#### UV BEHAVIOR OF THE EINSTEIN-YANG-MILLS SYSTEM

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# INTRODUCTION - RECENT RESULTS

Are there gravity corrections to the beta function of the Yang-Mills (YM) coupling constant?

perturbative calculations by several authors:

[Robinson and Wilczek (2006), Pietrykowski (2007), Toms (2007), Ebert, Plefka, Rodigast (2008)]

- only the former find a non-trivial result
- problematic either due to dimensional or cutoff regularization
- Tang and Wu ("loop regularization") find a non-zero correction [Tang and Wu (2008)]

# INTRODUCTION - THE ERG APPROACH

We analyzed the running of the gauge coupling constant in the Asymptotic Safety scenario for quantum gravity.

Main tools:

- FRGE for the effective average action
- the background field method

Advantages:

- sensitive to quadratic divergences
- background field method preserves gauge invariance Nevertheless:
  - resulting beta-function may still be gauge fixing dependent

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L THE YM-GRAVITY SYSTEM IN THE ERG APPROACH

TRUNCATION OF THE EFFECTIVE AVERAGE ACTION

#### THE TRUNCATION

$$\Gamma_k = \Gamma_k^{\rm EH} + \Gamma_k^{\rm YM} + \Gamma_k^{\rm gf} + S_{\rm gh} = \breve{\Gamma}_k + S_{\rm gh}$$

with

$$\Gamma_k^{\rm EH}[g] = \frac{Z_N(k)}{16\pi\hat{G}} \int \mathrm{d}^4 x \sqrt{g} \left(-R(g) + 2\bar{\lambda}(k)\right)$$
  
$$\Gamma_k^{\rm YM}[g,A] = \frac{Z_F(k)}{4\,\hat{g}_{\rm YM}^2} \int \mathrm{d}^4 x \sqrt{g} \, g^{\mu\rho} g^{\nu\sigma} F^a_{\mu\nu} F^a_{\rho\sigma}$$
  
$$\Gamma_k^{\rm gf}[\bar{h},\bar{a};\bar{g},\bar{A}] = \int \mathrm{d}^4 x \, \sqrt{\bar{g}} \, \left(\frac{Z_N(k)}{2\alpha_{\rm D}} \bar{g}^{\mu\nu} \mathrm{F}_{\mu} \mathrm{F}_{\nu} + \frac{Z_F(k)}{2\alpha_{\rm YM}} \mathrm{G}^a \mathrm{G}^a\right)$$

Motivation:

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- $\Gamma_k^{\rm EH}$  contains the all essential features of gravity close to the non-Gaussian fixed point (NGFP)
- ▶ in pure YM theory  $\Gamma_k^{\rm YM}$  approximates the perturbative 2-loop result within a few percent

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# GAUGE FIXING AND GHOST ACTION

Gauge conditions:

$$\mathcal{F}_{\mu}(\bar{h};\bar{g}) = \frac{1}{\sqrt{16\pi\,\hat{G}}} \left(\delta^{\beta}_{\mu}\bar{g}^{\alpha\gamma}\bar{D}_{\gamma} - \frac{1}{2}\bar{g}^{\alpha\beta}\bar{D}_{\mu}\right)\bar{h}_{\alpha\beta}$$

$$\mathbf{G}^{a}(\bar{a};\bar{g},\bar{A})=\hat{g}_{\mathrm{YM}}^{-1}\,\bar{g}^{\mu\nu}\bar{\mathcal{D}}_{\mu}\bar{a}_{\nu}^{a}$$

#### $S_{\rm gh}$ is then obtained by the Faddeev-Popov method

Further simplifications:

▶ neglect renormalization effects in the ghost sector → fluctuations in  $S_{\rm gh}$  are set to zero

• gauge parameters chosen to  $\alpha_{\rm D} = 1 = \alpha_{\rm YM}$ 

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FUNCTIONAL RENORMALIZATION GROUP EQUATION

$$\partial_t \Gamma_k = \frac{1}{2} \mathrm{STr} \Big[ \Big( \Gamma_k^{(2)} + \mathcal{R}_k(\Delta) \Big)^{-1} \partial_t \mathcal{R}_k(\Delta) \Big]$$

Inserting the truncation we have

$$\partial_t \Gamma_k = \frac{1}{2} \operatorname{Tr} \left[ \frac{\partial_t \mathcal{R}_k(\breve{\Delta})}{\breve{\Gamma}_k^{(2)} + \mathcal{R}_k(\breve{\Delta})} \right] - \operatorname{Tr} \left[ \frac{\partial_t \mathcal{R}_k^{\mathrm{gh}}(\Delta_{\mathrm{gh}})}{S_{\mathrm{gh}}^{(2)} + \mathcal{R}_k^{\mathrm{gh}}(\Delta_{\mathrm{gh}})} \right]$$

where

$$\mathcal{R}_k(x) = \mathcal{Z}_k k^2 R^{(0)}(x/k^2), \qquad \mathcal{R}_k^{\mathrm{gh}}(x) = \mathcal{Z}_k^{\mathrm{gh}} k^2 R^{(0)}(x/k^2)$$

- ▶ shape function  $R^{(0)}(y)$  with  $R^{(0)}(0) = 1$  and  $\lim_{y \to \infty} R^{(0)}(y) = 0$
- ►  $\mathcal{Z}_k$  and  $\mathcal{Z}_k^{\text{gh}}$  constant matrices in field space, chosen such that modes of  $\Gamma_k^{(2)} \to \Gamma_k^{(2)} + \mathcal{R}_k$  get shifted  $\zeta_k p^2 \to \zeta_k (p^2 + k^2 R^{(0)})$

UV BEHAVIOR OF THE EINSTEIN−YANG-MILLS SYSTEM ☐ THE YM-GRAVITY SYSTEM IN THE ERG APPROACH ☐ FUNCTIONAL RENORMALIZATION GROUP EQUATION

# Spectrally Adjusted Cutoff Operator

As cutoff operator we choose

$$\breve{\Delta} = \mathcal{Z}_k^{-1} \breve{\Gamma}_k^{(2)} \qquad \Delta_{\rm gh} = \mathcal{Z}_{\rm gh}^{-1} S_{\rm gh}^{(2)}$$

This has been called "spectrally adjusted" or "type III" cutoff in the literature.

- ► conceptually not as clear as the covariant Laplacian  $\square$  $\leftrightarrow$  additional terms on RHS due  $\partial_t$  acting on  $\breve{\Delta}$
- ► admits a simple spectral representation of the RHS → simplifies evaluation of the traces (no operator inversion necessary)

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## CALCULATION AND RG IMPROVEMENT

Calculational steps:

- $\blacktriangleright$  compute the Hessian  $\Gamma_k^{(2)}$  and  $S_{\rm gh}^{(2)}$
- equate background and full classical fields  $(\bar{g}_{\mu\nu} = g_{\mu\nu}, \ \bar{A}^a_\mu = A^a_\mu)$
- expand RHS in  $A^a_\mu$  to second order to extract the  $(F^a_{\mu\nu})^2$ -term (we may set  $g_{\mu\nu} = \delta_{\mu\nu}$ )
- ▶ all terms combine to an F<sup>2</sup>-contribution due to the built-in gauge invariance

Different degrees of RG improvement:

- ▶  $\partial_t$  acts only on explicit k-dependence  $\leftrightarrow$  1-loop calculation
- $\partial_t$  acts in addition on  $\mathcal{Z}_k$ -factors
- $\partial_t$  acts also on  $\check{\Gamma}^{(2)}_k$  in  $\check{\Delta}$

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#### 1-LOOP RESULT

Switching to dimensionless couplings

$$g_{\rm YM}^2(k) \equiv \frac{\hat{g}_{\rm YM}^2}{Z_F(k)}, \quad g(k) \equiv k^2 \frac{\hat{G}}{Z_N(k)}, \quad \lambda(k) \equiv k^{-2} \bar{\lambda}(k)$$

we obtain the 1-loop result

$$\partial_t g_{\rm YM}^2 = -\frac{6\,\Phi_1^1(0)}{\pi} \, g \, g_{\rm YM}^2 - \frac{11\,N}{24\,\pi^2} \, g_{\rm YM}^4$$

[Daum, U.H., Reuter (2009)]

where

$$\Phi_1^1(w) = \int_0^\infty \mathrm{d}z \, \frac{R^{(0)}(z) - zR^{(0)\prime}(z)}{z + R^{(0)}(z) + w}$$

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# CLASSICAL REGIME

Newton's constant

$$G(k) \approx G_0 = \text{const}, \quad g(k) = G_0 k^2$$

For an Abelian field (N = 0) we obtain

$$\partial_t g_{\rm YM}^2 = -\frac{6 \, \Phi_1^1(0)}{\pi} \, G_0 \, k^2 \, g_{\rm YM}^2$$

with solution

$$\begin{split} g_{\rm YM}^2(k) &= g_{\rm YM}^2(0) \cdot \exp\left(-\omega_{\rm YM}(k/m_{\rm Pl})^2\right) \\ &= g_{\rm YM}^2(0) \cdot \left[1 - \omega_{\rm YM}(k/m_{\rm Pl})^2 + O(k^4/m_{\rm Pl}^4)\right] \\ \end{split}$$
 where  $m_{\rm Pl} = G_0^{-1/2}$  and  $\omega_{\rm YM} = 3\Phi_1^1(0)/\pi$ .

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LDISCUSSION OF THE RESULT

#### Asymptotic Safety

Free Maxwell field does not destroy NGFP of the EH truncation. Therefore in the UV  $% \mathcal{A}^{(1)}$ 

$$g(k) 
ightarrow g^* \quad \Longrightarrow \quad G(k) = g^*/k^2 
ightarrow 0 \quad {\rm as} \quad k 
ightarrow \infty$$

which implies

$$\partial_t g_{\rm YM}^2 = - rac{6 \, \Phi_1^1(0)}{\pi} \, g^* \, g_{\rm YM}^2$$

with solution

$$g_{\rm YM}^2(k) \propto k^{-\Theta_{\rm YM}}, \quad \Theta_{\rm YM} = \frac{6 \, \Phi_1^1(0)}{\pi} \, g^*$$

- $\blacktriangleright$  total system has a NGFP with  $(g^*_{\rm YM}=0,g^*>0,\lambda^*>0)$
- gravity speeds up approach of asymptotic freedom in the YM sector (power law instead of logarithmic)

# SUMMARY

- we obtain a gravitational correction to the Yang-Mills beta function in a setting which both preserves gauge invariance and retains quadratic divergences
- ► the coefficient of the correction is scheme and gauge fixing dependent, but g<sub>YM</sub>(k) is not an observable quantity
- the form of the correction corresponds to the result of [Tang and Wu (2008)]
- the result is consistent with the Asymptotic Safety scenario for QEG, with vanishing gauge coupling at the NGFP
- RG improvements do not change the picture substantially