

# Phenomenological studies in the matrix models

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# §1 Introduction

- **Standard Model of Particle physics**

- Successful. Agrees well with experiments.
- Unsatisfactory as a final theory.
  - \* No quantum gravity
  - \* Too arbitrary

- ↑ **String phenomenology**

- Calabi-Yau, Orbifolds, Intersecting D-branes, etc.

- **String theory**

- Too many vacua. Landscape.
- We can not compare them dynamically.
- **Need an underlying, master theory**

# Candidates for nonperturbative formulations of String theory

- String field theories
- Gauge/gravity (AdS/CFT) dualities

- **Matrix Models**

T. Banks, W. Fischler, S.H. Shenker, L. Susskind, 1997

N. Ishibashi, H. Kawai, Y. Kitazawa, A. Tsuchiya, 1997

R. Dijkgraaf, E. P. Verlinde and H. L. Verlinde, 1997

# IIB (IKKT) matrix model

H. Kawai's talk

- Action

$$S_{\text{IIBMM}} = -\frac{1}{g_{\text{IIBMM}}^2} \text{Tr} \left( \frac{1}{4} [A_\mu, A_\nu] [A^\mu, A^\nu] + \frac{1}{2} \bar{\psi} \Gamma^\mu [A_\mu, \psi] \right) \quad (1)$$

- 

$$\int dA d\psi e^{-S_{\text{IIBMM}}}$$

$$N \rightarrow \infty$$

**Nonperturbative formulation of String theory**

# Spacetime structures from IIBMM

J. Nishimura's talk

HA, S. Iso, H. Kawai, Y. Kitazawa, T. Tada, 1998;

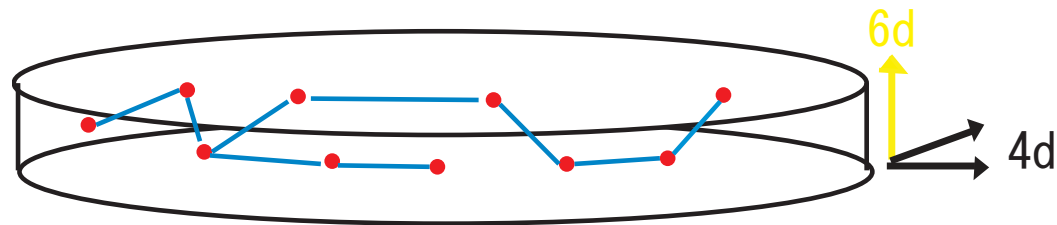
J. Nishimura, J. Ambjorn, K. Anagnostopoulos, T. Azuma, W. Bietenholtz, F. Hofheinz, T. Hotta, T. Okubo, F. Sugino, A. Tsuchiya, G. Vernizzi;

H. Kawai, S. Kawamoto, T. Kuroki, T. Matsuo, S. Shinohara

## Lorentzian IIBMM

S. Kim, J. Nishimura, A. Tsuchiya, 2012

- Eigenvalues distribution of  $A_\mu \longleftrightarrow$  Spacetime structure



$\Rightarrow$  4-dimensional spacetime?

Then, next, let's consider matter on the spacetime

SM  $\longleftarrow$  MM

## Importance of these studies:

1. Such a path may give us a guide for coming close from either side.
  - Study phenomenology beyond the SM
  - Justify/Modify the formulation of MM
2. We can, in principle, analyze dynamics and calculate everything, since MM has **definite action and measure**.  
An advantage that MM has over the ordinary string theories

# Mechanism to obtain chiral fermions

- An important ingredient of SM is the chiral fermions.

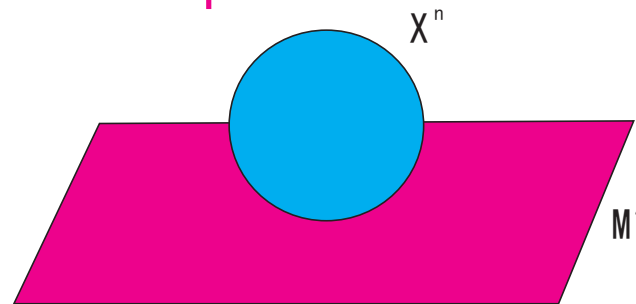
- Nontrivial topology in the extra dimensions

↓ Index Theorem

Chiral zero modes in the extra dimensions

↓  $D_{10} = D_4 + D_6$

Chiral massless fields on our spacetime



- This is the standard way.
  - Euler characteristics in CY
  - Boundary conditions in orbifolds
  - Intersection numbers in intersecting D-branes

# Chiral fermions and SM from MM

- Orbifolds

HA, S. Iso, T. Suyama, 2002

A. Chatzistavrakidis, H. Steinacker, G. Zoupanos, 2010

- Intersecting D-branes

A. Chatzistavrakidis, H. Steinacker, G. Zoupanos, 2011

J. Nishimura, A. Tsuchiya, 2013

HA, J. Nishimura, A. Tsuchiya, 2013

A. Tsuchiya's talk

↕ via T-duality

- Toroidal compactifications with magnetic fluxes  $\sim$  magnetized D-branes

HA, 2011, 2013, 2013



# Outline

✓ §1 Introduction

§2 Formulation of NC tori

§3 Matrix configurations for phenomenological models

§4 Probability distribution over the phenomenological models  
(Semiclassical analyses of MM)

§5 Conclusions and Discussions

## §2 Formulation of NC tori

# Spacetime and matter in matrices

- **Hermitian-matrix formulation**

HA, N. Ishibashi, S. Iso, H. Kawai, Y. Kitazawa, T. Tada, 2000

NC background

$$[p_\mu, p_\nu] = i\theta_{\mu\nu}$$

Decompose

$$A_\mu = p_\mu + a_\mu$$

- MM gives NC field theory.

$$\begin{aligned} & \text{tr} \left( [A_\mu, A_\nu]^2 + \bar{\psi} \Gamma_\mu [A_\mu, \psi] \right) \\ = & \int d^d x \left( (F_{\alpha\beta})^2 + (D_\alpha \phi_i)^2 + [\phi_i, \phi_j]^2 \right. \\ & \left. + \bar{\psi} \Gamma_\alpha D_\alpha \psi + \bar{\psi} \Gamma_i [\phi_i, \psi] \right)_* \end{aligned}$$

- **Unitary-matrix formulation**

J. Ambjorn, Y. M. Makeenko, J. Nishimura, R. J. Szabo, 2000

NC background

$$\Gamma_\mu \Gamma_\nu = \mathcal{Z}_{\mu\nu} \Gamma_\nu \Gamma_\mu$$

Decomposition

$$V_\mu = U_\mu \Gamma_\mu$$

- Correspondence between unitary and Hermitian matrices:

$$V_\mu \sim e^{iA_\mu/R}, \quad \Gamma_\mu \sim e^{ip_\mu/R}, \quad U_\mu \sim e^{ia_\mu/R}$$

- Twisted Eguchi-Kawai model gives NC plaquette action

$$\begin{aligned} & -\mathcal{N}\beta \sum_{\mu \neq \nu} \mathcal{Z}_{\nu\mu} \text{tr} \left( V_\mu V_\nu V_\mu^\dagger V_\nu^\dagger \right) \\ = & -\beta \sum_{\mu \neq \nu} \sum_x U_\mu(x) \star U_\nu(x + \epsilon \hat{\mu}) \star U_\mu(x + \epsilon \hat{\nu})^* \star U_\nu(x)^* \end{aligned}$$

- We will use finite-unitary-matrix formulation for NC tori.
- $M^4 \times T^6$  compactification

$$M^4 : A_\mu = x_\mu \otimes \mathbb{1}$$

$$T^6 : e^{iA_i/R} \sim \mathbb{1} \otimes V_i$$

with  $\mu = 0, \dots, 3$  and  $i = 4, \dots, 9$ .

- Alternatively, one can consider  $T^4 \times T^6$  with a huge anisotropy of sizes.
- Topological configurations  $V_i$  are given by using the Morita equivalence.

J. Ambjorn, Y. M. Makeenko, J. Nishimura, R. J. Szabo, 2000

R. J. Szabo, 2003

HA, J. Nishimura, Y. Susaki, 2009



- $\mathbb{I}_{p^1}, \dots, \mathbb{I}_{p^h}$  yield the gauge group  $U(p^1) \times U(p^2) \times \dots \times U(p^h)$ .
- The unitary matrix  $\Gamma_{l,i}^a$  represents NC  $T^2$  with magnetic flux  $q_l^a$ .
- $a = 1, \dots, h$  label the blocks.  
 $l = 1, 2, 3$  label  $l$ th  $T^2$  in  $T^6 = T^2 \times T^2 \times T^2$
- A topological configuration  $V_i$  is specified by the integers  $p^a$  and  $q_l^a$ .
- In fact,  $V_i$  are classical solutions for the unitary MM.

## The off-diagonal block $\psi^{ab}$ of adjoint fermion

$$\psi = \begin{pmatrix} \psi^{11} & \psi^{12} & \dots & \psi^{1h} \\ \psi^{21} & \psi^{22} & \dots & \psi^{2h} \\ \vdots & \vdots & \ddots & \vdots \\ \psi^{h1} & \psi^{h2} & \dots & \psi^{hh} \end{pmatrix}$$

- Bifundamental rep.  $(p^a, \bar{p}^b)$  under the gauge group  $U(p^a) \times U(p^b)$
- Topological charge, i.e., magnetic flux, in  $T^6$ :

$$\frac{1}{2} \text{Tr} [P^{aL} P^{bR} (\gamma + \hat{\gamma})] = p^a p^b \prod_{l=1,2,3} (q_l^a - q_l^b)$$

The Dirac index takes the correct values in a GW Dirac operator.

$\Rightarrow$  Generation number on our spacetime



## §3 Phenomenological models

- **Fermion content** = SM fermions with 3 generations
- SM gauge group  $\subset$  **Gauge group**  $\subset$   $U(8)$ 
  - (i)  $U(4) \times U_L(2) \times U_R(2)$  : Pati-Salam-like model
  - (ii)  $U_c(3) \times U(1) \times U_L(2) \times U_R(2)$
  - (iii)  $U(4) \times U_L(2) \times U(1)^2$
  - (iv)  $U_c(3) \times U(1) \times U_L(2) \times U(1)^2$

## (i) $U(4) \times U_L(2) \times U_R(2)$ model

Fermion embedding into matrix:

$$\psi = \begin{pmatrix} o & q & u & d \\ & l & \nu & e \\ \hline & o & o & \\ \hline & & & o \end{pmatrix}$$

$q$ : quark doublets

$l$ : lepton doublets

$u, d$ : quark singlets

$\nu, e$ : lepton singlets

$o$ : Vanishing index  $\Rightarrow$  No massless modes

## Flux $q_l^a$

$$q_l^{ab} = q_l^a - q_l^b \quad \dots T^2 \text{ flux}$$

$$q^{ab} = \prod_{l=1}^3 q_l^{ab} \quad \dots T^6 \text{ flux}$$

$T^6$  flux must take values

$$q^{ab} = \begin{pmatrix} 0 & -3 & 3 \\ & 0 & 0 \\ & & 0 \end{pmatrix}$$

4 solutions for  $T^2$  flux:

$\hat{q}_1^{ab}$	$\hat{q}_2^{ab}$	$\hat{q}_3^{ab}$
$\begin{pmatrix} -1 & 1 \\ & 2 \end{pmatrix}$	$\begin{pmatrix} 1 & 1 \\ & 0 \end{pmatrix}$	$\begin{pmatrix} 3 & 3 \\ & 0 \end{pmatrix}$
$\begin{pmatrix} -1 & -1 \\ & 0 \end{pmatrix}$	$\begin{pmatrix} 1 & 1 \\ & 0 \end{pmatrix}$	$\begin{pmatrix} 3 & -3 \\ & -6 \end{pmatrix}$
$\begin{pmatrix} -1 & -1 \\ & 0 \end{pmatrix}$	$\begin{pmatrix} 1 & -3 \\ & -4 \end{pmatrix}$	$\begin{pmatrix} 3 & 1 \\ & -2 \end{pmatrix}$
$\begin{pmatrix} -1 & -1 \\ & 0 \end{pmatrix}$	$\begin{pmatrix} 1 & 3 \\ & 2 \end{pmatrix}$	$\begin{pmatrix} 3 & -1 \\ & -4 \end{pmatrix}$

## (ii) $U_c(3) \times U(1) \times U_L(2) \times U_R(2)$ model

Fermion embedding into matrix:

$$\psi = \begin{pmatrix} o & o & q & u & d \\ o & l & \nu & e \\ o & o \\ o \end{pmatrix}$$

$T^6$  flux must take values

$$q^{ab} = \begin{pmatrix} 0 & 0 & -3 & 3 \\ 0 & -3 & 3 \\ 0 & 0 \\ 0 \end{pmatrix}$$

## 4 solutions for $T^2$ flux:

$\hat{q}_1^{ab}$	$\hat{q}_2^{ab}$	$\hat{q}_3^{ab}$
$\begin{pmatrix} -2 & -1 & -1 \\ & 1 & 1 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} -2 & -1 & -1 \\ & 1 & 1 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} 0 & -3 & 3 \\ & -3 & 3 \\ & & 6 \end{pmatrix}$
$\begin{pmatrix} 0 & -1 & 1 \\ & -1 & 1 \\ & & 2 \end{pmatrix}$	$\begin{pmatrix} 2 & 1 & 1 \\ & -1 & -1 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} 6 & 3 & 3 \\ & -3 & -3 \\ & & 0 \end{pmatrix}$
$\begin{pmatrix} 0 & -1 & 1 \\ & -1 & 1 \\ & & 2 \end{pmatrix}$	$\begin{pmatrix} 4 & 3 & 3 \\ & -1 & -1 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} 4 & 1 & 1 \\ & -3 & -3 \\ & & 0 \end{pmatrix}$
$\begin{pmatrix} 0 & -1 & 1 \\ & -1 & 1 \\ & & 2 \end{pmatrix}$	$\begin{pmatrix} 2 & 3 & 3 \\ & 1 & 1 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} -2 & 1 & 1 \\ & 3 & 3 \\ & & 0 \end{pmatrix}$

### (iii) $U(4) \times U_L(2) \times U(1)^2$ model

Fermion embedding into matrix:

$$\psi = \begin{pmatrix} o & q & u & d \\ & l & \nu & e \\ \hline & o & o & o \\ \hline & & o & o \\ \hline & & & o \end{pmatrix}$$

$T^6$  flux must take values

$$q^{ab} = \begin{pmatrix} 0 & -3 & 3 & 3 \\ & 0 & 0 & 0 \\ & & 0 & 0 \\ & & & 0 \end{pmatrix}$$



## 9 solutions for $T^2$ flux:

$\hat{q}_1^{ab}$	$\hat{q}_2^{ab}$	$\hat{q}_3^{ab}$
$\begin{pmatrix} -1 & 1 & -1 \\ & 2 & 0 \\ & & -2 \end{pmatrix}$	$\begin{pmatrix} 1 & 1 & -1 \\ & 0 & -2 \\ & & -2 \end{pmatrix}$	$\begin{pmatrix} 3 & 3 & 3 \\ & 0 & 0 \\ & & 0 \end{pmatrix}$
$\begin{pmatrix} -1 & 1 & -1 \\ & 2 & 0 \\ & & -2 \end{pmatrix}$	$\begin{pmatrix} 1 & 1 & 1 \\ & 0 & 0 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} 3 & 3 & -3 \\ & 0 & -6 \\ & & -6 \end{pmatrix}$
$\begin{pmatrix} -1 & 1 & 3 \\ & 2 & 4 \\ & & 2 \end{pmatrix}$	$\begin{pmatrix} 1 & 1 & 1 \\ & 0 & 0 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} 3 & 3 & 1 \\ & 0 & -2 \\ & & -2 \end{pmatrix}$
$\begin{pmatrix} -1 & 1 & -3 \\ & 2 & -2 \\ & & -4 \end{pmatrix}$	$\begin{pmatrix} 1 & 1 & 1 \\ & 0 & 0 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} 3 & 3 & -1 \\ & 0 & -4 \\ & & -4 \end{pmatrix}$
$\begin{pmatrix} -1 & -1 & 3 \\ & 0 & 4 \\ & & 4 \end{pmatrix}$	$\begin{pmatrix} 1 & 1 & 1 \\ & 0 & 0 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} 3 & -3 & 1 \\ & -6 & -2 \\ & & 4 \end{pmatrix}$
$\begin{pmatrix} -1 & -1 & -3 \\ & 0 & -2 \\ & & -2 \end{pmatrix}$	$\begin{pmatrix} 1 & 1 & 1 \\ & 0 & 0 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} 3 & -3 & -1 \\ & -6 & -4 \\ & & 2 \end{pmatrix}$
$\begin{pmatrix} -1 & -1 & 3 \\ & 0 & 4 \\ & & 4 \end{pmatrix}$	$\begin{pmatrix} 1 & -3 & 1 \\ & -4 & 0 \\ & & 4 \end{pmatrix}$	$\begin{pmatrix} 3 & 1 & 1 \\ & -2 & -2 \\ & & 0 \end{pmatrix}$
$\begin{pmatrix} -1 & -1 & -1 \\ & 0 & 0 \\ & & 0 \end{pmatrix}$	$\begin{pmatrix} 1 & -3 & 3 \\ & -4 & 2 \\ & & 6 \end{pmatrix}$	$\begin{pmatrix} 3 & 1 & -1 \\ & -2 & -4 \\ & & -2 \end{pmatrix}$
$\begin{pmatrix} -1 & -1 & -3 \\ & 0 & -2 \\ & & -2 \end{pmatrix}$	$\begin{pmatrix} 1 & 3 & 1 \\ & 2 & 0 \\ & & -2 \end{pmatrix}$	$\begin{pmatrix} 3 & -1 & -1 \\ & -4 & -4 \\ & & 0 \end{pmatrix}$

## (iv) $U_c(3) \times U(1) \times U_L(2) \times U(1)^2$ model

Fermion embedding into matrix:

$$\psi = \begin{pmatrix} o & o & q & u & d \\ & o & l & \nu & e \\ & & o & o & o \\ & & & o & o \\ & & & & o \end{pmatrix}$$

$T^6$  flux must take values

$$q^{ab} = \begin{pmatrix} 0 & 0 & -3 & 3 & 3 \\ & 0 & -3 & 3 & 3 \\ & & 0 & 0 & 0 \\ & & & 0 & 0 \\ & & & & 0 \end{pmatrix}$$

**No solution for  $T^2$  flux**

## **§4 Probability distribution**

**Semiclassical analyses of MM dynamics**

## Unitary MM

$$S_b = -\beta\mathcal{N} \sum_{i \neq j} \mathcal{Z}_{ji} \operatorname{tr} \left( \nu_i \nu_j \nu_i^\dagger \nu_j^\dagger \right) + \dots \quad (2)$$

with

$$\mathcal{Z}_{45} = \exp \left( 2\pi i \frac{s_1}{N_1} \right), \quad \mathcal{Z}_{67} = \exp \left( 2\pi i \frac{s_2}{N_2} \right), \quad \mathcal{Z}_{89} = \exp \left( 2\pi i \frac{s_3}{N_3} \right),$$

# Background configurations:

$$\mathcal{V}_i = \mathbb{1} \otimes V_i$$

with

$$\begin{aligned}
 V_{3+i} &= \left( \begin{array}{cccc} \Gamma_{1,i}^1 \otimes \mathbb{1}_{n_2^1} \otimes \mathbb{1}_{n_3^1} \otimes \mathbb{1}_{p^1} & & & \\ & \Gamma_{1,i}^2 \otimes \mathbb{1}_{n_2^2} \otimes \mathbb{1}_{n_3^2} \otimes \mathbb{1}_{p^2} & & \\ & & \dots & \\ & & & \Gamma_{1,i}^h \otimes \mathbb{1}_{n_2^h} \otimes \mathbb{1}_{n_3^h} \otimes \mathbb{1}_{p^h} \end{array} \right) \\
 V_{5+i} &= \left( \begin{array}{cccc} \mathbb{1}_{n_1^1} \otimes \Gamma_{2,i}^1 \otimes \mathbb{1}_{n_3^1} \otimes \mathbb{1}_{p^1} & & & \\ & \mathbb{1}_{n_1^2} \otimes \Gamma_{2,i}^2 \otimes \mathbb{1}_{n_3^2} \otimes \mathbb{1}_{p^2} & & \\ & & \dots & \\ & & & \mathbb{1}_{n_1^h} \otimes \Gamma_{2,i}^h \otimes \mathbb{1}_{n_3^h} \otimes \mathbb{1}_{p^h} \end{array} \right) \\
 V_{7+i} &= \left( \begin{array}{cccc} \mathbb{1}_{n_1^1} \otimes \mathbb{1}_{n_2^1} \otimes \Gamma_{3,i}^1 \otimes \mathbb{1}_{p^1} & & & \\ & \mathbb{1}_{n_1^2} \otimes \mathbb{1}_{n_2^2} \otimes \Gamma_{3,i}^2 \otimes \mathbb{1}_{p^2} & & \\ & & \dots & \\ & & & \mathbb{1}_{n_1^h} \otimes \mathbb{1}_{n_2^h} \otimes \Gamma_{3,i}^h \otimes \mathbb{1}_{p^h} \end{array} \right)
 \end{aligned}$$

- By plugging these configurations into the MM action (2),

$$S_b = -2\beta\mathcal{N}N_4N_5 \sum_{l=1}^3 \sum_{a=1}^h n_1^a n_2^a n_3^a p^a \cos \left( 2\pi \left( \frac{s_l}{N_l} + \frac{m_l^a}{n_l^a} \right) \right)$$

- Using the relation of the Morita equivalence,

$$\frac{s_l}{N_l} + \frac{m_l^a}{n_l^a} = \frac{q_l^a}{N_l n_l^a} = -\frac{1}{2r} \left( \frac{1}{N_l} - \frac{1}{n_l^a} \right)$$

$\Rightarrow$  The classical action takes the minimum value if and only if

$$q_l^a = 0 \Leftrightarrow n_l^a = N_l$$

for  $\forall a$  and  $\forall l$ .

## Small fluctuations around the minimum

- Configurations with  $|q_l^a| \ll N_l$
- The classical action is approximated as

$$\Delta S_b \simeq 4\pi^2 \beta \mathcal{N}^2 \sum_{l=1}^3 \frac{1}{(N_l)^4} \sum_{a=1}^h p^a (q_l^a)^2 \quad (3)$$

~ Instanton action



- **Comparing the IIB MM (1) and the unitary MM (2)**  
by using correspondence between Hermitian and unitary matrices:

$$\mathcal{V}_M \sim \exp \left( 2\pi i \frac{A_M}{\epsilon N_l} \right)$$

one finds a relation between the coupling constants

$$\frac{1}{2} \beta \mathcal{N} \left( \frac{2\pi}{\epsilon N_l} \right)^4 = \frac{1}{g_{\text{IIBMM}}^2}$$

⇒ The instanton action (3) becomes

$$\Delta S_b = \frac{A}{2\pi^2} \sum_{l=1}^3 \sum_{a=1}^h p^a (q_l^a)^2 .$$

with

$$A = \frac{\epsilon^4 \mathcal{N}}{g_{\text{IIBMM}}^2}$$

⇒ Scaling limits of fixing  $g_{\text{IIBMM}}^2 \mathcal{N}^\alpha / \epsilon^4$  with

(1)  $\alpha < -1 \Rightarrow A = 0$

(2)  $\alpha = -1 \Rightarrow A = \text{finite}$

(3)  $\alpha > -1 \Rightarrow A = \infty$

# How to take large-N limits and probability distribution over the string vacuum space

(1)  $g_{\text{IIBMM}}^2 \mathcal{N}^\alpha / \epsilon^4$  with  $\alpha < -1$

- All the topological sectors appear with equal probabilities.
- The estimation for the probability distribution over the string vacuum space reduces to number counting of the classical solutions.

(2)  $\alpha = -1$

- All the topological sectors appear, but with different probabilities  
 $\sim e^{-\Delta S_b}$ .

(3)  $\alpha > -1$

- Only a single topological sector survives.

# The instanton action in each phenomenological model

- Each solution in each phenomenological model specifies  $T^2$  fluxes  $q_l^{ab}$ .
- Since  $q_l^a$  are determined by

$$q_l^{ab} = q_l^a - q_l^b ,$$

$q_l^a$  have an arbitrariness of integer shifts  $q_l$ .

- The minimum value of the instanton actions within this arbitrariness  $q_l$

$$\Delta S_b = \frac{A}{2\pi^2} \sum_{l=1}^3 \sum_{a=1}^h p^a (q_l^a)^2$$

gauge group	integral $q_l$	fractional $q_l$
U(8)	0	0
U(4) $\times$ U <sub>L</sub> (2) $\times$ U <sub>R</sub> (2)	28	24
	44	40
	36	32
	36	32
U <sub>c</sub> (3) $\times$ U(1) $\times$ U <sub>L</sub> (2) $\times$ U <sub>R</sub> (2)	44	43
	40	39
	36	36
	28	27
U(4) $\times$ U <sub>L</sub> (2) $\times$ U(1) <sup>2</sup>	28	25
	40	37
	32	29
	36	32
	44	40
	40	37
	36	36
	40	37
	36	33

- Rather small instanton actions for the phenomenological models
- No substantial differences among the models

## §5 Conclusions and Discussions

## Summary

- We have considered the situations where the IIB MM is compactified on a torus with magnetic fluxes,
- exhausted all the matrix configurations that yield all the models whose gauge group is a subgroup of  $U(8)$ ,
- semiclassically estimated probability distribution of their appearance.



# Discussions

- Higgs field:

- Gauge fields in the extra dimensions give scalar fields on our spacetime

$$V_i = \begin{pmatrix} 0 & 0 & 0 \\ & 0 & \text{Higgs} \\ & & 0 \end{pmatrix}$$

- It is difficult to keep scalar fields massless against quantum corrections.  
**the naturalness/hierarchy problem**
  - \* How to realize **SUSY** and how to break it.
  - \* **Other resolutions** of the naturalness/hierarchy problem
- How the **electroweak symmetry breaking** occurs.
- Values of the **Yukawa couplings** and the flavor structures

- **Anomaly cancellations**

- Extra  $U(1)$  gauge groups  $\Rightarrow$  Anomaly
- May be canceled via the Green-Schwarz mechanism by exchange of the RR-fields

- **Extend the matrix configurations**

- by introducing the Wilson lines and tilting the tori

- **Compactifications**

- Other compactifications than tori
- Compactifications in MM:  
Matter and gravitons must propagate in the compactified space.

- **Ultimately,  
we hope to analyze full dynamics in the MM,  
and survey the probability distribution over the whole  
of the landscape!**

**Thank you**