LIGHT PARTICLES COUPLED TO PHOTONS

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CORFU WORKSHOP 2010

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### Motivation: The Axion Model

Consequence of Peccei-Quinn solution to strong CP-problem

 $\star$  Axion mass *m* and breaking scale *f* related

$$m = \frac{f_{\pi}m_{\pi}}{f} \frac{\sqrt{m_u m_d}}{m_u + m_d} = 0.6 \text{ eV} \frac{10^7 \text{ GeV}}{f}$$

★ Couplings to matter (i = p, n, e, etc.)

$$\mathcal{L}_{a\Psi\Psi} = \sum_{i} c_{i} \frac{1}{2f} (\bar{\Psi}_{i} \gamma^{\mu} \gamma_{5} \Psi_{i}) (\partial_{\mu} a)$$

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## Axion Coupling to Two Photons

Axion searches use  $a\gamma\gamma$  coupling

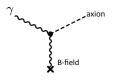


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# Axion-Photon Mixing in a Magnetic Field

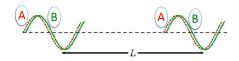


Interaction states different from propagation states:

$$|a'\rangle = \cos \varphi |a\rangle - \sin \varphi |\gamma\rangle$$
$$|\gamma'\rangle = \sin \varphi |a\rangle + \cos \varphi |\gamma\rangle$$

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### Coherent effect among A, B



 $|k_A - k_B| L \ll 2\pi$  (coherence)

## **Axion-Photon Coherence**

Propagation of a photon beam (energy E) traveling distance L

Conversion to axions:

$$Lm^2/E < 1$$
 (coherence)

Probability (in vacuum)

$$P(\gamma 
ightarrow a) = rac{1}{4} g_{a\gamma\gamma}^2 B_T^2 L^2$$

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Other (light) particles coupled to photons

Most experiments looking for axions use the fact that:

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- the mass is small
- it couples to two photons

⇒ Experiments are sensitive to other particles with these properties

# Other (light) particles coupled to photons

- Predicted in BSM models and/or needed
- Scalar particles
- quintessence fields in cosmology
- light spin 2 particles
- Axion has  $f_a m_a \sim f_\pi m_\pi$

But for other particles consider f and m independent parameters

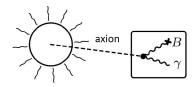
Name:

Axion Like Particles (ALPs)

or Weakly Interacting Sub-eV Particles (WISPs)

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## ALPs from the Sun



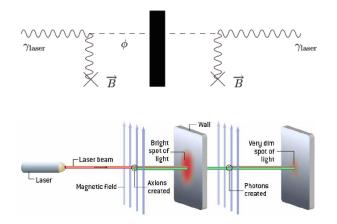
- Produce X-ray photons
- Valid for axion masses where conversion would be coherent
- CAST limit:  $g_{a\gamma\gamma} < 2.2 imes 10^{-10} \ {
  m GeV^{-1}}$  (95 % CL,  $m < 0.4 \, {
  m eV}$ )

#### Caveat:

Models where axion production suppressed at high T Need of Lab exps. (complementary to astro-constraints)

# ALPs in the Laboratory

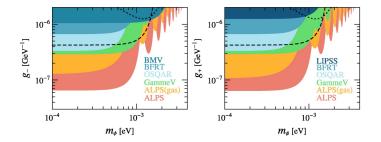
Light-Shining-through-a-Wall (LSW)



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### Experimental State of the Art

### Results from the ALPS collaboration at DESY



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### **New Forces**

Light bosons of mass mediate forces with range  $\lambda = mass^{-1}$ 

$$m \sim 10^{-5} \mathrm{eV} ~\leftrightarrow \lambda \sim 2 \mathrm{~cm}$$

Boson can be a Scalars or a Pseudoscalars

Scalar-mediated force among 1 and 2

$$V_{ss}(r) = rac{g_s^{(1)}g_s^{(2)}}{4\pi r}e^{-mr}$$

At a distances  $r \sim m^{-1}$ 

$$rac{V_{ss}}{V_{grav}}\sim (g^{(N)}_s)^2 rac{M_P^2}{m_N^2}$$

Equivalence principle restrict  $g_N$  to very tiny values

# Scalar-photon couplings and Equivalence Principle

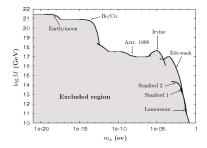


Coupling to protons but not to neutrons

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equivalence principle violation



(g = 1/M)

### **New Forces**

However, for pseudoscalar-exchange there is coupling to spin which weakens very much the new force

$$rac{V_{
m pp}}{V_{
m grav}} \sim (g_{
m p}^{(N)})^2 rac{M_{
m P}^2}{m_N^2} F$$
 $F = \sum_{
m spins} \left[ 7(ec{\sigma_1} \cdot ec{n})(ec{\sigma_2} \cdot ec{n}) - 2(ec{\sigma_1} \cdot ec{\sigma_2}) 
ight]$ 

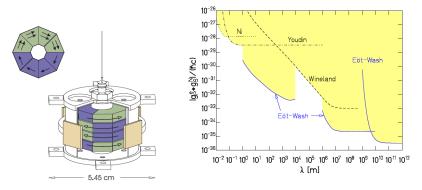
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For unpolarized bodies,  $F \rightarrow 0$ .

- In general, pheno point of view,
- $-S \times S \quad g_s g_s$
- $P \times P \quad g_s g_p$
- $-S \times P \quad g_p g_s$

### New Observational Constraints on New Forces

torsion pendulum containing  $9 \times 10^{22}$  polarized electrons



#### hep-ph/0606218

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Quantum states of neutrons in the Earth's gravitational field

• Bound states exist for neutrons trapped in the Earth's gravitational field.

Quantum mechanical expectation verified experimentally

OK without additional forces,

 $\Rightarrow$  get constraints on additional forces.

• proposal of exp. with polarized neutrons,  $\lambda = 1 - 200 \ \mu m$ 

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### Chamaleons

- Light scalar particles as quintessence
- Induce new force; impact in gravity experiments

The chamaleon idea:

$$V_{\text{eff}}(\phi) = V(\phi) + A(\phi) \rho_m$$

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 $A(\phi)$  arbitrary function,  $\rho_m$  ambient matter density

Particle mass / interaction range depend on  $\rho_m$ 

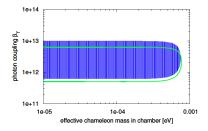
 $\Rightarrow$  Escape fifth-force constraints

## Chamaleons in the laboratory

Chameleons have generic coupling to photons  $\beta_{\gamma} = M_{Planck}/M$ 

Idea of detection:

- 1) Produce chamaleons in a magnetic field
- 2) Wait for desintegration into photons



We are testing models of Dark Energy in the lab !

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### Conclusions

- Many ideas in the subject of light particles coupled to photons
- New experiments undergoing
- The low-energy frontier of New Physics

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