

Localization and Reference Frames in κ -Minkowski Spacetime

Corfù Summer Institute: “Workshop on Quantum Geometry,
Field Theory and Gravity ”

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This talk is based on...

PRD

“Localization and Reference Frames in κ -Minkowski Spacetime”

(F.Lizzi, M. Manfredonia, F. Mercati, T. Poulain)

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Figure: Scan to open PDF version of the paper

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κ -Minkowski space time

Q.G. suggests classical space-time (at basis of GR and QFT) to be replaced with quantum structure.

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Non Commutative Geometry



κ -Minkowski

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Non Commutative Geometry



κ -Minkowski

$$[x^0, x^i] = i\lambda x^i, \quad [x^i, x^j] = 0,$$

$$i, j \in \{1, 2, 3\}$$

where $x^0 = ct$, sometimes $\lambda = 1/\kappa$, hence the name

κ -Minkowski **not invariant** under Poincaré

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Invariant under **non commutative generalization**: κ -Poincaré

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Invariant under **non commutative generalization**: κ -Poincaré

κ -Poincaré, κ -Minkowski \Rightarrow group manifold becomes a n.c. space



Principle of Non Localizability

$$\Delta x^0 \Delta x^i \geq \frac{\lambda}{2} |\langle x^i \rangle|$$

Somehow similar to Heisenberg principle

$$\Delta p_i \Delta q^j \geq \frac{\hbar}{2}$$

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Operator Representation for κ -Minkowski

We introduce the following representation as operators on $L^2(\mathbb{R}^3)$

$$\hat{x}^i \psi(x) = x^i \psi(x),$$

$$\hat{x}^0 \psi(x) = i\lambda \left(\sum_i x^i \partial_{x^i} + \frac{3}{2} \right) \psi(x) = i\lambda \left(r \partial_r + \frac{3}{2} \right) \psi(x).$$

which is compatible with κ -Minkowski commutators.

Caution!

Other representation are allowed. We want to consider those in future works.

The Time Eigenvalue Equation

Since \hat{x}^0 commutes with all spherical harmonics, we adopt polar basis (r, θ, φ) and we solved **eigenvalue equation**.

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$$i\lambda \left(r\partial_r + \frac{3}{2} \right) r^\alpha = i\lambda \left(\alpha + \frac{3}{2} \right) r^\alpha = \ell_\alpha r^\alpha, \quad \ell_\alpha = i\lambda \left(\alpha + \frac{3}{2} \right)$$

We want eigenvalues to be real!

$$\{\ell_\alpha \in \mathbb{R}\} \Leftrightarrow \left\{ \alpha = -\frac{3}{2} - i\tau, \tau \in \mathbb{R} \right\}$$

The adimensional τ is related to time $t = \frac{\lambda}{c}\tau$

Between Time and Space

Eigenfunction of Time

$$T_\tau[r] = \frac{r^{-\frac{3}{2}-i\tau}}{\lambda^{-i\tau}} = r^{-\frac{3}{2}} e^{-i\tau \log\left(\frac{r}{\lambda}\right)}$$

Between Time and Space

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$$T_\tau[r] = \frac{r^{-\frac{3}{2}-i\tau}}{\lambda^{-i\tau}} = r^{-\frac{3}{2}} e^{-i\tau \log(\frac{r}{\lambda})}$$

Self-adjointness of operators allow us to expand $\psi(r, \theta, \varphi)$ on functions $\{T_\tau[r]\}$

$$\psi(r, \theta, \varphi) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} d\tau r^{-\frac{3}{2}} e^{-i\tau \log(\frac{r}{\lambda})} \tilde{\psi}(\tau, \theta, \varphi)$$

Between Time and Space

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Of course this is not a Fourier transform...

...it is a **Mellin** transform!

$$\psi(r, \theta, \varphi) = \int_{-\infty}^{\infty} \frac{d\tau}{\sqrt{2\pi}} r^{-\frac{3}{2}} e^{-i\tau \log(\frac{r}{\lambda})} \tilde{\psi}(\tau, \theta, \varphi) = \mathcal{M}^{-1} \left[\tilde{\psi}(\tau, \theta, \varphi), r \right]$$

$$\tilde{\psi}(\tau, \theta, \varphi) = \int_0^{\infty} \frac{dr}{\sqrt{2\pi}} r^{\frac{1}{2}} e^{i\tau \log(\frac{r}{\lambda})} \psi(r, \theta, \varphi) = \mathcal{M} \left[\psi(r, \theta, \varphi), \frac{3}{2} + i\tau \right]$$

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Good News: It is still an isometry!

The above transformation preserves the norm

$$\int_0^{\infty} dr r^2 |\psi(r)|^2 = \int_{-\infty}^{\infty} d\tau |\tilde{\psi}(\tau)|^2$$

Physical Interpretation

Our Conjecture

We assume the usual theory of measurements and observables to stand in this framework.

As an example, the average time measured by an observer would be

$$\langle \hat{x}^0 \rangle_\psi = 4\pi \int r^2 dr \bar{\psi}(r) i\lambda \left(r\partial_r + \frac{3}{2} \right) \psi(r)$$

Points at the origin of space

We recall space-time uncertainty

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NOTE

$\langle x^i \rangle$ on the right hand side suggests **spatial localized states** to be possible at spatial origin.

We have to search for states that saturate uncertainty bound.

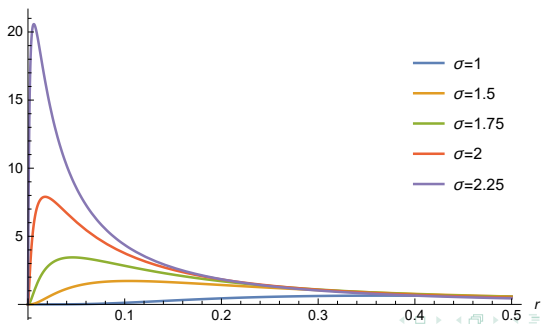
~~Gaussian~~ \rightarrow log-Gaussian

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log-Gaussians play the role of *coherent states*

$$L(r, r_0) = Ne^{-\frac{(\log r - \log r_0)^2}{\sigma^2}}$$

$L(r, r_0) |_{r_0 = \exp(-\sigma^2 \sigma^1)}$



The average values of r^n and $(\hat{x}^0)^n$ are

$$\langle \hat{r}^n \rangle_L = e^{\frac{\sigma^2}{8} n(n+6)} r_0^n; \quad \langle (\hat{x}^0)^n \rangle_L = \frac{1}{4\pi} \left(\frac{\lambda}{\sigma}\right)^n \begin{cases} 0 & n \text{ odd} \\ (n-1)!! & n \text{ even} \end{cases}$$

Eigenstate of the origin

There is a localized state $|o\rangle$ localized at origin ($\tau = 0, r = 0$).

Multiplication by $r^{i\tau'}$ shifts time

We have a one parameter family of states localized at spatial origin at different time

$$\{|o\rangle_\tau\} := \{|o\rangle_\tau \text{ is localized at } r = 0, \tau \neq 0\}$$

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κ -Poincaré quantum group

The noncommutative algebra of functions \mathcal{P}_κ , generated by $\Lambda^\mu{}_\nu$ and a^μ that leave κ -Minkowski commutation relations invariant under the transformation:

$$x^\mu \rightarrow x'^\mu = \Lambda^\mu{}_\nu \otimes x^\nu + a^\mu \otimes 1.$$

$$[x'^\mu, x'^\nu] = i\lambda (\delta^\mu{}_0 x'^\nu - \delta^\nu{}_0 x'^\mu)$$

κ -Poincaré commutation relations

$$[a^\mu, a^\nu] = i\lambda (\delta^\mu{}_0 a^\nu - \delta^\nu{}_0 a^\mu), \quad [\Lambda^\mu{}_\nu, \Lambda^\rho{}_\sigma] = 0,$$

$$[\Lambda^\mu{}_\nu, a^\rho] = i\lambda [(\Lambda^\mu{}_\sigma \delta^\sigma{}_0 - \delta^\mu{}_0) \Lambda^\rho{}_\nu + (\Lambda^\sigma{}_\nu \delta^0{}_\sigma - \delta^0{}_\nu) \eta^{\mu\rho}].$$

Hopf Algebra Structure

Coproduct $\Delta : \mathcal{P}_\kappa \rightarrow \mathcal{P}_\kappa \otimes \mathcal{P}_\kappa$:

$$\Delta(a^\mu) = a^\nu \otimes \Lambda^\mu_\nu + 1 \otimes a^\mu, \quad \Delta(\Lambda^\mu_\nu) = \Lambda^\mu_\rho \otimes \Lambda^\rho_\nu,$$

Antipode $S : \mathcal{P}_\kappa \rightarrow \mathcal{P}_\kappa$

$$S(a^\mu) = -a^\nu (\Lambda^{-1})^\mu_\nu, \quad S(\Lambda^\mu_\nu) = (\Lambda^{-1})^\mu_\nu,$$

Counit $\varepsilon : \mathcal{P}_\kappa \rightarrow \mathbb{C}$,

$$\varepsilon(a^\mu) = 0, \quad \varepsilon(\Lambda^\mu_\nu) = \delta^\mu_\nu,$$

Have to be homomorphisms with respect to the commutation relations

A representation for κ - Poincaré

Long story short, a representation on $L^2(\mathfrak{so}(3, 1))$ turns out to be non faithful. Thus we enlarged the Hilbert space with 3 more coordinates $q^i \in \mathbb{R}^3: L^2(SO(3, 1)) \Rightarrow L^2(\mathfrak{so}(3, 1) \times \mathbb{R}^3)$

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A faithful (although complicated) representation

$$a^\rho \phi = i\lambda \delta^\rho_0 \left(\frac{3}{2} \phi(q, \omega) + q^i \frac{\partial \phi}{\partial q^i} \right) + \delta^\mu_i q^i \phi$$

$$- i\lambda : [(\Lambda^\mu_\sigma \delta^\sigma_0 - \delta^\mu_0) \Lambda^\rho_\nu + (\Lambda^\sigma_\nu \delta^0_\sigma - \delta^0_\nu) \eta^{\mu\rho}] \Lambda^\nu_\alpha \frac{\partial}{\partial \omega^\mu_\alpha} : \phi,$$

$$\Lambda^\mu_\nu \phi = \Lambda^\mu_\nu(\omega) \phi = (\exp \omega)^\mu_\nu \phi,$$

where $\phi = \phi(q, \omega)$ with $q \in \mathbb{R}^3$ and $\omega \in \mathfrak{so}(3, 1)$.

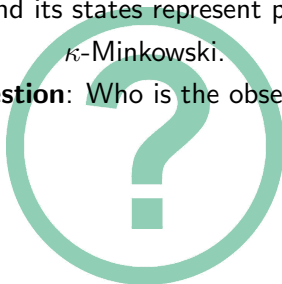
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Observers and Reference Frames

Algebra of \hat{x}^i, \hat{x}^0 and its states represent position and time in κ -Minkowski.

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Question: How do we change the observer?

State of the observer!

Algebra of a 's and Λ 's represent the state of a translated and Lorentz transformed observer

Identity Transformation state

We define the state $|o\rangle_{\mathcal{P}}$ of \mathcal{P}_{κ} such that

$$\mathcal{P}\langle o|f(a, \Lambda)|o\rangle_{\mathcal{P}} = \varepsilon(f)$$

The state returns the value of the function on the identity transformation.

It describes the Poincaré transformation between two coincident observers

Physical Interpretation

Consider two different observers

Alice

(at the origin)

$$\mathbb{1} \otimes x^\nu$$

Bob

(in relative motion)

$$\Lambda^\mu{}_\nu \otimes x^\nu + a^\mu \otimes \mathbb{1}$$

both coordinates are $\hat{x}^\mu, \hat{x}'^\mu \in \mathcal{P}_\kappa \otimes \mathcal{M}_\kappa$



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We take direct sum of the representation of \mathcal{P}_κ and \mathcal{M}_κ .



States of a transformed observer

$$\begin{array}{ll} |g\rangle \in L^2(SO(3,1) \times \mathbb{R}_q^3) & |\psi\rangle \in L^2(\mathbb{R}^3) \\ \text{(related to the observer)} & \text{(related to the event)} \end{array}$$

Generic element of Hilbert space

$$|g, \psi\rangle = |g\rangle \otimes |\psi\rangle$$

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Expectation values measured by Bob are related to those measured by Alice through relation

$$\begin{aligned} \langle x'^{\mu} \rangle &= \langle g | \otimes \langle \psi | (\Lambda^{\mu}_{\nu} \otimes x^{\nu} + a^{\mu} \otimes 1) | g \rangle \otimes | \psi \rangle \\ &= \langle g | \Lambda^{\mu}_{\nu} | g \rangle \langle \psi | x^{\nu} | \psi \rangle + \langle g | a^{\mu} | g \rangle \end{aligned}$$

(averages transforms with averaged κ -Poincaré transformation rule)

Example: Transforming the origin state

The origin state for Alice is:

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...and for Bob:

$$|g, o\rangle = |g\rangle \otimes |o\rangle$$

We want to know what Bob measures with the coordinates centred on his reference frame

$$\langle x'^{\mu} \rangle = \langle g | \Lambda^{\mu}_{\nu} | g \rangle \langle o | x^{\nu} | o \rangle + \langle g | a^{\mu} | g \rangle \langle o | o \rangle,$$

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$$\langle x'^{\mu} \rangle = \langle g | a^{\mu} | g \rangle$$

The expectation value of coordinates x' is fully determined by the expectation value of translation a

This is natural: the different observers are comparing positions, not directions.

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Conclusions and Outlooks

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- We adopted observables interpretation and discussed physical meaning
- We gave representation of κ -Poincaré group as operators on Hilbert space $L^2(SO(3, 1) \times \mathbb{R}_q^3)$
- We implemented all the above results to introduce transformations between different observers
- We developed a new framework based on a representation of $\mathcal{P}_\kappa \otimes M_\kappa$ as operators on $L^2(SO(3, 1) \times \mathbb{R}_q^3 \times \mathbb{R}^3)$ in which a state has informations about both the event and the observer who is staring at it.

Future Perspectives

- Introduce the whole phase space or at least some notion of **momentum space**

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- Introduce the whole phase space or at least some notion of **momentum space**
- Introduce **dynamics** and evolution parameter for observers and events
- Better understand the role of \hat{x}_0 and time t in the dynamics.



**THANK YOU
FOR
YOUR
ATTENTION!
ANY QUESTIONS?**