3} \operatorname{Re} E J + T_2 + \operatorname{Re}(\overline{V}_2 \lrcorner \Omega) \right)$$

$$d\Omega = W_1 J \wedge J + W_2 \wedge J + \overline{W}_5 \wedge \Omega + v \wedge (E\Omega - 2V_2 \wedge J + S)$$

Supersymmetry Equations

$$\psi_+ + i\psi_- \equiv \chi_1 \otimes \chi_2^t$$

+

$$\chi_1 \otimes \chi_2^t \propto \sum_p \chi_2^t \gamma_{m_1 \dots m_p} \chi_1 \gamma^{m_1 \dots m_p}$$

+

$$\gamma^{m_1 \dots m_k} \rightarrow dx^{m_1} \wedge \dots \wedge dx^{m_k}$$

⇓

polyforms

$$\psi_{\pm}(v, J, \Omega \parallel \varphi)$$

Supersymmetry Equations

$$\delta_{\epsilon_{1,2}}\psi = 0 = \delta_{\epsilon_{1,2}}\lambda$$

⇓

constraints on $\{\chi_1, \chi_2\}$

⇓

$$d_H(e^{A-\Phi}\psi_{\mp}) = 0$$

$$d_H(e^{2A-\Phi}\psi_{\pm}) \mp 2me^{A-\Phi}\psi_{\mp} = \frac{1}{8}e^{3A} \star_7 \lambda F$$

$$(\psi_{\mp} \wedge \lambda F)_7 = \mp \frac{m}{2} e^{-\Phi} \text{vol}_7$$

Classification

- | We obtain a set of constraints on the intrinsic torsion of the G-structure and expressions for the supergravity fields in terms of the geometric data.
- | This allows for charting the AdS_3 landscape and the discovery of new solutions.

Classification

| A family of solutions for the strict $SU(3)$ -structure case, were examined in [AP, Prins '19]: the internal manifold M_7 is a $U(1)$ fibration over a conformally Kähler base, and they feature a varying axio-dilaton, a primitive $(2, 1)$ -form flux $H + ie^\Phi F_3$, and five-form flux F_5 .

$$\nabla^2(R - 2|\partial\Phi|^2) - \frac{1}{2}R^2 + R_{ij}R^{ij} + 2|\partial\Phi|^2R - 4R_{ij}\partial^i\Phi\bar{\partial}^j\Phi - \frac{8}{3}e^{-\Phi}H_{ijk}^{(2,1)}(H^{(1,2)})^{ijk} = 0$$

| The solutions of [Kim '05], [Donos, Gauntlett, Kim '08], [Benini, Bobev '13], [Benini, Bobev, Cricigno '15], [Couzens, Martelli, Schafer-Nameki '17], with $\mathcal{N} = (2, 0)$ supersymmetry, belong in this family.

GK geometries

[Gauntlett, Kim '07]

Y_{2n+1} consisting of a metric, a scalar function B and a closed two-form F .

The metric on Y_{2n+1} has a unit norm Killing vector ξ , defining a foliation \mathcal{F}_ξ of Y_{2n+1}

$$\xi = \frac{2}{n-2} \partial_z, \quad \eta = \frac{n-2}{2} (dz + P)$$

The metric on Y_{2n+1} then has the form

$$ds_{2n+1}^2 = \eta^2 + e^B ds_{2n}^2$$

where ds_{2n}^2 is a Kähler metric transverse to \mathcal{F}_ξ .

GK geometries

This Kähler metric, with transverse Kähler two-form J , Ricci two-form $\rho = dP$ and Ricci scalar R , determines all of the remaining fields.

$$e^B = \frac{(n-2)^2}{8} R, \quad F = -\frac{2}{n-2} J + d(e^{-B}\eta)$$

These off-shell geometries become solutions provided that the transverse Kähler metric satisfies the non-linear partial differential equation

$$\square R = \frac{1}{2} R^2 - R_{ij} R^{ij}.$$

One can define an extremal problem which is dual to c -extremization. [Couzens, Gauntlett, Martelli, Sparks '18]

Part II

$\mathcal{N} = (2, 0) \text{ AdS}_3 \times M_7$

$$\epsilon_1 = \sum_{I=1}^2 \zeta^I \otimes \chi_1^I \otimes \begin{pmatrix} 1 \\ -i \end{pmatrix}, \quad \epsilon_2 = \sum_{I=1}^2 \zeta^I \otimes \chi_2^I \otimes \begin{pmatrix} 1 \\ \pm i \end{pmatrix} ; \quad \nabla_\mu \zeta^I = \frac{1}{2} m \gamma_\mu \zeta^I$$

$$d_H(e^{A-\Phi} \Psi_{\mp}^{[IJ]}) = \mp \frac{c}{16} \delta^{IJ} F_{\pm}$$

$$d_H(e^{2A-\Phi} \Psi_{\pm}^{[IJ]}) \mp 2m e^{A-\Phi} \Psi_{\mp}^{[IJ]} = \frac{1}{8} e^{3A} \star_7 \lambda f_{\pm} \delta^{IJ}$$

$$d_H(e^{-\Phi} \Psi_{\pm}^{[IJ]}) = \frac{1}{16} \epsilon^{IJ} (\tilde{\xi} \wedge + \iota_{\xi}) F_{\pm}$$

$$d_H(e^{3A-\Phi} \Psi_{\mp}^{[IJ]}) = \pm \frac{1}{16} \epsilon^{IJ} (\tilde{\xi} \wedge + \iota_{\xi}) e^{3A} \star_7 \lambda F_{\pm}$$

where

$$\Psi^{IJ} \equiv \chi_1^I \otimes \chi_2^{J\dagger}$$

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- | The six-dimensional space transverse to the Killing vector supports an $SU(2)$ -structure characterized by (z, j_2, ω_2) . The $SU(2)$ -structure “lives” on a four-dimensional subspace which is complex.

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- | The six-dimensional space transverse to the Killing vector supports an $SU(2)$ -structure characterized by (z, j_2, ω_2) . The $SU(2)$ -structure “lives” on a four-dimensional subspace which is complex.
- | The problem of finding solutions to the equations of motion consists of solving PDEs coming from the Bianchi identities for the p-form fields.

$$\mathcal{N} = (2, 0) \text{ AdS}_3 \times M_7$$

In order to make progress we have imposed an additional isometry. The metric of the internal space reads

$$ds_7^2 = \frac{e^{2A}}{4m^2} (d\psi + \mathcal{A})^2 + e^{\Phi-3A} \left[e^{2u} D\varphi^2 + e^{\Phi-A} dy^2 + e^{\frac{1}{2}(5A-\Phi)-u} ds^2(M_4) \right]$$

with the metric on M_4 Kähler at fixed y coordinate.

$$\mathcal{N} = (2, 0) \text{ AdS}_3 \times M_7$$

Taking

$$ds^2(M_4) = e^{2f_1(y)} ds^2(\Sigma_1) + e^{2f_2(y)} ds^2(\Sigma_2)$$

we have solved the differential equations coming from the Bianchi identities and found explicit solutions: (i) one class for non-zero Romans mass (ii) two classes for zero Romans mass

$$\mathcal{N} = (2, 0) \text{AdS}_3 \times M_7$$

$$ds^2(M_4(\vec{x}, y)) = e^{2f(y)} ds^2(\mathcal{M}_4(\vec{x})),$$

$$ds^2 = e^{2A} \left[ds^2(\text{AdS}_3) + \frac{1}{4m^2} (d\psi + P_K)^2 + (d\varphi + t_1 \Sigma)^2 + e^{\Phi-5A} \left(e^{\Phi-A} dy^2 + ds^2(\mathcal{M}_4) \right) \right]$$

$$e^{-4A} = \frac{R_K}{8m^2} \sqrt{2f_0 y + c}, \quad e^{-2\Phi} = \frac{(2f_0 y + c)^{5/4} \sqrt{R_K}}{2\sqrt{2}m},$$

where the metric on \mathcal{M}_4 is Kähler and satisfies the master equation

$$\square_K R_K - \frac{1}{2} R_K^2 + R^{mn} R_{mn} = 8m^2 t_1^2 |d\Sigma|^2.$$

For $t_1 = 0$ we have a five-dimensional GK geometry and an extremal problem can be setup for the calculation fo the central charge.

Part III

Compactifying Higher Dimensional Field Theories

- | Large classes of theories in lower dimensions.
- | Their properties admit a description in terms of the geometry and topology of the compact manifold.

D3-branes on a Riemann surface

| We considered D3-branes compactified on a Riemann surface with a twist, preserving $\mathcal{N} = (2, 2)$ supersymmetry and flowing to a two-dimensional SCFT.

| $SU(4) \rightarrow U(1)^3 = U(1)_L \times U(1)_R \times U(1)_F$

D3-branes on a Riemann surface

$$S^5_{U(1)^3} \rightarrow M_7$$
$$\downarrow$$
$$\Sigma_g$$

D3-branes on a Riemann surface

[Couzens, Martelli, Schafer-Nameki '17]

$$\frac{1}{L^2} ds^2 = \frac{\sqrt{y}}{\sin \zeta} [ds^2(\text{AdS}_3) + ds^2(X_7)]$$

$$ds^2(X_7) = \cos^2 \zeta (d\psi_1 + \sigma)^2 + \sin^2 \zeta d\psi_2^2 + \frac{\sin^2 \zeta}{4y^2 \cos^2 \zeta} dy^2 + \frac{\sin^2 \zeta}{y} g^{(4)}(y, x)_{ij} dx^i dx^j$$

| SU(2)-structure (J, Ω) on the four-dimensional base; $g^{(4)}$ is Kähler

| $R_L = \partial_{\psi_1} - \partial_{\psi_2}$, $R_R = \partial_{\psi_1} + \partial_{\psi_2}$

| The geometry is supported by F_5

D3-branes on a Riemann surface

- | Assumption: $g^{(4)}$ contains an additional flavour $U(1)$
- | We have reduced the torsion conditions under this assumption. The solution is determined by a potential D satisfying a

$$(\partial_{X_1}^2 + \partial_{X_2}^2)D = 16y^2 \left(\partial_y^2 D \partial_\Theta D - (\partial_y \partial_\Theta D)^2 \right) e^{\partial_y D}$$

D3-branes on a Riemann surface

$$\frac{1}{L^2} ds^2 = \sqrt{\frac{yg}{h}} \left[ds^2(\text{AdS}_3) + \frac{h}{g} d\psi_2^2 + \frac{he^{2A}}{yg} (dX_1^2 + dX_2^2) + \frac{h}{yg} ds^2(\mathcal{M}_4) \right],$$

$$ds^2(\mathcal{M}_4) = \frac{1}{4} g_{ij} du^i du^j + h^{ij} \eta_i \eta_j, \quad e^{2A} = 4y^2 e^{\partial_y D} g,$$

$$g_{ij} \equiv -\partial_i \partial_j D, \quad h_{ij} \equiv -\partial_i \partial_j (D + y(\log y - 1)),$$

$$\eta_1 \equiv d\psi_1 + \star_2 d_2(\partial_y D), \quad \eta_2 \equiv \frac{1}{2} (d\phi + \star_2 d_2(\partial_\Theta D)),$$

with

$$u^i = \{y, \Theta\}, \quad g \equiv \det(g_{ij}), \quad h \equiv \det(h_{ij}).$$

D3-branes on a Riemann surface

★ similar system for D4–D8/O8 [Bah, AP, Weck '18] and M5 –branes [Bah '15] on Riemann surfaces

D3-branes on a Riemann surface

Riemann surface of constant curvature:

$$e^{2A} = f(y, \Theta) e^{2A_0(X_1, X_2)}$$

$$\frac{1}{L^2} ds^2 = \sqrt{\Lambda} [ds^2(\text{AdS}_3) + 4e^{4\nu} c^+ c^- e^{2A_0} (dX_1^2 + dX_2^2)] + \frac{1}{\sqrt{\Lambda}} ds^2(\mathcal{M}_5),$$

$$ds^2(\mathcal{M}_5) = d\mu_0^2 + \frac{1}{c^+} d\mu_+^2 + \frac{1}{c^-} d\mu_-^2 + \frac{1}{c^+} \mu_+^2 \eta_+^2 + \frac{1}{c^-} \mu_-^2 \eta_-^2 + \mu_0^2 d\psi_2^2,$$

$$\Lambda \equiv \mu_0^2 + \frac{(m^+)^2}{c^+} \mu_+^2 + \frac{(m^-)^2}{c^-} \mu_-^2,$$

$$\eta_{\pm} \equiv \frac{1}{2} [(1 \pm \epsilon) d\psi_1 \pm d\phi + V], \quad V = \kappa (\partial_{X_2} \tilde{A}_0 dX_1 - \partial_{X_1} \tilde{A}_0 dX_2).$$

$g = 0$: non-compact $g = 1$: $\text{AdS}_3 \times S^2 \times T^4$ $g > 1$: compact

D3-branes on a Riemann surface

$$c_{\text{sugra}} = \frac{3}{2G_N^{(3)}} = \frac{c_L + c_R}{2} = 3N^2(g - 1)$$

D3-branes on a Riemann surface

topological disc of non-constant curvature

$$\frac{1}{L^2} ds^2 = \sqrt{W} H(x)^{\frac{1}{3}} [ds^2(\text{AdS}_3) + ds^2(\Sigma_2)] + \frac{1}{\sqrt{W}} \sum_{I=1}^3 (X^I)^{-1} [d\mu_I^2 + \mu_I^2 (d\phi_I + A_I)^2],$$

$$W = \sum_{I=1}^3 X^I \mu_I^2, \quad A_I = \frac{x - x_0}{x + 3K_I} d\varphi, \quad X^I = \frac{H(x)^{\frac{1}{3}}}{x + 3K_I},$$

$$ds^2(\Sigma_2) = \frac{1}{4P(x)} dx^2 + \frac{P(x)}{H(x)} d\varphi^2,$$

where μ_I embed a unit radius two-sphere into \mathbb{R}^3 and the functions of x are

$$H = (x + 3K_1)(x + 3K_2)(x + 3K_3), \quad P = H - (x - x_0)^2, \quad K_1 = K_2 = K, \quad K_3 = -\frac{1}{3}x_0$$

Limiting case of [\[Boido, Ipiña, Sparks '21\]](#)

D3-branes on a Riemann surface

| $x = x_- : \mathbb{R}^2/\mathbb{Z}_k$

| $x = x_0 : \text{regular}$

| $(x, \mu_3) = (x_0, 0) : \text{flavour D3-branes smeared over } S^3$

$$c_{\text{sugra}} = 3N^2 \frac{M^2}{4k(1+2M)}$$

Future Directions

- | Study the field theory of D3-branes on a topological disc and reproduce the holographic central charge.
- | Find topological disc solutions for the D4-O8/D8 and M5 -brane configurations.
- | Classify & construct $\mathcal{N} = (2, 0)$ solutions in the “time-like class”.
- | A generalized geometry formulation of the master equation and the gravity dual of c-extremization (beyond GK geometries).

The End. Thank you!