Gravity as a gauge theory in non-commutative spaces

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- Gravity as gauge theory in 3+1 and 2+1 dimensions
- ② Gauge theory on nc spaces
- Application on the nc dS3 and fuzzy sphere

Gravity in four dimensions as a gauge theory

Use of alternative approach of GR - vielbein formulation

4-d gravity described as a gauge theory of the Poincare group ISO(3,1) \downarrow consists of 10 generators:

see for details Kibble - Stelle '85, Utiyama '56, McDowell-Mansuri '79

- 4 of local translations P_a
- 6 Lorentz transformations M_{ab}

The generators satisfy the commutation relations:

$$[M_{ab}, M_{cd}] = \eta_{ac}M_{db} - \eta_{bc}M_{da} - \eta_{ad}M_{cb} + \eta_{bd}M_{ca}$$
$$[P_a, M_{bc}] = \eta_{ab}P_c - \eta_{ac}P_b$$
$$[P_a, P_b] = 0$$

where
$$\eta_{ab} = \text{diag}(-1, +1, +1, +1)$$
.



- Gauging: For each generator \leadsto introduction of a gauge field:
 - Vielbein $e_{\mu}^{\ a}$ corresponding to translations
 - Spin connection $\omega_{\mu}^{\ ab}$ corresponding to Lorentz transformations
- Therefore, the gauge connection is expanded as:

$$A_{\mu} = e_{\mu}^{\ a}(x)P_{a} + \frac{1}{2}\omega_{\mu}^{\ ab}(x)M_{ab}$$

• A_{μ} transforms in the adjoint rep:

$$\delta A_{\mu} = \partial_{\mu} \epsilon + [A_{\mu}, \epsilon] \,,$$

where ϵ is a parameter valued in **iso**(3, 1):

$$\epsilon = \xi^a(x)P_a + \frac{1}{2}\lambda^{ab}(x)M_{ab}$$

• The transformations of the gauge fields, e, ω are:

$$\begin{split} \delta e_{\mu}^{\ a} &= \partial_{\mu} \xi^a - e_{\mu}^{\ b} \lambda^a_{\ b} + \omega_{\mu}^{\ ab} \xi_b \\ \delta \omega_{\mu}^{\ ab} &= \partial_{\mu} \lambda^{ab} - \lambda^a_{\ c} \omega_{\mu}^{\ cb} + \lambda^b_{\ c} \omega_{\mu}^{\ ca} \end{split}$$

• Curvature tensors are obtained using the standard formula:

$$R_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu} + [A_{\mu}, A_{\nu}]$$

• Writing $R_{\mu\nu}=R_{\mu\nu}{}^aP_a+\frac{1}{2}R_{\mu\nu}{}^{ab}M_{ab},$ we obtain:

$$\begin{split} R_{\mu\nu}{}^{a} &= \partial_{\mu}e_{\nu}{}^{a} - \partial_{\nu}e_{\mu}{}^{a} + e_{\mu}{}^{b}\omega_{\nu b}{}^{a} - e_{\nu}{}^{b}\omega_{\mu b}{}^{a} \\ R_{\mu\nu}{}^{ab} &= \partial_{\mu}\omega_{\nu}{}^{ab} - \partial_{\nu}\omega_{\mu}{}^{ab} - \omega_{\mu}{}^{cb}\omega_{\nu}{}^{a}{}_{c} + \omega_{\mu}{}^{ac}\omega_{\nu c}{}^{b} \end{split}$$

Action is needed to complete the picture:

- Built out of Poincare invariants
- Analogy with Y-M theory suggests an action of the form:

$$S = \int d^4 \xi R_{ab}^{\ cd} R^{ab}_{\ cd}$$

• The right choice is:

$$\mathcal{S}_E = \frac{1}{16\pi G} \int d^4x \sqrt{-g} R_{ab}^{ab} ,$$

which can be written as:

$$S_E = \frac{1}{16\pi G} \int d^4x \sqrt{-g} e^{\mu}_{\ a} e^{\nu}_{\ b} (\partial_{\mu}\omega_{\nu}^{\ ab} - \partial_{\nu}\omega_{\mu}^{ab} + \omega_{\mu}^{\ ac}\omega_{\nu c}^{\ b} - \omega_{\nu}^{\ ac}\omega_{\mu c}^{\ b})$$

- → Functional of both the vielbeins and the spin connections
- → First order formulation of GR equations



Varying with respect to the fields → e.o.m.:

- with respect to $\omega \leadsto$ torsion-free condition
 - ✓ Torsion-free condition holds when scalars coupled to gravity
 - X Torsion non-zero when spinors coupled to gravity
- with respect to $e \rightsquigarrow \text{Einstein field equations (no matter)}$

Therefore, we conclude:

- Form of Einstein action: $A^2(dA + A^2)$
- Such action does not exist in gauge theories
- Gravity *cannot be* considered as a gauge theory.

Gravity in three dimensions as a gauge theory

The Einstein action for 3-d gravity is:

Witten '88

$$S = \int_{M} \epsilon^{\mu\nu\rho} \left(e_{\mu a} (\partial_{\nu} \omega_{\rho}^{\ a} - \partial_{\rho} \omega_{\nu}^{\ a} + \epsilon_{abc} \omega_{\nu}^{\ b} \omega_{\rho}^{\ c}) \right)$$

- Consideration of e and ω as gauge fields
- The above action, S, is $AdA + A^3$ in general form
- \bullet Interpretation of ${\mathcal S}$ as a Chern-Simons 3-form

Commutation relations of ISO(2,1):

$$[J_a, J_b] = \epsilon_{abc} J^c \quad [J_a, P_b] = \epsilon_{abc} P^c \quad [P_a, P_b] = 0$$

Construction of a gauge theory for ISO(2, 1):

- Gauge field Lie valued one form: $A_{\mu} = e_{\mu}^{\ a} P_a + \omega_{\mu}^{\ a} J_a$
- Infinitesimal gauge parameter: $u = \rho^a P_a + \tau^a J_a$
- Transformation of A_{μ} under a gauge trans/tion: $\delta A_{\mu} = -D_{\mu}u$
- Covariant derivative: $D_{\mu} = \partial_{\mu} u + [A_{\mu}, u]$



ullet Standard procedure \leadsto transformations of the fields:

$$\delta e_{\mu}^{\ a} = -\partial_{\mu} \rho^{a} - e_{\mu}^{\ b} \tau^{c} \epsilon_{abc} - \omega_{\mu}^{\ b} \rho^{c} \epsilon_{abc}$$
$$\delta \omega_{\mu}^{\ a} = -\partial_{\mu} \tau^{a} - \omega_{\mu}^{\ b} \tau^{c} \epsilon_{abc}$$

• Curvature tensor \leadsto commutator of covariant derivatives:

$$F_{\mu\nu} = [D_{\mu}, D_{\nu}] = P_a(\partial_{\mu}e_{\nu}^{\ a} - \partial_{\nu}e_{\mu}^{\ a} + \epsilon^{abc}(\omega_{\mu b}e_{\nu c} + e_{\mu b}\omega_{\nu c}))$$
$$+ J_a(\partial_{\mu}\omega_{\nu}^{\ a} - \partial_{\nu}\omega_{\mu}^{\ a} + \epsilon^{abc}\omega_{\mu b}\omega_{\nu c})$$

• IF we had considered ISO(2,1) gauge theory on a 4-d manifold, Y, we would form a topological invariant:

$$Tr(T^aT^b)\,\int F^aF^b$$

• Calculations lead to the expression:

$$\frac{1}{2} \int_{Y} \epsilon^{\mu\nu\rho\sigma} (\partial_{\mu} e_{\nu}^{\ a} - \partial_{\nu} e_{\mu}^{\ a} + \epsilon^{abc} (\omega_{\mu b} e_{\nu c} + e_{\mu b} \omega_{\nu c}) \times (\partial_{\rho} \omega_{\sigma a} - \partial_{\sigma} \omega_{\rho a} + \epsilon_{ade} \omega_{\rho}^{\ d} \omega_{\sigma}^{\ e})$$

- Integrand can be written as a total derivative
- Integral reduces to integral on the 3-d boundary of Y, M.
- This integral is by definition the Chern-Simons action:

$$S_{CS} = \int_{M} \epsilon^{\mu\nu\rho} \left(e_{\mu a} (\partial_{\nu} \omega_{\rho}^{\ a} - \partial_{\rho} \omega_{\nu}^{\ a} + \epsilon_{abc} \omega_{\nu}^{\ b} \omega_{\rho}^{\ c}) \right)$$

- Therefore, ISO(2,1) Chern Simons action is identical to 3-d Einstein action
- 3-d gravity is a Chern Simons gauge theory of ISO(2,1)

3-d Gravity with cosmological constant as a gauge theory

Generalization of the previous case, with action:

$$S = \int_{M} \epsilon^{\mu\nu\rho} \left(e_{\mu a} (\partial_{\nu} \omega_{\rho}^{\ a} - \partial_{\rho} \omega_{\nu}^{\ b}) + \epsilon_{abc} e_{\mu}^{\ a} \omega_{\nu}^{\ b} \omega_{\rho}^{\ c} + \frac{1}{3} \lambda \epsilon_{abc} e_{\mu}^{\ a} e_{\nu}^{\ b} e_{\rho}^{\ c} \right)$$

We note:

- Not Minkowski but dS or AdS depending on the sign of λ
- Their symmetry is *not* ISO(2,1) but SO(3,1) and SO(2,2)
- Since 3-d gravity \leftrightarrow gauging ISO(2,1) \Rightarrow 3-d gravity with $\lambda \to$ gauging SO(3,1) and SO(2,2), respectively (?)

First thing to do, generalization of the algebra:

$$[J_a,J_b]=\epsilon_{abc}J^c \quad [J_a,P_b]=\epsilon_{abc}P^c \quad [P_a,P_b]=\lambda\epsilon_{abc}J^c$$



• Repeating the procedure of the gauging → generalization of the transformations of the fields:

$$\delta e_{\mu}^{\ a} = -\partial_{\mu} \rho^{a} - e_{\mu}^{\ b} \tau^{c} \epsilon_{abc} - \omega_{\mu}^{\ b} \rho^{c} \epsilon_{abc}.$$

$$\delta \omega_{\mu}^{\ a} = -\partial_{\mu} \tau^{a} - \omega_{\mu}^{\ b} \tau^{c} \epsilon_{abc} - \lambda \epsilon^{abc} e_{\mu b} \rho_{c}$$

• The expression for the curvature is:

$$F_{\mu\nu} = P_a(\partial_\mu e_\nu^{\ a} - \partial_\nu e_\mu^{\ a} + \epsilon_{abc}(\omega_\mu^{\ b} e_\nu^{\ c} + \omega_\nu^{\ c} e_\mu^{\ b}))$$
$$+ J_a(\partial_\mu \omega_\nu^{\ a} - \partial_\nu \omega_\mu^{\ a} + \epsilon^{abc}(\omega_{\mu b}\omega_{\nu c} + \lambda e_{\mu b} e_{\nu c}))$$

- Chern Simons 3—form is precisely the Einstein action.
- E.o.m. \rightsquigarrow vanishing of the field strength tensor:
 - vanishing of the coefficient of $P_a \Rightarrow \omega$ is Levi Civita connection (torsionless condition)
 - vanishing of the coefficient of $J_a \Rightarrow$ Einstein equation with cosmological constant



Gauge theories on nc spaces

- Algebra $\mathcal A$ of operators $X^\mu \to \mathrm{nc}$ space with Madore Schupp Wess '00
- Operators X^{μ} satisfy the comm relation $[X_{\mu}, X_{\nu}] = i\theta_{\mu\nu}, \theta_{\mu\nu}$ not specified.
- Introduction of nc gauge theories through covariant nc coordinates, defined as: $\mathcal{X}_{\mu} = X_{\mu} + A_{\mu}$, obeying a covariant gauge transformation rule: $\delta \mathcal{X}_{\mu} = [\epsilon, \mathcal{X}_{\mu}] \ (\sim \text{cov der})$
- A_{μ} transforms as: $\delta A_{\mu} = -[X_{\mu}, \epsilon] + [\epsilon, A_{\mu}] \ (\sim \text{gauge connection})$
- A_{μ} is used to define an nc covariant field strength which defines the nc gauge theory:

$$F_{\mu\nu} = \left[\mathcal{X}_{\mu}, \mathcal{X}_{\nu}\right] - i\bar{\theta}_{\mu\nu} \,, \quad F_{\mu\nu} = \left[\mathcal{X}_{\mu}, \mathcal{X}_{\nu}\right] - C_{\mu\nu\rho}\mathcal{X}_{\rho} \,,$$

cases of constant and linear noncommutativity.



- Gauge theory could be abelian or nonabelian:
 - Abelian if ϵ is a function in \mathcal{A} .
 - Nonabelian if ϵ is matrix valued.
- ▷ In nonabelian case, where are the gauge fields valued?
- Let us consider the relation:

$$[\epsilon,A] = [\epsilon^A T^A, A^B T^B] = \frac{1}{2} \{\epsilon^A, A^B\} [T^A, T^B] + \frac{1}{2} [\epsilon^A, A^B] \{T^A, T^B\} \,,$$

• Cannot restrict to a matrix algebra - last term neither 0 nor algebra element in nc.

see Prof. Castellani's lectures

- There are two options to overpass the difficulty:
 - Consider the universal enveloping algebra
 - Extending the generators and/or fixing the rep so that the anticommutators close.
- > We employ the second option.



3d gravity with cosmological constant in nc

- The cov coord should accommodate info of no vielbein and spin connection (analogy with the gauging of Poincare/(A)dS group)

 Nair '03, Abe Nair '03, Nair '06
- We consider the 3-d case with positive λ .
- The relevant isometry group is SO(3,1) (SL(2, **C**) the corresponding spin group)
- Nonabelian group \rightarrow focus on the spinor rep with generators: $\Sigma_{AB} = \frac{1}{2} \gamma_{AB} = \frac{1}{4} [\gamma_A, \gamma_B], A, B = 1 \dots 4.$

Due to the product relation:

$$\gamma_{AB}\gamma^{CD} = 2\delta^{[C}_{[B}\delta^{D]}_{A]} + 4\delta^{[C}_{[B}\gamma_{A]}{}^{D]} + i\varepsilon_{AB}{}^{CD}\gamma_5 \ , \label{eq:gamma_AB}$$

one finds the commutation and anticommutation relations:

$$[\gamma_{AB}, \gamma_{CD}] = 8\eta_{[A[C}\gamma_{D]B]}$$

$$\{\gamma_{AB}, \gamma_{CD}\} = 4\eta_{C[B}\eta_{A]D}\mathbf{1} + 2i\epsilon_{ABCD}\gamma_{5}$$



- \bullet γ_5 and **1** have to be included in the algebra
- Extension by these two elements \rightarrow 8-dimensional algebra \rightarrow SL(2, **C**) to GL(2, **C**) with generators $\{\gamma_{AB}, \gamma_5, i\mathbf{1}\}$

P. Aschieri -L. Castellani '09

- In SO(3) notation we have the generators γ_{ab} and $\gamma_a = \gamma_{a4}$ with $a = 1, \ldots, 3$. We can also define: $\tilde{\gamma}^a = \epsilon^{abc} \gamma_{bc}$.
- Commutation and anticommutation relations for γ and $\tilde{\gamma}$:

$$\begin{split} & [\tilde{\gamma}^a,\tilde{\gamma}^b] = -4\epsilon^{abc}\tilde{\gamma}_c \,, \ [\gamma_a,\tilde{\gamma}_b] = -4\epsilon_{abc}\gamma^c \,, \ [\gamma_a,\gamma_b] = \epsilon_{abc}\tilde{\gamma}^c \,, \ [\gamma^5,\gamma^{AB}] = 0 \\ & \{\tilde{\gamma}^a,\tilde{\gamma}^b\} = -8\eta^{ab}\mathbf{1} \,, \ \{\gamma_a,\tilde{\gamma}^b\} = 4i\delta^b_a\gamma_5 \,, \ \{\gamma_a,\gamma_b\} = 2\eta_{ab}\mathbf{1} \,, \ \{\gamma^5,\gamma^{AB}\} = 0 \end{split}$$

 \bullet We consider $GL(2, \mathbf{C})$ as the gauge group. The covariant coordinate is:

$$\mathcal{X}_{\mu} = e_{\mu}{}^{a}(X) \otimes \gamma_{a} + \omega_{\mu}{}^{a}(X) \otimes \widetilde{\gamma}_{a} + \mathcal{A}_{\mu} \otimes i\mathbf{1} + \widetilde{A}_{\mu}(X) \otimes \gamma_{5} ,$$
with $\mathcal{A}_{\mu} = e_{\mu}{}^{a}X_{a} + A_{\mu}(X)$.



• Gauge parameter expands in a similar way:

$$\epsilon = \xi^a(X) \otimes \gamma_a + \lambda^a(X) \otimes \widetilde{\gamma}_a + \epsilon_0(X) \otimes i\mathbf{1} + \widetilde{\epsilon}_0(X) \otimes \gamma_5$$

• Trans of component fields derive from $\delta \mathcal{X}_{\mu} = [\epsilon, \mathcal{X}_{\mu}]$ and are:

$$\begin{split} \delta e_{\mu}^{\ a} &= -i[\mathcal{A}_{\mu}, \xi^{a}] - 2\{\xi_{b}, \omega_{\mu c}\}\epsilon^{abc} - 2\{\lambda_{b}, e_{\mu c}\}\epsilon^{abc} + i[\epsilon_{0}, e_{\mu}^{\ a}] \\ \delta \omega_{\mu}^{\ a} &= -i[\mathcal{A}_{\mu}, \lambda^{a}] + \frac{1}{2}\{\xi_{b}, e_{\mu c}\}\epsilon^{abc} - 2\{\lambda_{b}, \omega_{\mu c}\}\epsilon^{abc} + i[\epsilon_{0}, \omega_{\mu}^{\ a}] \\ \delta \mathcal{A}_{\mu} &= -i[\mathcal{A}_{\mu}, \epsilon_{0}] - i[\xi_{a}, e_{\mu}^{\ a}] + 4i[\lambda_{a}, \omega_{\mu}^{\ a}] - i[\tilde{\epsilon}_{0}, \widetilde{A}_{\mu}] \\ \delta \widetilde{A}_{\mu} &= -i[\mathcal{A}_{\mu}, \tilde{\epsilon}_{0}] + 2i[\xi_{a}, \omega_{\mu}^{\ a}] + 2i[\lambda_{a}, e_{\mu}^{\ a}] + i[\epsilon_{0}, \widetilde{A}_{\mu}] \end{split}$$

• If we consider: $e_{\mu}^{\ a}=\delta_{\mu}^{\ a},\,\omega_{\mu}^{\ a}=0,\,\tilde{A}_{\mu}=0$ (Y-M limit), we obtain:

$$\delta A_{\mu} = -i[X_{\mu}, \epsilon_0] + i[\epsilon_0, A_{\mu}]$$

recovering the trans rule for a nc Y-M gauge field.

• If we consider: $A_{\mu} = 0$, $[A_{\mu}, f] \to \partial_{\mu} f$ (comm limit), we obtain field trans of 3-d comm case.

After a redefinition:

$$\gamma_a \to \frac{2i}{\sqrt{\lambda}} P_a, \tilde{\gamma}_a \to -4J_a, 4\lambda^a \to \lambda^a, \xi^a \frac{2i}{\sqrt{\lambda}} \to -\xi^a, e^a_\mu \to \frac{\sqrt{\lambda}}{2i} e^a_\mu, \omega^a_\mu \to -\frac{1}{4} \omega^a_\mu$$

Curvature tensors are:

$$\mathcal{R}_{\mu\nu} = [\mathcal{X}_{\mu}, \mathcal{X}_{\nu}] - \epsilon_{\mu\nu\rho} \mathcal{X}_{\rho}$$

The curvature tensor can be expanded as:

$$\mathcal{R}_{\mu\nu} = T^a_{\mu\nu} \otimes \gamma_a + R^a_{\mu\nu} \otimes \tilde{\gamma}_a + F_{\mu\nu} \otimes i\mathbf{1} + \tilde{F}_{\mu\nu} \otimes \gamma_5$$

We end up with the various tensors:

$$\begin{split} T^{a}_{\mu\nu} &= i[\mathcal{A}_{\mu}, e_{\nu}^{\ a}] - i[\mathcal{A}_{\nu}, e_{\mu}^{\ a}] - 2\{e_{\mu b}, \omega_{\nu c}\}\epsilon^{abc} - 2\{\omega_{\mu b}, e_{\nu c}\}\epsilon^{abc} - \epsilon_{\mu\nu}{}^{\rho}e_{\rho}^{\ a} \\ R^{a}_{\mu\nu} &= i[\mathcal{A}_{\mu}, \omega_{\nu}^{\ a}] - i[\mathcal{A}_{\nu}, \omega_{\mu}^{\ a}] - 2\{\omega_{\mu b}, \omega_{\nu c}\}\epsilon^{abc} + \frac{1}{2}\{e_{\mu b}, e_{\nu c}\}\epsilon^{abc} - \epsilon_{\mu\nu}{}^{\rho}\omega_{\rho}^{\ a} \\ F_{\mu\nu} &= i[\mathcal{A}_{\mu}, \mathcal{A}_{\nu}] - i[e_{\mu}^{\ a}, e_{\nu a}] + 4i[\omega_{\mu}^{\ a}, \omega_{\nu a}] - i[\widetilde{A}_{\mu}, \widetilde{A}_{\nu}] - \epsilon_{\mu\nu}{}^{\rho}\mathcal{A}_{\rho} \\ \widetilde{F}_{\mu\nu} &= i[\mathcal{A}_{\mu}, \widetilde{A}_{\nu}] - i[\mathcal{A}_{\nu}, \widetilde{A}_{\mu}] + 2i[e_{\mu}^{\ a}, \omega_{\nu a}] + 2i[\omega_{\mu}^{\ a}, e_{\nu a}] - \epsilon_{\mu\nu}{}^{\rho}\widetilde{A}_{\rho} \end{split}$$

- If we consider the comm limit: same tensors as comm case
- If we consider the Y-M limit, we obtain:

$$F_{\mu\nu} = i[\mathcal{A}_{\mu}, \mathcal{A}_{\nu}] - \epsilon_{\mu\nu}{}^{\rho}\mathcal{A}_{\rho} \,, \quad \mathcal{A}_{\mu} \to X_{\mu} + A_{\mu} \,,$$

field strength tensor with \mathcal{A} interpreted as a cov coord.



Final step \rightarrow write down the action:

$$\mathcal{S} = \frac{1}{2} \mathrm{Trtr}(\epsilon^{\mu\nu\rho} \mathcal{X}_{\mu} \mathcal{X}_{\nu} \mathcal{X}_{\rho} - \mathcal{X}_{\mu} \mathcal{X}^{\mu}) = \frac{1}{4} \mathrm{Trtr}(\epsilon^{\mu\nu\rho} \mathcal{X}_{\mu} \mathcal{R}_{\nu\rho}) \,,$$

where

- Tr is trace over matrices
- tr is trace over the algebra

Using the following form of the algebra trace:

$$\operatorname{tr}(\gamma_a \gamma_b) = 4\eta_{ab}, \quad \operatorname{tr}(\tilde{\gamma}_a \tilde{\gamma}_b) = -16\eta_{ab},$$

it takes the form:

$$S = \text{Tr}\epsilon^{\mu\nu\rho} (e_{\mu a} T^a_{\nu\rho} - 4\omega_{\mu a} R^a_{\nu\rho} - \mathcal{A}_{\mu} F_{\nu\rho} + \tilde{A}_{\mu} \tilde{F}_{\nu\rho}).$$



Varying the action wrt \mathcal{X} , we obtain:

$$\mathcal{R}_{\nu\rho} + \frac{1}{3} \epsilon_{\nu\rho}^{\ \mu} \mathcal{X}_{\mu} = 0 \,,$$

which decompose to the set of e.o.m.:

$$\begin{split} T^a_{\nu\rho} - \frac{1}{3} \epsilon_{\nu\rho}{}^{\mu} e_{\mu}{}^a &= 0 \; , \quad R^a_{\nu\rho} - \frac{1}{3} \epsilon_{\nu\rho}{}^{\mu} \omega_{\mu}{}^a = 0 \; , \\ F_{\nu\rho} - \frac{1}{3} \epsilon_{\nu\rho}{}^{\mu} \mathcal{A}_{\mu} &= 0 \; , \quad \widetilde{F}_{\nu\rho} - \frac{1}{3} \epsilon_{\nu\rho}{}^{\mu} \widetilde{A}_{\mu} = 0 \; . \end{split}$$

Comments

- Derive e.o.m. after variation wrt the gauge fields expecting the same expressions
- Working on a model for gravity on a fuzzy sphere
- Next step include matter fields
- Purpose is to learn the tools and move on to more realistic scenarios

Thank you!