

Classical dynamics on fuzzy space¹

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¹FG Scholtz, Phys. Rev. D 98 (2018) 104058

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- 2 Quantum Mechanics on Fuzzy Space
- 3 Classical Dynamics
 - Path Integral Action
 - Equations and Constants of Motion
 - Features of the Equations of Motion
- 4 Underpinning Lorentz Geometry
- 5 Concluding Remarks

- The fuzzy sphere commutation relations are

$$[\hat{x}_i, \hat{x}_j] = 2i\lambda\varepsilon_{ijk}\hat{x}_k.$$

where λ is the non-commutative length parameter. These commutation relations respect the rotational symmetry. The Casimir operator $\hat{x}^2 = \hat{x}_i\hat{x}_i$ is associated with the square of the radial distance and its eigenvalues are determined by the $su(2)$ representation under consideration: $j(j+1)$, $j = 0, 1/2, \dots$

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- Fuzzy space is the collection of fuzzy spheres with each allowed radius appearing once.

- A concrete realisation of fuzzy space is provided by the Schwinger construction, which utilises two sets of boson creation and annihilation operators to build a representation of $su(2)$:

$$[\hat{a}_\alpha, \hat{a}_\beta^\dagger] = \delta_{\alpha\beta} \quad \text{and} \quad [\hat{a}_\alpha, \hat{a}_\beta] = [\hat{a}_\alpha^\dagger, \hat{a}_\beta^\dagger] = 0, \quad \alpha, \beta = 1, 2.$$

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- The Casimir operator reads $\hat{x}^2 = \hat{x}_i \hat{x}_i = \lambda^2 \hat{n}(\hat{n} + 2)$ with $\hat{n} = \hat{a}_1^\dagger \hat{a}_1 + \hat{a}_2^\dagger \hat{a}_2$ from which it is clear that each $su(2)$ representation occurs precisely once.

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- We denote this realisation of fuzzy space by \mathcal{H}_{FS} .

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- The quantum Hilbert space \mathcal{H}_Q is defined as the algebra of bounded operators on \mathcal{H}_{FS} generated by the coordinates (the operators that commute with \hat{x}^2)² and have a finite norm with respect to a weighted Hilbert-Schmidt inner product³:

$$\mathcal{H}_Q = \left\{ \psi : [\psi, \hat{n}] = 0, \text{tr}_{FS} (\psi^\dagger \hat{r} \psi) < \infty \right\}.$$

Here tr_{FS} denotes the trace over \mathcal{H}_{FS} and $\hat{r} = \lambda(\hat{n} + 1)$. States in \mathcal{H}_Q are denoted $|\psi\rangle$.

²N Chandra et al, J.Phys. A: Math.Theor 47 (2014) 445203

³The choice of weight is motivated by the requirement that the projector that projects on all states in \mathcal{H}_{FS} with radius less than or equal to R (the analogue of the characteristic function of a ball with radius R) gives the volume of a sphere of radius R for large R (V. Gáliková and P. Prešnajder, 2013 *J. Math. Phys.* **54** 052102) □

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- Quantum observables are identified with self-adjoint operators acting on \mathcal{H}_Q . These include:

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- From here it is standard quantum mechanics.
- Quantum observables are identified with self-adjoint operators acting on \mathcal{H}_Q . These include:
 - The coordinates that act through left multiplication as

$$\hat{X}_i|\psi\rangle = |\hat{x}_i\psi\rangle,$$

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 - The coordinates that act through left multiplication as

$$\hat{X}_i|\psi\rangle = |\hat{x}_i\psi\rangle,$$

- The angular momentum operators which act adjointly according to

$$\hat{L}_i|\psi\rangle = |\frac{\hbar}{2\lambda}[\hat{x}_i, \psi]\rangle \quad \text{with} \quad [\hat{L}_i, \hat{L}_j] = i\hbar\varepsilon_{ijk}\hat{L}_k.$$

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- The non-commutative analogue of the Laplacian, which gives the kinetic energy, is defined as

$$\hat{\Delta}_\lambda|\psi\rangle = |-\frac{1}{\lambda\hat{r}}[\hat{a}_\alpha^\dagger, [\hat{a}_\alpha, \psi]]\rangle.$$

It commutes with the angular momenta and is symmetric on \mathcal{H}_Q .

- The Hamiltonian is given by

$$\hat{H} = -\frac{\hbar^2}{2m}\hat{\Delta} + V(\hat{R})$$

with \hat{R} the radius operator that acts as

$$\hat{R}|\psi\rangle = |\lambda(\hat{n} + 1)\psi\rangle, \hat{n} = \mathbf{a}_\alpha^\dagger \mathbf{a}_\alpha.$$

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- The angular momentum operators commute with the Hamiltonian and are therefore conserved, but there is a further important conserved quantity, which is the operator $\hat{\Gamma}$ that acts as follows

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$$\hat{\Gamma}|\psi\rangle = |[\mathbf{a}_\alpha^\dagger \mathbf{a}_\alpha, \psi]\rangle.$$

- It can easily be checked that Γ commutes with the Hamiltonian. Note that $\hat{\Gamma}|\psi\rangle = 0, \forall \psi \in \mathcal{H}_Q$.

- The formalism above has been used to solve the free particle and particle in a well ⁴, to develop scattering theory on non-commutative spaces⁵ and the statistical physics⁶. The main results are:

⁴N Chandra et al, J.Phys. A: Math.Theor 47 (2014) 445203

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- The free particle spectrum is given by

$$E_{\vec{k}} = \frac{2\hbar^2}{m\lambda^2} \sin^2 \left(\frac{|\vec{k}|\lambda}{2} \right) \leq \frac{2\hbar^2}{m\lambda^2}, \quad |\vec{k}| \in [0, \pi/\lambda)$$

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- For $|\vec{k}|\lambda \ll 2$

$$E_{\vec{k}} = \frac{\hbar^2 |\vec{k}|^2}{2m}$$

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- For $|\vec{k}|\lambda \ll 2$

$$E_{\vec{k}} = \frac{\hbar^2 |\vec{k}|^2}{2m}$$

- Each single particle state occupies a finite volume $V_0 = 4\pi\lambda^3$. For Fermions this is an excluded volume.

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- The resulting equation of state has striking consequences for the mass-radius relationship of a white dwarf:

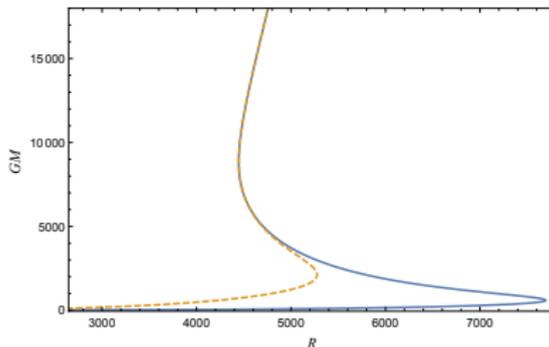


Figure: Mass-radius relationship for white dwarf at two temperatures in arbitrary units.

From Quantum to Classical

- Now that we have a quantum theory with a short length scale, manifest rotational symmetry and the appropriate low energy limit, we may ask what is the underpinning classical dynamics, i.e. how may Newton dynamics be altered? To answer this question, we must compute the path integral action and extract the equations of motion.

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- To do this, we enlarge the quantum Hilbert space \mathcal{H}_q to include all Hilbert-Schmidt operators acting on \mathcal{H}_{FS} and not just those commuting with the Casimir. We denote this enlarged space by \mathcal{H}_q^0 . Clearly $\mathcal{H}_q \subset \mathcal{H}_q^0$. From the definition of \mathcal{H}_q , states that belong to the subspace \mathcal{H}_q must satisfy the constraint

$$\hat{\Gamma}|\psi\rangle = 0.$$

Note that since $\hat{\Gamma}$ is conserved, initial states that satisfy this condition, will do so at all times.

- Let $|\ell\rangle$ be a set of overcomplete coherent states, i.e.

$$\int d\mu(\ell) |\ell\rangle \langle \ell| = \mathbf{1},$$

then the transition amplitude can be represented as a path integral

$$\langle \ell_f, t_f | \ell_i, t_i \rangle = \int_{\ell(t_i)=\ell_i}^{\ell(t_f)=\ell_f} [d\mu(\ell)] e^{\frac{i}{\hbar} S},$$

with action

$$S = \int_{t_i}^{t_f} dt \langle \ell(t) | i\hbar \frac{\partial}{\partial t} - H | \ell(t) \rangle.$$

- We can easily construct a set of coherent states on \mathcal{H}_q^0 :

$$|z_\alpha, w_\alpha\rangle = |z_\alpha\rangle\langle w_\alpha|.$$

where $|z_\alpha\rangle$ is a Glauber coherent state on \mathcal{H}_c and

$$\int \frac{d\bar{z}_\alpha dz_\alpha d\bar{w}_\alpha dw_\alpha}{\pi^4} |z_\alpha, z_\alpha\rangle\langle z_\alpha, w_\alpha| = \mathbf{1}_q.$$

Note though that in general $\Gamma|z_\alpha, w_\alpha\rangle \neq 0$. However, if we want to compute transition amplitudes between states that satisfy $\Gamma|\psi\rangle = 0$, we can safely use them to insert the identity at intermediate times in a time slicing procedure as Γ is conserved. Indeed, if this is done, the Γ must appear as a conserved quantity in the resulting action and we must simply require it to vanish to satisfy the condition of physicality of the initial state.

- The general result for the path integral action can now be applied to obtain

$$S = \int_{T_i}^{T_f} dT \left[\frac{i}{2} \left(\bar{z}_\alpha \dot{z}_\alpha - \dot{\bar{z}}_\alpha z_\alpha + \bar{w}_\alpha \dot{w}_\alpha - \dot{\bar{w}}_\alpha w_\alpha \right) - H \right]$$

Here

$$H = (f_1(R) \bar{z}_\alpha z_\alpha - f_2(R) (\bar{z}_\alpha w_\alpha + z_\alpha \bar{w}_\alpha) + f_3(R) \bar{w}_\alpha w_\alpha) + W(R).$$

■ with

$$R = \bar{z}_\alpha z_\alpha$$

$$f_1(R) = \frac{1}{2} \langle z_\alpha | \frac{1}{\hat{n} + 2} | z_\alpha \rangle,$$

$$f_2(R) = \frac{1}{2} \langle z_\alpha | \frac{1}{\sqrt{(\hat{n} + 1)(\hat{n} + 2)}} | z_\alpha \rangle,$$

$$f_3(R) = \frac{1}{2} \langle z_\alpha | \frac{1}{\hat{n} + 1} | z_\alpha \rangle,$$

$$W(R) = \frac{1}{e_0} \langle z_\alpha | V(\hat{R}) | z_\alpha \rangle + 2f_3(R) \equiv \tilde{V}(R) + 2f_3(R).$$

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- Here we introduced the time and energy scales

$$t_0 = \frac{m\lambda^2}{\hbar}, \quad e_0 = \frac{\hbar}{t_0},$$

and the dimensionless quantities

$$T = \frac{t}{t_0}, \quad X_i = \frac{x_i}{\lambda}, \quad E = \frac{e}{e_0}.$$

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- The equation of motion are easily found:

$$\dot{z}_\alpha = -i \frac{\partial H}{\partial \bar{z}_\alpha},$$

$$\dot{\bar{z}}_\alpha = i \frac{\partial H}{\partial z_\alpha},$$

$$\dot{w}_\alpha = i \frac{\partial H}{\partial \bar{w}_\alpha},$$

$$\dot{\bar{w}}_\alpha = -i \frac{\partial H}{\partial w_\alpha}.$$

- There are five constants of motion

$$\begin{aligned}\Gamma &= \bar{z}_\alpha z_\alpha - \bar{w}_\alpha w_\alpha, \\ L_j &= \bar{z}_\alpha \sigma_{\alpha\beta}^{(j)} z_\beta - \bar{w}_\alpha \sigma_{\alpha\beta}^{(j)} w_\beta, \\ E &= H(z, \bar{z}, w, \bar{w}).\end{aligned}$$

- There are five constants of motion

$$\Gamma = \bar{z}_\alpha z_\alpha - \bar{w}_\alpha w_\alpha,$$

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$$E = H(z, \bar{z}, w, \bar{w}).$$

- It is important to note that we must require $\Gamma = 0$.

- We are interested in the equation of motion for the physical coordinates of a particle $X_i = \langle z_\alpha | \hat{X}_i | z_\alpha \rangle$. These can be extracted from the results above and read in the limit $R \gg 1$ for the dimensionless coordinates

$$\ddot{\vec{X}}_\pm = \frac{W'(R)}{R} \left[(\vec{X} \times \dot{\vec{X}}) \pm \sqrt{1 - \dot{\vec{X}} \cdot \dot{\vec{X}}} \vec{X} \right].$$

The dimensionless conserved quantities are

$$\vec{L}_\pm = \sqrt{1 - \dot{\vec{X}} \cdot \dot{\vec{X}}} (\vec{X} \times \dot{\vec{X}}) \pm ((\vec{X} \times \dot{\vec{X}}) \times \dot{\vec{X}}),$$

$$E_\pm = 1 \pm \sqrt{1 - \dot{\vec{X}} \cdot \dot{\vec{X}}} + W(R).$$

- The dimensionful version of the equation of motion is

$$\ddot{\vec{x}}_{\pm} = \frac{w'(r)}{mr} \left[\frac{m\lambda}{\hbar} (\vec{x} \times \dot{\vec{x}}) \pm \sqrt{1 - \left(\frac{m\lambda}{\hbar}\right)^2 \dot{\vec{x}} \cdot \dot{\vec{x}} \vec{x}} \right],$$

and the conserved quantities

$$\begin{aligned} \vec{\ell}_{\pm} &= \hbar \vec{L} = m \left[\sqrt{1 - \left(\frac{m\lambda}{\hbar}\right)^2 \dot{\vec{x}} \cdot \dot{\vec{x}}} (\vec{x} \times \dot{\vec{x}}) \right. \\ &\quad \left. \pm \frac{m\lambda}{\hbar} \left((\vec{x} \times \dot{\vec{x}}) \times \dot{\vec{x}} \right) \right], \end{aligned}$$

$$\mathbf{e}_{\pm} = \frac{\hbar^2}{m\lambda^2} \left[1 \pm \sqrt{1 - \left(\frac{m\lambda}{\hbar}\right)^2 \dot{\vec{x}} \cdot \dot{\vec{x}}} \right] + w(r).$$

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 - They predict a limiting speed of $v_0 = \frac{\hbar}{m\lambda}$,
 - They predict a cut-off in kinetic energy of $e_k \leq \frac{2\hbar^2}{m\lambda^2}$,

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- They predict two branches, depending on the energy,
- One branch reduces to standard Newton dynamics at speeds $v \ll v_0$.
- From

$$\vec{L}_{\pm} \cdot \dot{\vec{X}} = 0, \quad \vec{L}_{\pm} \cdot \ddot{\vec{X}}_{\pm} = 0, \quad \vec{L}_{\pm} \cdot \vec{X} = \mp \vec{L} \cdot \vec{L} \equiv \mp L^2,$$

we conclude that the motion is still planar, but displaced as $\vec{L} \cdot \vec{X} \neq 0$ as for central Newtonian dynamics.

Effective Radial Potential for Gravity

- From the conserved energy, we can construct an effective radial potential. In the case of gravity this reads in dimensionless units ($\beta = \frac{GMm^2\lambda}{\hbar^2} > 0$)

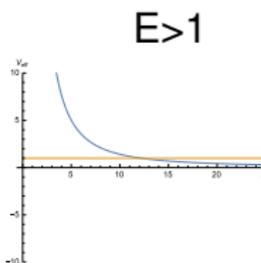
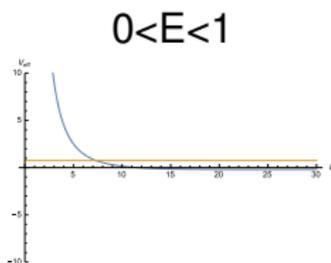
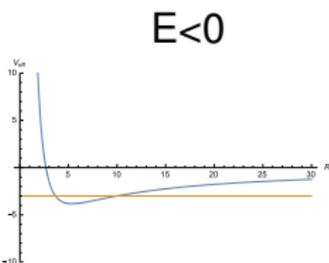
$$\dot{R}^2 + \frac{2(E-1)\beta}{R} + \frac{\beta^2 + L^2}{R^2} \equiv \dot{R}^2 + V_{\text{eff}} = E(2-E).$$

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- We immediately observe that the energy must be limited by $E < 2$ for this to have a solution.



Precession in a Gravitational Potential

- The equations of motion in general lead to precession of the orbitals. In the case of gravity one can compute this to leading order in the non-commutative parameter, with the result

$$\Delta\phi = \pi + \frac{\pi GM}{8a(1 - \epsilon^2)v_0^2}.$$

where a is the length of the semi-major axis and ϵ the eccentricity.

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- This is remarkably similar to the GR result

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Precession in a Gravitational Potential

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- The equations of motion in general lead to precession of the orbitals. In the case of gravity one can compute this to leading order in the non-commutative parameter, with the result

$$\Delta\phi = \pi + \frac{\pi GM}{8a(1 - \epsilon^2)v_0^2}.$$

where a is the length of the semi-major axis and ϵ the eccentricity.

- This is remarkably similar to the GR result

$$\Delta\phi = \pi + \frac{3\pi GM}{c^2 a(1 - \epsilon^2)},$$

- This comparison must be done with care as the noncommutative result is a noncommutative perturbation of flat space and not curved space as in the case of GR.

Stable Circular Orbitals in a Gravitational Potential 1/2

- Let us make the following ansatz for circular orbitals

$$x(t) = r \sin \theta \cos(\omega t), \quad y(t) = r \sin \theta \sin(\omega t), \quad z(t) = r \cos \theta.$$

Note that θ is time-independent.

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- We can now compute the radial dependence of the velocity (velocity curve):

$$v(r) = v_0 \sqrt{\frac{2}{1 + \sqrt{1 + 4 \left(\frac{r}{r_0}\right)^2}}},$$

$$\cot \theta = \sqrt{\frac{2}{\sqrt{1 + 4 \left(\frac{r}{r_0}\right)^2} - 1}}.$$

where v_0 is the limiting velocity and $r_0 = \frac{GM}{v_0^2}$.

Stable Circular Orbitals in a Gravitational Potential 2/2

- We note the following interesting behaviour

$$v(r) = v_0, \quad r \ll r_0,$$

$$v(r) = \sqrt{\frac{GM}{r}}, \quad r \gg r_0.$$

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- In this scenario Newton dynamics applied to orbits of stars close to the centre will lead to a severe underestimation of included mass

- The equations of motion can explicitly be written in covariant form

$$\frac{d^2 x^\mu}{d\tau^2} + \tilde{\Gamma}_{\lambda\nu}^\mu \frac{dx^\lambda}{d\tau} \frac{dx^\nu}{d\tau} = S_{\lambda\nu}^\mu \frac{dx^\lambda}{d\tau} \frac{dx^\nu}{d\tau}$$

where $\tilde{\Gamma}_{\lambda\nu}^\mu$ are the Levi-Civita connections and $d\tau$ the proper time of the metric (to leading order in λ)

$$g_{\mu\nu} = \begin{pmatrix} 1 - \frac{\lambda}{r} & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}.$$

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- Note: We use as fiducial frame for the connections the one in which the metric has the above form.

- The $S_{\lambda\nu}^{\mu}$ are the differences of two sets of connections and explicitly read in the fiducial frame (to leading order in λ)

$$S_{00}^i = \frac{\lambda}{2r^3} \left(1 + \left(\frac{mv_0^2 - e}{mv_0^2} \right)^2 \right) x^i,$$

$$S_{0j}^0 = -\frac{\lambda}{2r^3} x_j,$$

$$S_{0j}^i = \frac{\lambda (e - mv_0^2)}{2mv_0^2 r^3} \epsilon_{ijk} x^k,$$

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- Note that all the $S_{\lambda\nu}^{\mu}$ vanish in the commutative limit and that $g_{\mu\nu}$ reduces to the Minkowski metric.

Concluding Remarks

- The emergence of a mass dependent limiting speed is problematic. This can be bypassed by assuming that λ is mass dependent such that v_0 is mass independent.

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- The more exiting scenario is one where $v_0 < c$, perhaps even locally, in which case there are observational consequences, but it is not clear that this can be done consistently.

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- The more exiting scenario is one where $v_0 < c$, perhaps even locally, in which case there are observational consequences, but it is not clear that this can be done consistently.
- The emergence of the Lorentz geometry is a surprise.
- It is not that easy to introduce a short length scale and preserve the rotational symmetry and any such construction may share the features above. We therefore expect these to be rather generic.