

# Soft walls

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Based on collaboration with:

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Outline

Introduction to  
SW

The model

Graviton  
fluctuations

Radion  
fluctuations

The Higgs  
background

EW5W

Conclusion

## Outline

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The outline of this talk is

## Outline

- ▶ Introduction
- ▶ The soft-wall model
- ▶ Graviton fluctuations
- ▶ Radion fluctuations
- ▶ The Higgs background
- ▶ EWSB
- ▶ Conclusion

# INTRODUCTION

- ▶ Warped extra dimensions are useful to solve long-standing problems: hierarchy, flavor,...
- ▶ Also the AdS/CFT correspondence might deal with non-perturbative theories: technicolor, QCD,...
- ▶ We will concentrate on general 5D theories with a metric

## Proper coordinates

$$ds^2 = e^{-2A(y)} \eta_{\mu\nu} dx^\mu dx^\nu + dy^2$$

$$\eta_{\mu\nu} = (-, +, +, +, +)$$

or in

## Conformally flat coordinates

$$ds^2 = e^{-2A(z)} (\eta_{\mu\nu} dx^\mu dx^\nu + dz^2), \quad \frac{dz}{dy} = e^A$$

- ▶ The AdS metric  $A_{AdS} = ky$  has conformal invariance
- ▶ The theory requires UV completion which translates into a UV brane at  $y = 0$
- ▶ Conformal invariance has to be broken to generate a mass gap

## IR brane

Conformal invariance is normally broken by an IR brane (RS1<sup>a</sup>) at  $y = y_c$ . It can be stabilized by the GW mechanism<sup>b</sup>: it requires a stabilizing scalar in the gravitational background

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<sup>a</sup>L. Randall and R. Sundrum, hep-ph/9905221

<sup>b</sup>W. Goldberger and M. Wise, hep-ph/9907447

- ▶ There is another way of breaking the conformal invariance if the scalar field has a singularity at  $y = y_s$  which replaces the IR brane

- ▶ We call this a soft wall

## Soft-walls

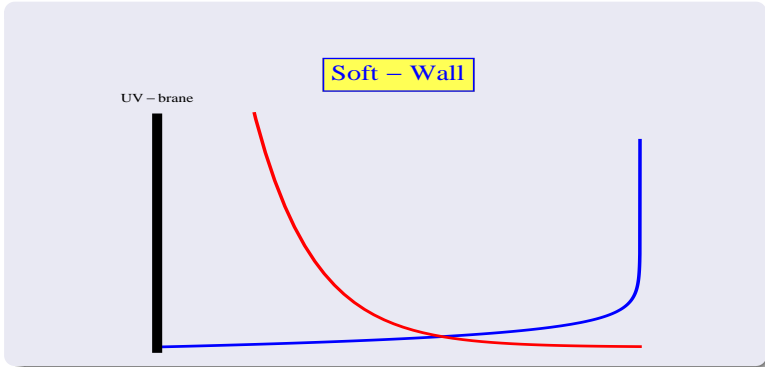
There is no IR brane and the extra dimension is non-compact but of finite length

$$\int e^{-A(z)} dz \equiv y_s = \int_0^{y_s} dy < \infty$$

- ▶ This implies that the IR brane is replaced by a

Naked curvature singularity at  $y = y_s$  where  $A(y_s) \rightarrow \infty$

- ▶ Stabilizing the distance  $y_s$  is similar to stabilizing the brane-to-brane distance  $y_c$  by the GW mechanism



- ▶ The stabilizing field has a divergence at  $y = y_s$
- ▶ The metric backreacts and vanishes at the singularity
- ▶ No IR brane is required

# THE MODEL

- ▶ We will introduce a scalar field  $\phi$  with some boundary condition (BC) at the UV brane @  $y = 0$ :  $\phi_0$
- ▶ We want naturalness to be fulfilled with an **exponential** relation between  $ky_s$  and  $\phi_0$  as

$$ky_s \sim e^{\nu \phi_0}$$

In this way a hierarchy can be **naturally** generated with values  $\nu, \phi_0 \simeq \mathcal{O}(1)$

- ▶ The presence of  $\phi(y)$  will **backreact** on the AdS metric  $A_{AdS} = ky$  providing a modification far from the UV brane. **Solving the exact EOM is required**
- ▶ We will solve the Einstein EOM in the bulk by <sup>1</sup>

"Superpotential" method (non supersymmetric models)

$$A'(y) = W(\phi), \quad \phi'(y) = \partial W / \partial \phi$$

$$V(\phi) = 3(\partial W / \partial \phi)^2 - 12W^2$$

<sup>1</sup>O. DeWolfe, D.Z. Freedman, S.S. Gubser and A. Karch,  
hep-th/9909134

- ▶ The model <sup>2</sup> is defined by the

## Superpotential

$$W(\phi) = k(1 + e^{-\nu\phi})$$

- ▶ It has an analytical solution provided by the

## Gravitational background

$$A(y) = ky - \frac{1}{\nu^2} \log \left( 1 - \frac{ky}{ky_s} \right)$$

$$\phi(y) = \frac{1}{\nu} \log[\nu^2(ky_s - ky)]$$

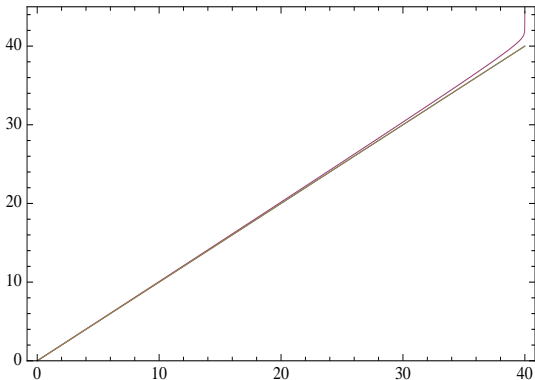
- ▶ The length  $y_s$  is related to the boundary field  $\phi_0$

## Singularity location

$$ky_s = \frac{1}{\nu^2} e^{-\nu\phi_0}$$



- ▶ The metric  $A(y)$  separates from the AdS metric only near the singularity. E.g. for  $\nu = 2$ ,  $k y_s = 40$



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**The model**

## Graviton fluctuations

## Radion fluctuations

## The Higgs background

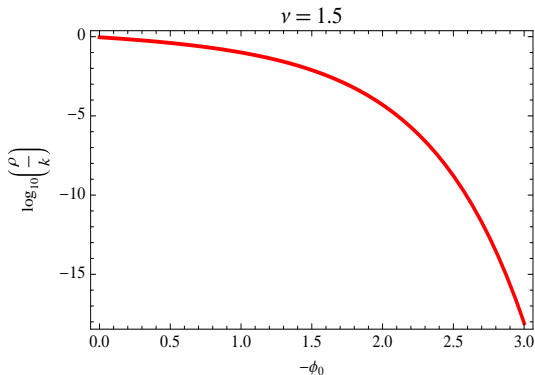
## EWSB

## Conclusion

- ▶ The low scale  $\rho$  can be related to  $\phi_0$  as a double exponential

$$\rho = k \exp \left\{ -\frac{1}{\nu^2} \left( e^{-\nu \phi_0} - \nu \phi_0 - \log \nu^2 \right) \right\}$$

- ▶ A big hierarchy can be naturally obtained with  $|\phi_0| \simeq \mathcal{O}(1)$



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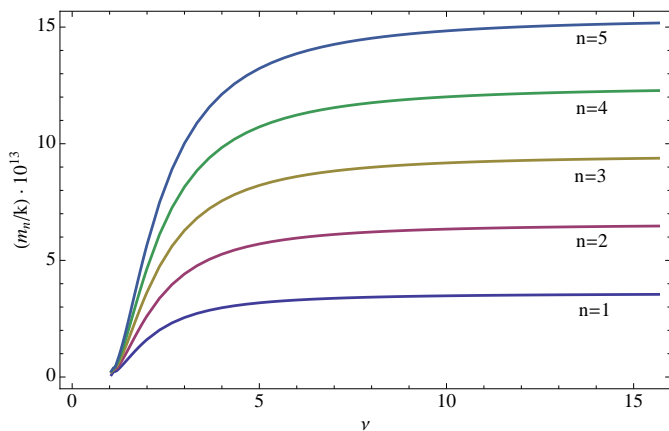
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# GRAVITON FLUCTUATIONS

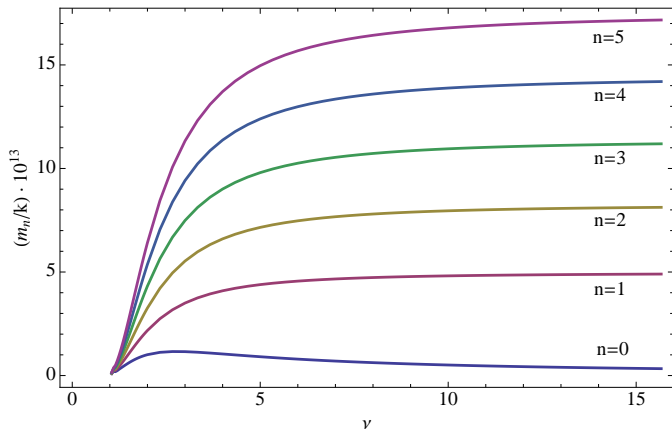
- ▶ We have solved numerically the EOM
- ▶ For  $ky_s = 30$  the first levels mass spectra :



- ▶ Level spacing shrinks in the conformal limit  $\nu \rightarrow 1$  (RS is  $\nu \rightarrow \infty$ )

# RADION FLUCTUATIONS

- ▶ For  $ky_s = 30$  the first levels mass spectra:



The radion/graviton ratio

$$m_{\text{radion}}/m_{\text{grav}}^{(1)} \ll 1$$

# THE HIGGS BACKGROUND

- ▶ In soft-walls there is no IR brane and thus the Higgs has to propagate in the bulk
- ▶ The Higgs doublet in the Standard Model (SM) can be described by

$$H(x, y) = \frac{1}{\sqrt{2}} \begin{bmatrix} 0 \\ h(y) + \xi(x, y) \end{bmatrix} e^{i\vec{\chi}(x, y) \cdot \vec{\sigma}}$$

- ▶ We will assume the Higgs does not perturb the previous mechanism for fixing the radion mass.
- ▶ We then impose the

## Regularity condition

$h(y)$  is regular at  $y_s$

- ▶ In that case the Higgs does not backreact on the gravitational metrics

- ▶ The Higgs background satisfies the equation of motion

## Bulk EOM

$$h'' - 4A'h' = \frac{\partial V}{\partial h}, \quad V(h) = \text{bulk potential}$$

and

## Boundary conditions

$$h'(0) = \left. \frac{\partial \lambda_0}{\partial h} \right|_{y=0}, \quad \lambda_0(h) = \text{UV brane potential}$$

- ▶ I will present a model where EW symmetry is broken at the UV brane

- ▶ EWSB is triggered by the potentials

## UV breaking

$$V(H) = a(a - 4)|H|^2$$

and

$$\lambda_0(H) = M_0|H|^2 + \gamma_0|H|^4$$

- ▶ The bulk EOM is solved by a linear combination of Whittaker functions

$$h(y) = c_W e^{2A} W_{\frac{-4}{(a-2)\nu^2}, \frac{4-\nu^2}{2\nu^2}} [2(a-2)(y_s - y)] \\ + c_M e^{2A} M_{\frac{-4}{(a-2)\nu^2}, \frac{4-\nu^2}{2\nu^2}} [2(a-2)(y_s - y)]$$

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- ▶ Regularity and BC at UV imply (for  $a \geq 2$ )

## BC at UV (for $a \geq 2$ )

$$\text{Regularity} \Rightarrow c_W = 0$$

$$\text{and for } h_0 = c_M M \frac{-4}{(a-2)\nu^2}, \frac{4-\nu^2}{2\nu^2} [2(a-2)y_s]$$

$$h_0^2 \simeq \frac{4 - a - M_0}{\gamma_0}$$

- ▶  $h_0(\gamma_0)$  is fixed by the EW condition

## EWSB condition

$$v_{SM}^2 = \int_0^{y_s} h^2(y) e^{-2A(y)} dy$$

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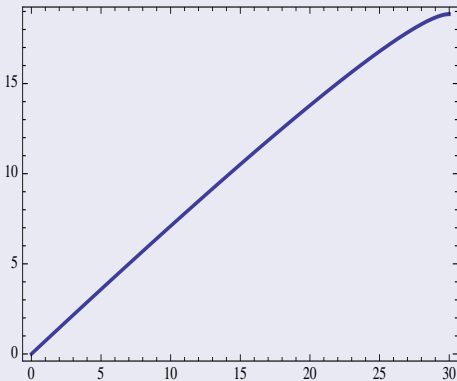
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The Higgs background looks like:

$\log_{10} h(y)/h(0)$  as a function of  $ky$  for  
 $a = 2.1, \nu = 1.5, y_s = 30/k$



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# EWWSB IN THE GAUGE SECTOR

We will illustrate the mechanism with an abelian example

- ▶ The Higgs is defined as

$$H(x, y) = \frac{1}{\sqrt{2}} [h(y) + \xi(x, y)] e^{ig_5 \chi(x, y)}$$

- ▶ The action is invariant under 5D gauge transformations

## 5D gauge transformations

$$A_M(x, y) \rightarrow A_M(x, y) + \frac{1}{g_5} \partial_M \alpha(x, y)$$

$$\chi(x, y) \rightarrow \chi(x, y) + \frac{1}{g_5} \alpha(x, y)$$

- ▶ We will take the 5D gauge condition

$$\partial^\mu A_\mu - m_A^2 \chi + (e^{-2A} A_5)' = 0, \quad m_A(y) = g_5 h(y) e^{-A(y)}$$

- ▶ The 4D theory is invariant under  $\alpha(x) = \alpha(x, y)/f(y)$  gauge transformations and contains:



$$A_\mu(x, y) = \frac{a_\mu(x) \cdot f(y)}{\sqrt{y_s}}$$



$$G(x, y) = m_A^2 \chi - (e^{-2A} A_5)' = \frac{m_f G(x) \cdot f(y)}{\sqrt{y_s}}$$



$$K(x, y) = \chi' - A_5 = \frac{K(x) \cdot \eta(y)}{\sqrt{y_s}}$$

- ▶ With profiles

## Profiles

$$m_f^2 f + (e^{-2A} f')' - m_A^2 f = 0, \quad \text{Neumann BC}$$

$$m_\eta^2 \eta + \left[ m_A^{-2} \left( e^{-2A} m_A^2 \eta \right)' \right]' - m_A^2 \eta = 0, \quad \text{Dirichlet BC}$$

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- ▶ We can find an approximation for the **light gauge boson** mode in the limit where the breaking is small and thus there is a light mode with almost constant profile

## Analytical approximation

$$f_A^0(y) = 1 - \delta_A + \delta f_A(y)$$

$$\delta f_A(y) = \int_0^y dy' e^{2A(y')} \int_0^{y'} dy'' \left[ m_A^2(y'') - m_{f_A^0}^2 \right]$$

$$\delta_A = \frac{1}{y_s} \int_0^{y_s} dy \delta f_A(y)$$

- ▶ The light mode mass

## Mass of light mode

$$m_{f_A^0}^2 = \frac{1}{y_s} \int_0^{y_s} m_A^2(y) dy$$

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# ELECTROWEAK CONSTRAINTS

- ▶ In our 5D model (for fixed values of the parameters  $\nu, y_s, \dots$ ) we have the free parameters ( $g_5, g'_5, h_0, a$ ) which fix the physical spectrum of zero mode masses
- ▶ Once we have fixed the condition  $m_{f_Z} = m_Z^{(ph)}$  the eigenvalue  $m_{f_W}$  is a prediction of the theory
- ▶ The parameter

$$\rho_0 = \frac{m_{f_W}^2}{c_W^2 m_{f_Z}^2} \equiv 1 - \Delta\rho = 1 - s_W^2 \tilde{\delta}_Z$$

can deviate from unity which amounts to a violation of the custodial symmetry (CS)

$$\tilde{\delta}_V = \frac{m_V^2}{k^2} y_s \int_0^{y_s} \left( \Omega - \frac{y}{y_s} \right)^2 e^{2A} dy$$
$$\Omega(y) = \frac{U(y)}{U(y_s)}, \quad U'(y) = h^2(y) e^{-2A(y)}$$

- ▶ We will be assuming here (not necessarily an assumption) that fermions are localized on the UV brane in which case

$$g_V = g_V^{SM} f_V(0) \equiv g_V [1 - \delta_V]$$

- ▶ The latter changes the definition of the Fermi constant measured in the  $\mu$ -decay and the  $Z$  widths which constrain the

## EWPT Parameters

$$\delta_Z = \frac{m_Z^2}{k^2} y_s \int_0^{y_s} dy e^{2A(y)} \left(1 - \frac{y}{y_s}\right) \left(\Omega - \frac{y}{y_s}\right)$$

$$\delta_W = c_W^2 \delta_Z$$

through the observables  $\bar{s}_\ell^2, \Gamma_{\ell+\ell-}, \dots$

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- ▶ We can also express the departure with respect to the SM predictions in the language of the usual parameters  $(S, T, U)$
- ▶ It turns out that

$(S, T, U)$  parameters

$$\alpha(m_Z)T = s_W^2 \tilde{\delta}_Z$$

$$\frac{\alpha(m_Z)}{4s_W^2 c_W^2} S = -2\delta_Z$$

$$\frac{\alpha(m_Z)}{4s_W^2} (S + U) = -2\delta_W$$

- ▶ Or using the relation  $\delta_W = c_W^2 \delta_Z$

$$\alpha(m_Z)T = s_W^2 \tilde{\delta}_Z, \quad \alpha(m_Z)S = -8s_W^2 c_W^2 \delta_Z, \quad \alpha(m_Z)U \simeq 0$$

- ▶ The strongest constraint is on  $T$

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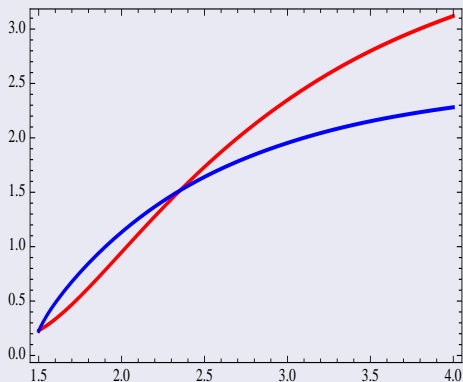
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# NUMERICAL RESULTS

We can obtain a lower bound on  $m_{KK}$  from the bounds on  $T$  and  $S$

$$\alpha T, \alpha S < 10^{-3}, y_s = 35/k, a \simeq 2: m_{KK}/\text{TeV} \text{ Vs } \nu$$



No explicit CS is required!

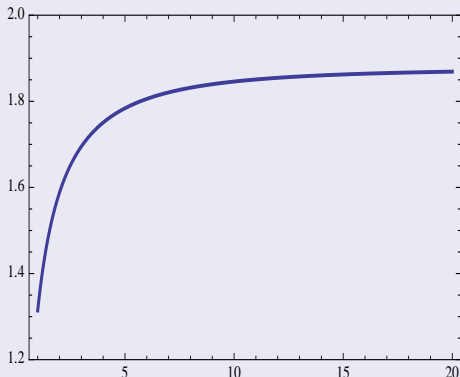


One of the reasons why the bounds go down when  $\nu \rightarrow 1$  is

$$h(y) \sim e^{a_{\text{eff}}(\nu, a)y}$$

and [for RS  $a_{\text{eff}}(\infty, a) = a$ ]

$a_{\text{eff}}$  Vs  $\nu$  for  $a = 2$



and the Higgs profile is **much less IR localized than in the RS case**

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# CONCLUSION

- ▶ We have proposed a set of soft-wall models with AdS geometry near the UV brane
- ▶ The large hierarchy is generated without fine-tuning by a background scalar field
- ▶ The limit  $\nu \rightarrow \infty$  is RS
- ▶ For  $\nu = 1$  the spectrum is continuum above a mass gap  $\rho$  and it can model unparticles
- ▶ We propose models of EWSB with a bulk Higgs
- ▶ Electroweak constraints can be satisfied in a way similar to the SM: no extra custodial symmetry has to be introduced
- ▶ Indirect and direct constraints can be satisfied for KK-masses of  $\mathcal{O}(1)$  TeV
- ▶ The model is Higgsless and KK modes unitarize  $WW$  scattering (Pokorski's talk)
- ▶ One can also break EWS in the bulk with a light Higgs at the price of some fine tuning

## BACKUP SLIDES

- ▶ Expression for  $S, T$  simplify in the RS case as

$(S, T, U)$  parameters

$$\alpha(m_Z) T^{RS} = s_W^2 \frac{m_Z^2}{\rho^2} (ky_s) \frac{(a-1)^2}{a(2a-1)} + \dots$$

$$\alpha(m_Z) S^{RS} = 2s_W^2 \frac{m_W^2}{\rho^2} \frac{a^2 - 1}{a^2} + \dots$$

- ▶  $T$  is volume enhanced while  $S$  is not
- ▶ No analytical expression exists for soft-wall metric