Collider signals of models with radiatively-induced neutrino masses

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in collaboration with **A. Santamaria** and **J. Alcaide.** Based on hep-ph:1709.XXXXX.



Where is new physics? (the LHC paradox)



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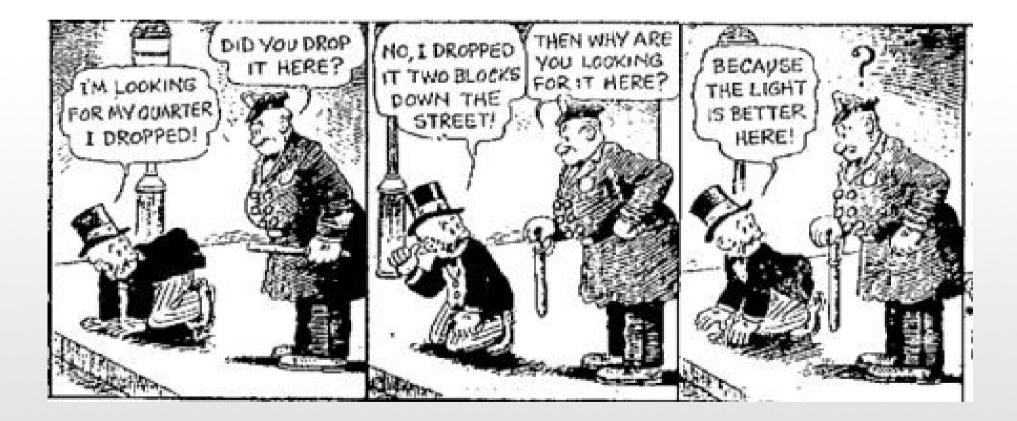
The hierarchy problem suggests that there should be new physics at the TeV scale

All of the (around 2.000) papers published by ATLAS and CMS have seen nothing

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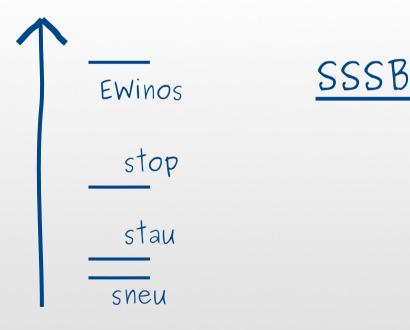


SUSY searches assume R-parity conservation, lightest neutralino is dark matter

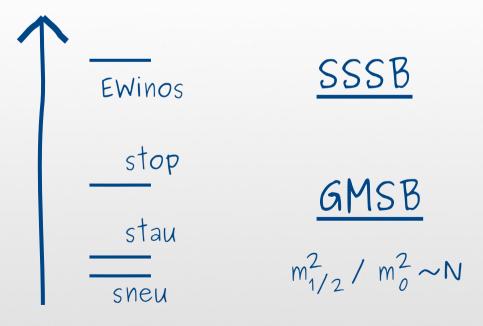
 $M_{\tilde{f}} > 1 \text{ TeV}$













EWinos	SSSB	
stop	GMSB	M7 > 350 GeV
stau sneu	$m_{1/2}^2 / m_0^2 \sim N$	[MC, Delgado, Nardini, Quirós, 1702.07359]

VLQ searches assume decays into SM particles only

T ______ t ____ MT>1 TeV

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MT > 600 GeV

[MC, 1705.03013]

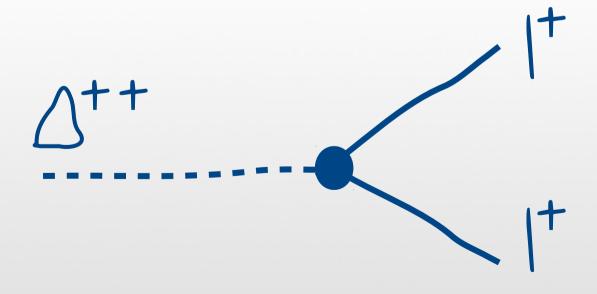


Strongly influenced by the MCHM. Already in the NMCHM:

4d UV completion, DM, BAU, ... and most models are non-minimal! MT > 600 GeV

[MC, 1705.03013]

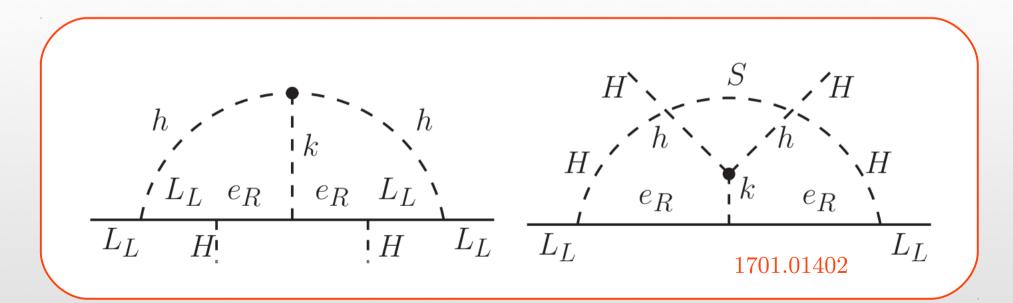






Scalar exotic decays are probably present

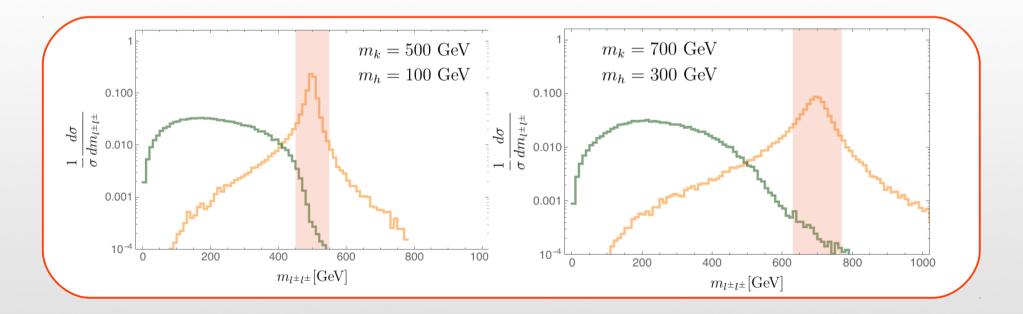
- Models with radiatively-induced neutrino masses do contain new (scalar) particles at the TeV scale
- LNV scalars must coupled linearly to non-leptonic fields



The status of current analyses

Many searches for doubly-chaged scalars (e.g. 1211.6312, 1210.5070, 1207.2666, ATLAS-CONF-2016-051, ...)

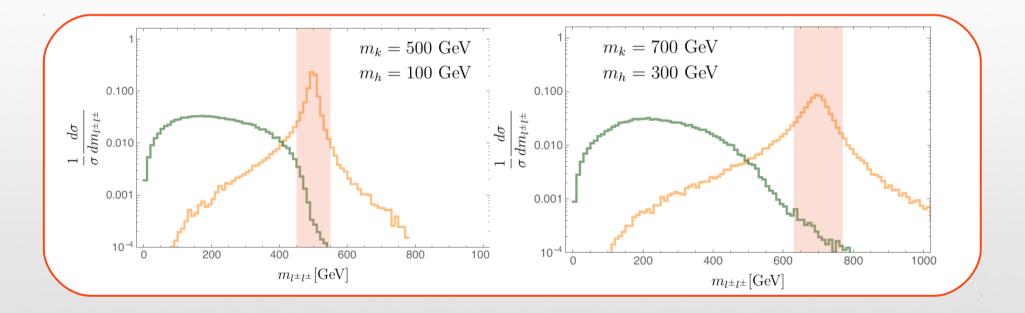
They all set a narrow cut on $m_{\ell^{\pm}\ell^{\pm}}$



The status of current analyses

SUSY searches (e.g. 1602.09058) are also non-constraining (checked with CheckMate 2)

Broad-scope analyses are also not sensitive



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The ATLAS Collaboration

Abstract

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We need to know where to look

- We propose a broad but still sensitive search for doublycharged scalars in neutrino models
- We consider three (orthogonal) signal regions: 21, 31, 41



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$$\begin{split} m_T^2 &= \left[\sqrt{(p_T^{\ell^{\pm}\ell^{\pm}})^2 + m_{\ell^{\pm}\ell^{\pm}}^2 + E_T^{\text{miss}}} \right]^2 \\ &- \left[p_x^{\ell^{\pm}\ell^{\pm}} + E_x^{\text{miss}} \right]^2 - \left[p_y^{\ell^{\pm}\ell^{\pm}} + E_y^{\text{miss}} \right]^2 \end{split}$$

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$$m_{T2} = \min_{\mathbf{q}_T} \left\{ \max \left[p_T^{L_1} E_T^{\text{miss}} - \mathbf{p}_T^{L_1} \cdot \mathbf{q}_T, \right. \\ \left. p_T^{L_2} E_T^{\text{miss}} - \mathbf{p}_T^{L_2} \cdot \left(\mathbf{E}_T^{\text{miss}} - \mathbf{q}_T \right) \right] \right\}.$$

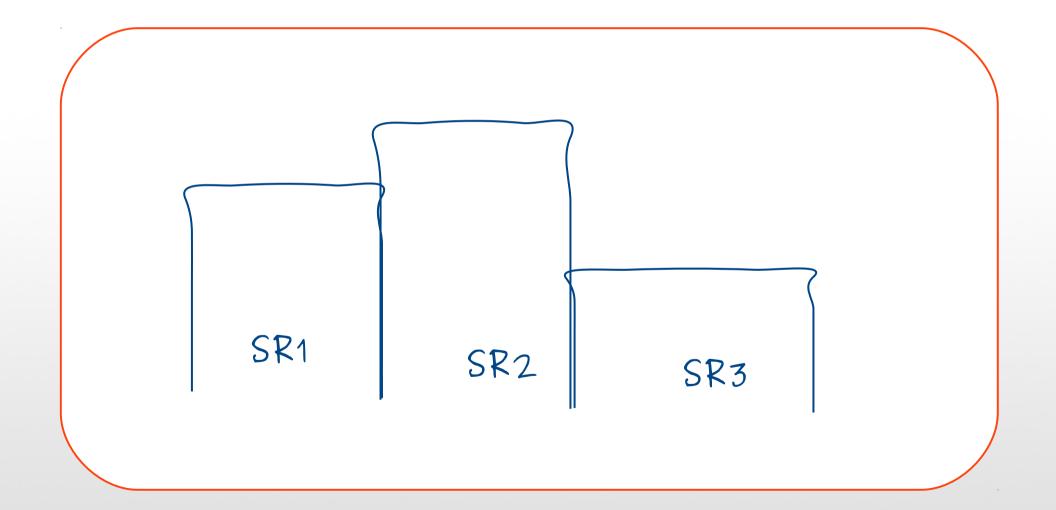
In addition, we consider the observables $m_{\ell^{\pm}\ell^{\pm}}, S_T$. For each SR, we look for the category with highest S/\sqrt{B}

$m_{T2} \backslash S_T > [\text{GeV}]$	100	200	300	400	500	600	700	800	900
100									
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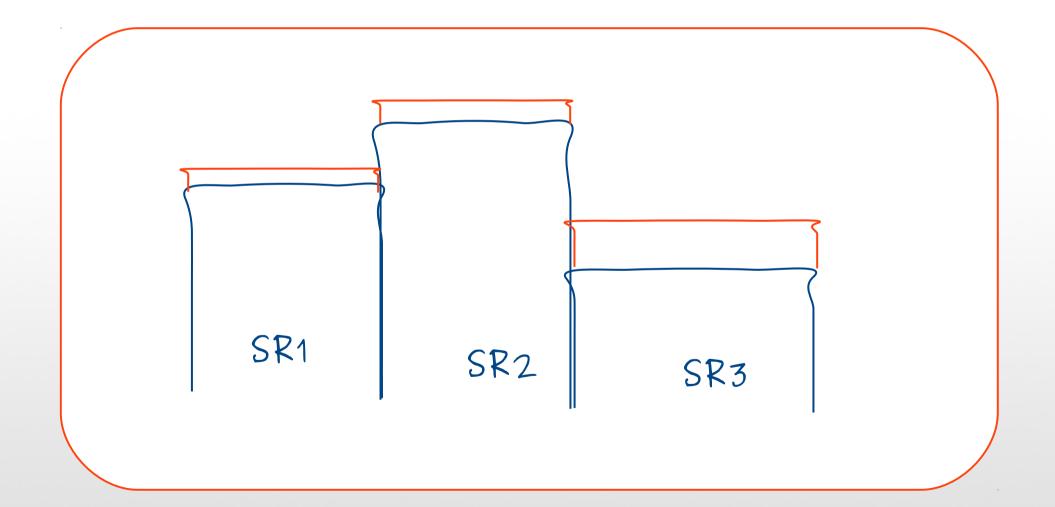
r	$m_{T2} \setminus S_T > [\text{GeV}]$	100	200	300	400	500	600	700	800	900
	100	19	5.5	1.7	0.65	0.2	0.11	0.063	0.056	0.047
	200	1.1	0.97	0.55	0.21	0.08	0.051	0.036	0.031	0.02
	300	0.19	0.19	0.17	0.058	0.047	0.03	0.02	0.02	0.017
	400	0.035	0.035	0.034	0.034	0.024	0.021	0.022	0.02	0.017
	500	0.021	0.021	0.021	0.021	0.021	0.018	0.017	0.017	0.017
	600	0.018	0.018	0.018	0.018	0.018	0.018	0.017	0.017	0.017
	700	0.017	0.017	0.017	0.017	0.017	0.017	0.017	0.017	0.017
	800	0.017	0.017	0.017	0.017	0.017	0.017	0.017	0.017	0.017
	900	0	0	0	0	0	0	0	0	0

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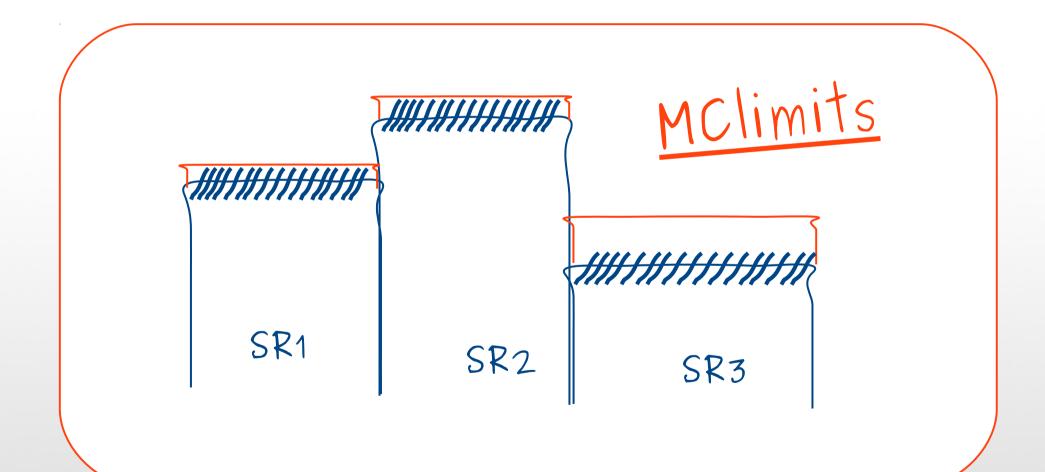


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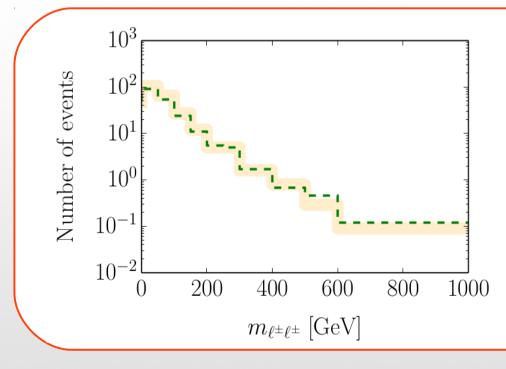
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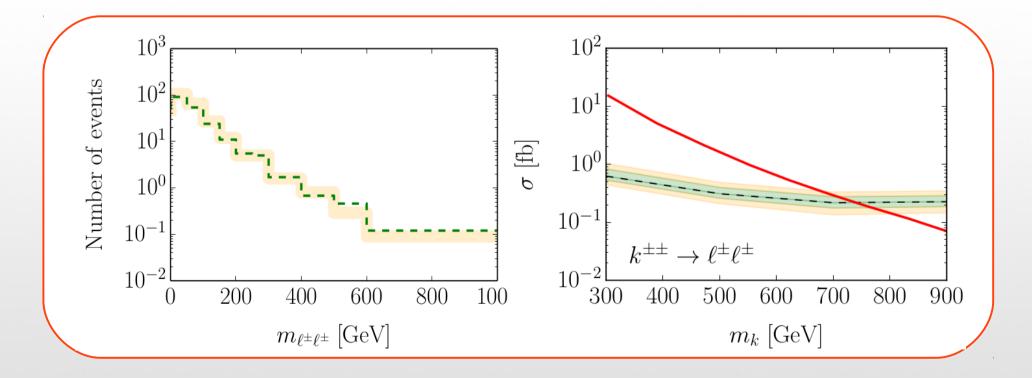
First, the background simulation is validated

We consider Z+jets, tt, WW, WZ, ZZ, WWW, WWZ, WZZ, ZZZ, ttW, ttZ [NLO accuracy]



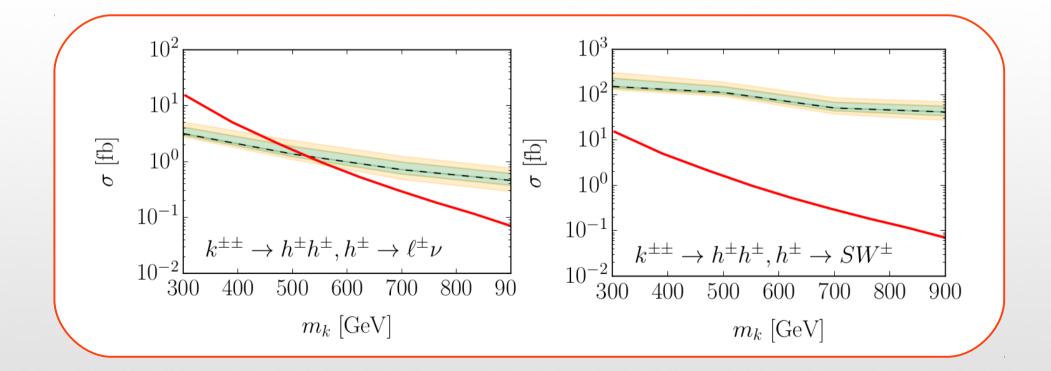
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Some constraints

We assume 100% into a given final state, and a total collected luminosity of 37/fb



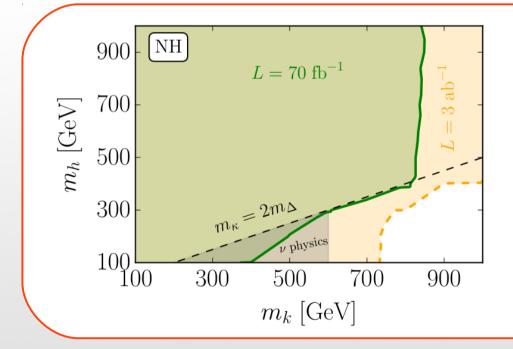
Implications for the Zee-Babu model

- Extend the SM with a singly-charged scalar and a doubly-charged one:
- Most of the parameter space already constrained by neutrino and low-energy data

$$L = L_{\rm SM} + f^{ab} \overline{\tilde{L}_{aL}} L_{Lb} h^+ + g^{ab} \overline{e_a^c} e_b k^{++} - \mu k^{++} h^- h + \text{h.c.} + \cdots$$

Implications for the Zee-Babu model

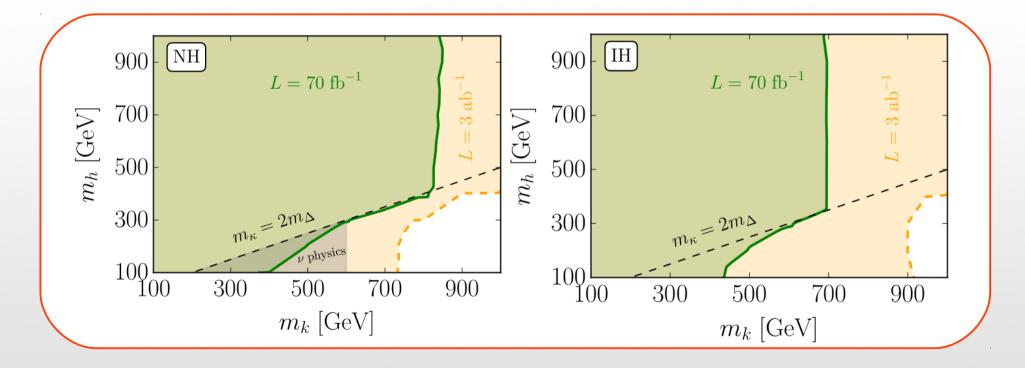
For example, in the normal hierarchy: $g_{11} = g_{22} = 0.1$, $g_{12} = g_{13} = g_{33} = 0.001$, $f_{12} = f_{13} = 0.01$, $f_{23} = 0.02$,





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Conclusions



- Realistic models of new physics are not necessarily ruled out (not at all!). New searches must be worked out,
- We proposed a broad analysis sensitive to scalars appearing in neutrino models. Background estimations are provided,
- New modes not previously constrained can be already tested with current data. Others (mainly those arising in DM models), are still far from reacheable
- For the Zee-Babu model, we can probe regions not previously tested by neutrino physics, nor by low-energy experiments

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Thank you very much for your attention!



