Geometric flows: string cosmology, gravitational instantons and non-relativistic gravity

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Highlights

Framework and motivations

Reminder on Ricci flows

Strings, sigma models and RG flows

Gravitational instantons and geometric flows

Geometric flows in non-relativistic gravity

Last word

Geometric flows

Irreversible continuous evolution of Riemanian geometry $g_{ij}(t)$ – *non-linear parabolic equation driven by some tensor* $S_{ij}[g_{ij}]$

$$\frac{\partial g_{ij}}{\partial t} = S_{ij}$$

- Plethora of flows: Ricci, Einstein, Calabi (even D), Cotton (3D), Bach (4D), ...
- Plethora of behaviours: infinite or finite time, convergence towards canonical metrics, singular end-points with topology change, ...

Ricci flows: $S_{ij} = -R_{ij}$

- Introduced by R. Hamilton in 1982 as a tool for proving Poincaré's (1904) and Thurston's (late 70s) 3D conjectures
- Driven by Hamilton's programme culminating in Perel'man's proof (2002–03)
- Appeared to be relevant in physics

Ricci flows and generalizations arise in problems related to gravity with time foliation

- In non-critical string theory Ricci flow can mimic time evolution as an RG flow towards the IR – cosmology
- In 4D self-dual gravitational instantons with homogeneous spatial sections: time evolution is a Ricci flow of the 3D homogeneous leaves
- In non-relativistic gravity with covariance explicitly broken to foliation-preserving diffeomorphisms and with detailed-balance dynamics: time evolution is a geometric flow of the 3D space

Aim: review these properties

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Basic features

Ricci-flow equations

$$\frac{\partial g_{ij}}{\partial t} = -R_{ij}$$

Volume is not preserved along the flow

 $\frac{\mathrm{d}V}{\mathrm{d}t} = \frac{1}{2} \int \mathrm{d}^D x \sqrt{\det g} g^{ij} \frac{\partial g_{ij}}{\partial t} = -\frac{1}{2} \int \mathrm{d}^D x \sqrt{\det g} R$

Consequence:

- ► positive curvature → space contracts
- negative curvature \rightarrow space expands
- Killing vectors are preserved in time: the isometry group can only grow or remain unaltered

Example

- At initial time: $R_{ii}^{(0)} = ag_{ii}^{(0)}$ with a constant
- Subsequent evolution: linear rescaling

$$egin{aligned} \mathsf{g}_{ij}(t) &= (1 - \mathsf{a}t) \mathsf{g}_{ij}^{(0)} \ \mathsf{R}_{ij}(t) &= \mathsf{R}_{ij}^{(0)} \end{aligned}$$

Properties

- $a > 0 \Rightarrow$ uniform contraction \rightarrow singularity at t = 1/a
- $a < 0 \Rightarrow$ indefinite expansion

Rôle in Poincaré's and Thurston's conjectures

Poincaré's conjecture: any closed 3-manifold with trivial fundamental group is homeomorphic to a 3-sphere Similar theorem proven for D > 4 in the 60s and for D = 4 in 1982

Thurston's geometrisation conjecture: extension when the fundamental group is not trivial

- In direct relation with the classification of 3-manifolds (obtained e.g. as quotients by discrete isometry groups)
- Consequence: any 3-manifold can be decomposed in locally homogeneous components

Ricci flow is governed by a non-linear heat-like equation

Smoothes the initial geometry and brings it to a simpler form where the conjectures can be checked

Singularities and degeneracies can appear (e.g. pinching cycles) with topology changes

Must be kept under control: Perel'man's "surgery"

Remarks

To avoid the trivial singularity of shrinking volume: normalised flow

$$rac{\partial g_{ij}}{\partial t} = -R_{ij} + rac{\langle R
angle}{D}g_{ij}$$

Consequence: $\frac{dV}{dt} = 0$ although the eqs. are identical to the ordinary ones after appropriate rescaling of g_{ij} and t

Other generalisations to the Ricci flow exist – Perel'man's flow is not pure Ricci: an extra scalar participates

The case of homogeneous 3-manifolds

Why?

- Appear in the final stage of Hamilton's programme for Thurston's geometrisation
- Building blocks for
 - Lorentzian 4-manifolds (cosmological solutions)
 - Euclidean 4-manifolds (gravitational instantons)
- Ricci flows can be studied by analytic methods sometimes related to remarkable integrable systems

Spaces M_3 admitting a group G_{hom} acting transitively [in 3D see Scott, 1983]

Cosets or products thereof

 $\begin{array}{l} H_3, H_2 \times S^1, S^2 \times S^1\\ H_3 = SL(2,\mathbb{C})/SU(2), H_2 = SL(2,\mathbb{R})/U(1), S^2 = SU(2)/U(1)\\ (\dim G_{\mathrm{hom}} = 6, 4, 4 > 3 \Rightarrow \mathrm{action} \ multiply \ \mathrm{transitive}) \end{array}$ $\bullet \ \dim G_{\mathrm{hom}} = 3 \Rightarrow \mathrm{action} \ simply \ \mathrm{transitive:} \ \mathcal{M}_3 \ \mathrm{is} \ \mathrm{locally} \ G_{\mathrm{hom}}\\ \bullet \ 3 \ \mathrm{linearly} \ \mathrm{independent} \ \mathrm{Killing} \ \mathrm{vectors} \ \mathrm{tangent} \ \mathrm{to} \ \mathcal{M}_3:\\ [\xi_i, \xi_j] = c^i_{\ jk} \xi_k\\ \bullet \ \mathrm{left-invariant} \ \mathrm{Maurer-Cartan} \ \mathrm{forms} \ \sigma^i: \ \mathrm{d}\sigma^i = \frac{1}{2} c^i_{\ ik} \sigma^j \wedge \sigma^k \end{array}$

$$\blacktriangleright \ c^{k}_{ij} = -\epsilon_{ij\ell} m^{\ell k} + \delta^{k}_{i} a_{j} - \delta^{k}_{j} a_{i} \Rightarrow c^{i}_{ij} = 2a_{j}$$

3D Ghom are Bianchi groups [Bianchi 1897; Taub, 1951]

Bianchi classes

class A unimodular: I, II, $VI_{h=-1}$, $VII_{h=0}$, VIII, IX T_3 , Heisenberg, $E_{1,1}$, E_2 , $SL(2, \mathbb{R})$, SU(2)class B non-unimodular: III, IV, V, $VI_{h\neq-1}$, $VII_{h\neq0}$

Geometry

- ► The most general metric: $ds^2 = g^{ij}\sigma^i\sigma^j = \delta_{ij}\theta^i\theta^j$ $\theta^i = \Theta^i{}_i\sigma^j, g_{ij} = \delta_{k\ell}\Theta^k{}_i\Theta^\ell{}_i (\Theta^i{}_i \text{ are coordinate-independent})$
- ► Minimalistic (diagonal) ansatz $\Theta^{i}_{j} = \gamma_{j} \delta^{i}_{j}$: $ds^{2} = \sum_{i} \gamma_{i} (\sigma^{i})^{2}$

not always the most general if the σ^i s are in a canonical form

Behaviour under Ricci flow

Bianchi IX class: 3-spheres

• Left-invariant Maurer–Cartan forms of $G_{hom} \equiv SU(2)$

$$\left\{egin{aligned} \sigma^1 &= \sin artheta \sin \psi \, \mathrm{d} arphi + \cos \psi \, \mathrm{d} artheta \ \sigma^2 &= \sin artheta \cos \psi \, \mathrm{d} arphi - \sin \psi \, \mathrm{d} artheta \ \sigma^3 &= \cos artheta \, \mathrm{d} arphi + \mathrm{d} \psi \end{aligned}
ight.$$

with $0 \le \vartheta \le \pi$, $0 \le \varphi \le 2\pi$, $0 \le \psi \le 4\pi$ (Euler angles) $d\sigma^i + \frac{1}{2} \epsilon^i{}_{ik} \sigma^j \wedge \sigma^k = 0$

- The diagonal ansatz is the most general
 - If two γ s are equal the isometry group is promoted to $SU(2) \times U(1)$ (axial symmetry)
 - Full isotropy requires γ s be *all* equal

The Ricci-flow equations ($^{\cdot} = d/d\tau = \gamma_1\gamma_2\gamma_3 d/dt$)

$$2rac{\gamma_1}{\gamma_1}=(\gamma_2-\gamma_3)^2-\gamma_1^2$$
 and cyclic perms.

Typical behaviour

At large T $\gamma_i \approx 1/\sqrt{\tau} \Rightarrow$ convergence toward the round sphere of vanishing radius

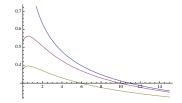


Figure: Behaviour of γ_i s for Bianchi IX

Remarks

• If torsion is added (H = dB), the flow

$$\frac{\partial g_{ij}}{\partial t} = -R_{ij}^{(-)}$$

converges towards a finite-radius round sphere

 Other Bianchi classes of homegeneous 3-manifolds exhibit different behaviours (absence of convergence, pancake degeneracies, cigar degeneracies, ...) [Isenberg, Jackson, 1992]

A parenthesis: 19th century integrable systems

Bianchi IX Ricci-flow equations are a remarkable integrable system

[Sfetsos, unpubl.; Bakas, Orlando, Petropoulos, 2006; Bourliot, Estes, Petropoulos, Spindel, 2009]

- Darboux equations on "triply orthogonal surfaces" solved by Halphen using modular forms [Darboux 1878; Halphen, 1881]
- Darboux-Halphen system studied extensively by mathematicians over the recent years [Takhtajan, 1992; Maciejewski, Strelcyn, 1995; Chakravarty, Halburd, Ablowitz, 2003]
- Darboux-Halphen appeared in the framework of Bianchi IX self-dual gravitational instantons and in the scattering of SU(2) BPS monopoles [Gibbons, Pope, 1979; Manton, 1981; Atiyah, Hitchin, 1985; Gibbons, Manton, 1986]

Typical system appearing in general self-dual Yang-Mills reductions – *in the late '70s all integrable systems were even thought to be SDYM reductions* [Ward, 1985] – *now in geometric flows*

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String propagation in curved backgrounds

Non-linear sigma models (g_{MN} Lorentzian or Euclidean)

$$S = \frac{1}{2\pi\alpha'} \int d^2 z \, g_{MN}(x) \partial x^M \bar{\partial} x^N$$

potentially B_{MN} (axion), Φ (dilaton)

Not scale-invariant: g_{MN} flows under RG [Friedan, 1985]

µ: RG 2D mass scale

 $\mu^{-1\partial/\partial\mu^{-1}}g_{MN} = -\beta [g_{MN}] \equiv -R_{MN} + O(\alpha')$ $\bullet \ t = -\log\mu; \ \text{RG time} \begin{cases} UV: t \to -\infty \\ IR: t \to +\infty \end{cases}$ $\text{RG flow} \equiv \text{Ricci flow}$

► \exists fixed-points – possibly exact 2D CFTs (all-order α')

Application: string cosmology

Original aim: find time-dependent string backgrounds – exact or at $O(\alpha')$ – address the usual problems

Keep in mind

- Around the big-bang: high curvature \rightarrow exact solutions
- After the inflation era: adiabatic evolution → O(α') enough higher O(g_s) due to V(Φ)

plus moduli problem, phase transitions, ...

The FRWL paradigm: assume $M_4 = T \times M_3$ *with* M_3 *homogeneous (even maximally symmetric)*

$$\mathrm{d}s^2 = -\mathrm{d}t^2 + \exp 2\sigma(t)\mathrm{d}\Omega^2$$

- ▶ In general relativity: fluid, Λ_c , ... → FRWL eqs. for σ , p, q
- In string theory: less freedom
 - moduli, fluxes, ...
 - internal manifold
 - curvature and dilaton set the validity of approximations

Example: linear-dilaton solution (exact)

$$\begin{aligned} {\rm d} s^2 &= -{\rm d} t^2 + {\rm d} \Omega_{S^3}^2 + {\rm d} s_{\rm int}^2 \\ \Phi &= \Phi_0 - \bar Q t, \; H = 2 \omega_{S^3} \end{aligned}$$

 $(a(\tau) = \bar{Q}^2 \tau^2$ in Einstein frame) Usually appears in limiting regimes

Central question

Could RG flow mimic cosmological evolution with IR fixed points being steady-state universes?

- Irreversibility and averaging hypothesis [Carfora, Piotrkowska, 1995]
- Hints from Liouville theory in non-critical strings [1989–1991: Das, David, Kawai, Polyakov, Sen, Wadia, ...]
- Popular, not a priori justified
 - RG-flow equations are 1st-order in time: $\partial g_{ij}/\partial t = -\beta \left[g_{ij} \right]$
 - genuine evolution is second-order: $\beta \left[g_{\mu\nu}\right] = 0$

Analysed in various contexts to set under which circumstances this could be valid [Tseytlin, 1992; Schmidhuber, Tseytlin, 1994; Bakas, Orlando, Petropoulos, 2006]

Strategy

Promote a flowing 3D spatial sigma model

 $g_{ij}\partial x^i\bar\partial x^j$

into a scale-invariant 4D space–time sigma model with a genuine time coordinate along which the space–time is foliated

$$-\partial t\bar{\partial}t + g_{ij}(t)\partial x^{i}\bar{\partial}x^{j} + R_{2}\Phi(t)$$

• Write $\beta \left[g_{\mu\nu} \right] = \beta \left[\Phi \right] = 0$ where $g_{\mu\nu} = \{-1, g_{ij}\}$

• Compare to $\frac{\partial g_{ij}}{\partial t} = -\beta [g_{ij}]$

Simplest situation: 3D space \equiv round sphere (isotropic Bianchi IX $\rightarrow \gamma_1 = \gamma_2 = \gamma_3 = L^2 \exp 2\sigma$) with torsion $H = 2/L\omega_{S_1^3}$

► 3D:
$$\frac{d\sigma}{d\log\mu^{-1}} = -\frac{2e^{-2\sigma}}{L^2} (1 - e^{-4\sigma}) \Rightarrow \text{IR fixed-point } \sigma = 0$$

(radius *L*)

- 4D: define $Q \equiv -\dot{\Phi} + \frac{3}{2}\dot{\sigma}$
 - full equations:

$$\begin{split} \dot{Q} &= -\frac{3}{2}\dot{\sigma}^2 \\ \ddot{\sigma} &+ 2Q\dot{\sigma} = -\frac{2\mathrm{e}^{-2\sigma}}{L^2} \left(1 - \mathrm{e}^{-4\sigma}\right) \end{split}$$

- classical motion of a particle at coordinate σ in a potential

$$V = V_0 - \frac{e^{-2\sigma}}{L^2} \left(1 - \frac{e^{-4\sigma}}{3} \right)$$

with a friction force $-2Q\dot{\sigma}$ due to the dilaton motion (like the inflaton [lectures by Lazarides and Sasaki])

4D dynamics: steady-state universe \equiv linear dilaton solution

 $\sigma = 0, \ \Phi = \Phi_0 - \bar{Q}t$

space \equiv IR fixed point of the 3D sigma model

4D dynamics: full regime \rightarrow friction \rightarrow the kinetic term $\ddot{\sigma}$ becomes subdominant \rightarrow first-order 3D equation with

 $t = -2Q \log \mu$

4D genuine time evolution \simeq RG i.e. Ricci flow in 3D

Evolution towards the RG regime

- At early times the 4D genuine time evolution is
 - oscillatory if $Q_0 < \sqrt{8}/L$ (large initial curvature)
 - monotonic if $Q_0 > \sqrt{8}/L$ (small initial curvature)
- ► At late times (3D Ricci-flow regime) → monotonic

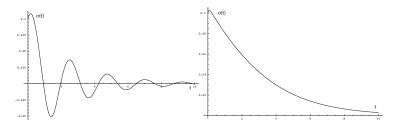


Figure: Typical behaviour for $\sigma(t)$ in the non-linear system

General result [Schmidhuber, Tseytlin, 1994]

- RG fixed points in $D \equiv$ steady-state universes in D+1
- ► In the vicinity of the fixed points: evolution ≡ generalised Ricci flow
- Far or between fixed points: dynamics reminiscent of Zamolodchikov's c-theorem
- Origin: dilaton motion (necessary for CFT) \rightarrow friction

Limitations

- Usual approximations
- Case of marginal lines in 3D sigma models

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Framework

Gravitational instantons: positive action conjecture, tools to handle non-perturbative transitions in quantum gravity, ...^[70s]

- Non-singular solutions of Euclidean Einstein's equations
- Hard task in general easier with some specific ansatz
 - \blacktriangleright 4D geometry foliated with 3D leaves: $\mathcal{M}_4 = \mathcal{T} \times \mathcal{M}_3$
 - ▶ homegeneous spatial sections $M_3 \exists G_{hom}$ of Bianchi type
 - self-duality in vacuum \rightarrow first-order equations consistent if G_{hom} algebra is unimodular

Connection with Ricci flows of homogeneous 3D spaces [Cvetič, Gibbons, Lii, Pope, 2001; Bourliot, Estes, Petropoulos, Spindel 2009]

- ► The gravitational instantons on M₄ are in one-to-one correspondence with the flat SU(2) connections on M₃
- ► In all unimodular cases Euclidean-time evolution of the leaves is a geometric flow on M₃ driven by Ricci plus a flat SU(2) gauge connection

Again the dynamics of Ricci flows seems to capture – exactly – higher-dimensional gravitational set-ups

Solving 4D Euclidean Einstein equations

Usual set-up: orthonormal frame $ds^2 = \delta_{ab}\theta^a \theta^b$ invariant under local SO(4) transformations, connection one-form $\omega^a{}_b$, curvature two-form $\mathcal{R}^a{}_b$

- Riemann tensor: $\mathcal{R}^{a}_{\ b} = \mathsf{d}\omega^{a}_{\ b} + \omega^{a}_{\ c} \wedge \omega^{c}_{\ b}$
- (Anti-)self-dual-curvature solutions:

•
$$\mathcal{R}^{a}_{\ b} = \pm \tilde{\mathcal{R}}^{a}_{\ b}$$
 & Cyclic id. \Rightarrow vacuum eqs.

• $\mathcal{R}^{a}_{\ b} = \pm \tilde{\mathcal{R}}^{a}_{\ b} \Leftrightarrow \omega^{a}_{\ b} = \pm \tilde{\omega}^{a}_{\ b}$ up to local SO(4) rotations $(\tilde{\omega}^{a}_{\ b} = \frac{1}{2} \epsilon^{a}_{\ bc}{}^{d} \omega^{c}_{\ d}) \Rightarrow$ several branches of 1st order equations

Curvature and connection \in **6** *(antisymmetric) of* SO(4) – *reducible as* (**3**, **3**) *under the decomposition* $SO(4) \cong SU(2)_{sd} \otimes SU(2)_{asd}$

In the geometries $\mathcal{M}_4 = \mathcal{T} \times \mathcal{M}_3$ the decomposition is simple

Curvature two-form

$$S_i = \frac{1}{2} \left(\mathcal{R}_{0i} + \frac{1}{2} \epsilon_{ijk} \mathcal{R}^{jk} \right)$$

$$A_i = \frac{1}{2} \left(\mathcal{R}_{0i} - \frac{1}{2} \epsilon_{ijk} \mathcal{R}^{jk} \right)$$

Connection one-form

$$\Sigma_i = \frac{1}{2} \left(\omega_{0i} + \frac{1}{2} \epsilon_{ijk} \omega^{jk} \right)$$

$$A_i = \frac{1}{2} \left(\omega_{0i} - \frac{1}{2} \epsilon_{ijk} \omega^{jk} \right)$$

Relations

•
$$S_i = d\Sigma_i - \epsilon_{ijk}\Sigma^j \wedge \Sigma^k$$

• $A_i = dA_i + \epsilon_{ijk}A^j \wedge A^k$

 $\{S_i, \Sigma_i\}$ vectors of $SU(2)_{sd}$ and singlets of $SU(2)_{asd}$ and vice-versa for $\{A_i, A_i\}$

Self-dual gravitational instantons

 $\mathcal{A}_i = 0 \leftrightarrow \mathsf{flat} \ SU(2)$ connections on \mathcal{M}_3

Further assumption \mathcal{M}_3 homogeneous of Bianchi type locally G_{hom} flat SU(2) connections on $\mathcal{M}_3 \leftrightarrow$ homomorphisms $G_{hom} \rightarrow SU(2)$ *Example: Bianchi IX* $G_{hom} \equiv SU(2)$

- 1. Trivial homomorphism: $G_{\text{hom}} \rightarrow \text{identity of } SU(2), A_i = 0 \ \forall i$ Eguchi–Hanson branch e.g.
- 2. Isomorphism (rank-3): $A_i = \sigma^i/2 \quad \forall i \text{ Taub-NUT branch e.g.}$

Note: non-singular real self-dual gravitational instantons exist only for unimodular – class A – Bianchi groups and for Bianchi III

The view from the leaf: geometric flows

4D self-dual gravitational instantons foliated in time with Bianchi-A homogeneous spaces

$$\mathsf{d}s^2 = \mathsf{d}t^2 + \sum_i \left(\gamma_i \sigma^i\right)^2$$

- γ_i(T): 1st order differential equations branches: choice of flat SU(2) anti-self-dual Levi−Civita connection A_i (A_i = 0)
- Example: Bianchi IX

$$\begin{cases} 2\frac{\dot{\gamma}_1}{\gamma_1} = (\gamma_2 - \gamma_3)^2 - \gamma_1^2 + 2\tilde{\lambda}\gamma_2\gamma_3 \text{ and cyclic} \\ \\ A_i = (1 - \tilde{\lambda})\frac{\sigma^i}{2} \end{cases}$$

- $\tilde{\lambda} = 0$: Taub-NUT Darboux-Halphen system
- $\tilde{\lambda} = 1$: Eguchi–Hanson Lagrange system (Euler top)

The evolution of the leaves in Euclidean time is covariant – resembles a flow due to the privileged rôle of time in $\mathcal{M}_4 = \mathcal{T} \times \mathcal{M}_3$ – what kind of flow?

Consider a pure 3D set-up: same Bianchi-class M_{3S} – parameter t

- Metric: $d\tilde{s}^2 = \sum_i \gamma_i(t) (\sigma^i)^2$
- SU(2) gauge connection:

$$\tilde{A} = \tilde{A}_i \sqrt{\gamma_i} \sigma^i = \tilde{A}_i^j T_j \sqrt{\gamma_i} \sigma^i$$

$$\tilde{F} = d\tilde{A} + [\tilde{A}, \tilde{A}]$$

Define a generalised Ricci flow

$$egin{aligned} &\left(rac{\partial ilde{g}_{ij}}{\partial t}=- ilde{R}_{ij}-rac{1}{2} ext{tr}\left(ilde{A}_{i} ilde{A}_{j}
ight)\ &\left(rac{\partial ilde{A}_{i}}{\partial t}=0, \quad ilde{F}=0 \end{aligned}$$

- Dynamics in parametric time = coordinate-time evolution in the corresponding gravitational instanton – for each branch defined by a flat connection
- Example: Bianchi IX $\rightarrow \tilde{A}_i = -\frac{\tilde{\lambda} T_i \sigma^i}{\sqrt{\gamma_i}}$
 - $\tilde{\lambda} = 0$: pure Ricci flow (Darboux–Halphen) Taub–NUT
 - $\tilde{\lambda} = 1$: Ricci plus gauge (Lagrange) Eguchi–Hanson

Comments

The dynamics of self-dual gravitational instantons in vacuum is a geometric flow driven by Ricci plus flat SU(2) gauge connection

- ► Ã is a background SU(2) gauge field inherited from the anti-self-dual part of the 4D Levi-Civita connection
- ▶ The geometric flow *is not* gauge invariant not supposed to be
- The gauge field
 - does not flow $(\dot{\tilde{A}} = 0)$
 - its strength is set to $\tilde{\tilde{F}} = 0$

Towards a Ricci plus Yang–Mills flow with $\tilde{F} \neq 0 \& \dot{\tilde{A}} \neq 0$

- We must relax the self-duality condition: $A_i \neq 0$
- A similar condition is needed to provide 1st order equations
 - introduce a 4D cosmological constant Λ_c
 - impose self-duality on the Weyl tensor to solve $R_{ab} = \Lambda_c g_{ab}$

In the presence of cosmological constant 4D gravitational instantons with homogeneous spatial sections are equivalent to geometric flows of 3D leaves with a background SU(2) gauge field \tilde{A} s.t. $\tilde{A} \otimes \tilde{F} \neq 0$ Example in Bianchi IX: the Fubini–Study solution (metric on $\mathbb{C}P_2$)

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3D geometric flows in homogeneous spaces versus self-duality in Einstein dynamics in 4D space–times with time foliation and homogeneous spatial sections

- ▶ Role of 4D: $SO(4) \cong SU(2) \times SU(2) \Rightarrow$ reduction is sd \oplus asd
- ▶ Role of the 3D homogeneity: $G_{\text{hom}} \rightarrow SU(2) \Rightarrow$ gauge choice
- Role of the self-duality: effectively reduces the system from 4D to 3D off-shell though: "g_{ij} runs"

Hard to generalise in arbitrary dimension (except perhaps is some hyper-Kähler or quaternionic spaces)

Horava–Lifshitz gravity: similar properties \rightarrow *similar conclusions* [Horava 2008-09]

- Similar spaces: foliation $\mathcal{M}_4 = \mathcal{T} \times \mathcal{M}_3$
- Major difference: explicit breaking of the diffeomorphism invariance – previously this breaking was spontaneous
- Similar constraint: detailed balance previously self-duality
- ▶ Similar effect in 4D \rightarrow 3D: the dynamics is governed by combined Ricci and Cotton flows
- Consequences:
 - valid in $D + 1 \rightarrow D$ dimensions with other flows
 - important issue of degrees of freedom and the appearance of Einstein gravity in the IR
 - richer landscape of fixed points not always isotropic [Bakas, Bourliot, Lüst, Petropoulos 2009]

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Ricci and more general geometric flows appear in a wide palette of physical problems in diverse dimensions

Here: selection of topics related to gravity and time evolution with some fundamental questions: (i) Relation time evolution/RG flow/Liouville field? (ii) Relation energy evolution/holography?