

# Strongly-coupled Anisotropic Gauge Theories And Holography

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Based on works with:  
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in progress.

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# Outline

- 1 Introduction
- 2 The Theory
- 3 Stable and Physical Theories
- 4 Phase Transitions
- 5 Probing the Theory
- 6 Conclusions

# Briefly on AdS/CFT

- Gauge/Gravity duality: A way to map and answer **quantum questions** to **gravity geometric** questions.
- The initial AdS/CFT correspondence:  $\mathcal{N} = 4$  sYM on flat space  $\Leftrightarrow AdS_5 \times S^5$ , is the **harmonic oscillator** of the gauge/gravity dualities.
- The theory is simple: **Conformal, Maximally Supersymmetric, No Temperature...**
- Since the discovery of the initial correspondence, there is an extensive research aiming to construct more realistic gauge/gravity dualities.

**Gauge/Gravity Dualities** with: **Less/No Supersymmetry; Broken conformal symmetry, confinement; fundamental matter**(probe and backreacting Dq branes); etc.

✓ We study **Anisotropic** theories in Gauge/Gravity correspondence.

# Why? Attempts for Realizations in Nature

The existence of **strongly coupled anisotropic systems**.

- The expansion of the plasma along the longitudinal beam axis at the earliest times after the collision results to **momentum anisotropic plasmas**.
- Strong **Magnetic Fields** in strongly coupled theories.
- New interesting phenomena in presence on such fiels, i.e. **inverse magnetic catalysis**.  
*eg: (Bali, Bruckmann, Endrodi, Fodor, Katz, Krieg et al. 2011)*
- Anisotropic low dimensional **materials** in condensed matter.

# Why? More:

- Weakly coupled vs strongly coupled anisotropic theories.  
(Dumitru, Strickland, Romatschke, Baier,... 2008,...)
- Consistent top-down models. Properties of the supergravity solutions, that are dual to the anisotropic theories.
- Black hole solutions that are AdS in UV flowing to Lifshitz-like in IR :
  - ★ Why there is a fixed scaling parameter  $z$  for such solutions?  
(Azeyanagi, Li, Takayanagi, 2009)
  - ★ Other systems that have fixed scaling IR solution (e.g. in Heavy quark density). Why?  
(Kumar 2012; Faedo, Kundu, Mateos, Tarrío 2014)
  - ★ New flows to Hyperscaling violation IR backgrounds?

# Why? Even More:

- Ⓢtriking Features! Several **Universality Relations** predicted for the isotropic theories are **violated**!
  - ★ **Shear viscosity over entropy density ratio** takes **parametrically** low values  $\frac{\eta}{s} < \frac{1}{4\pi}$ !  
 (Rebhan, Steineder 2011; Jain, Samanta, Trivedy 2015; D.G., Gursoy, Pedraza, 2017;...)
  - ★ **Langevin coefficients** inequality for heavy quark motion in the **anisotropic** plasma gets inverted  $\kappa_L > \kappa_T$  .  
 (Gursoy, Kiritsis, Mazzanti, Nitti 2010; D.G, Soltanpanahi, 2013a, 2013b )
  - ★ ...
  - ★ Implications to QGP **hydrodynamic simulations**.

# The Introduction of the Theory in One Page:

- Strongly coupled **anisotropic** theory.
- How the theory looks like and how to obtain it?
  - ✓ 4d  $SU(N)$  gauge theory in the large  $N_c$ -limit.
  - ✓ Its dynamics are affected by a **scalar operator**  $\mathcal{O} \sim \text{Tr} F^2$ .
  - ✓ Anisotropy is introduced by **another operator**  $\tilde{\mathcal{O}} \sim \theta(x_3) \text{Tr} F \wedge F$  with a space dependent coupling.
  - ✓ On the gravity dual side we have a "backreacting" scalar field depending on spatial directions, the **axion**; and a non-trivial **dilaton**.
- Eventually the gravity dual theory is an **Einstein-Axion-Dilaton theory** in 5 dimensions with a non-trivial potential.
  - ✓ Solutions are RG flows:  
AdS in UV  $\Rightarrow$  **Anisotropic (Hyperscaling Lifshitz-like)** in IR.
- **Formally interesting** theories. And a solid ground to study **strongly coupled phenomena** in presence of **anisotropy**.

## Parts of the Theory Timeline-Related bibliography:

### Non-Confining Anisotropic Theories:

*(Azeyanagi, Li, Takayanagi, 2009; Mateos, Trancanelli, 2011; Jain, Kundu, Sen, Sinha, Trivedi, 2015;...)*

### Confining Anisotropic Theories:

*(D.G., Gursoy, Pedraza, 2017 )*

Similar ideas in different context. For example:

*(Gaiotto, Witten 2008; Chu, Ho, 2006; Choi, Fernandez, Sugimoto 2017;...)*



# How is Anisotropy introduced? A Pictorial Representation:

- For the Lifshitz-like IIB Supergravity solutions

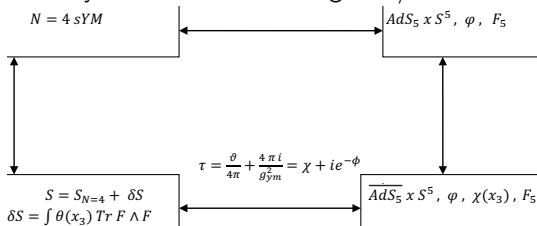
$$ds^2 = u^{2z}(dx_0^2 + dx_i^2) + u^2 dx_3^2 + \frac{du^2}{u^2} + ds_{S^5}^2.$$

Introduction of additional branes:

(Azeyanagi, Li, Takayanagi, 2009)

	$x_0$	$x_1$	$x_2$	$x_3$	$u$	$S^5$
D3	X	X	X	X		
D7	X	X	X			X

- Which equivalently leads to the following AdS/CFT deformation.



- $dC_8 \sim \star d\chi$  with the non-zero component  $C_{x_0 x_1 x_2 S^5}$ .

- Possible to compactify  $x_3$  to get dual to the pure Chern-Simons gauge theory.

# The Anisotropic Theory

The generalized **Einstein-Axion-Dilaton action** with a **potential** for the dilaton and an **arbitrary coupling** between the axion and the dilaton:

$$S = \frac{1}{2\kappa^2} \int d^5x \sqrt{-g} \left[ R - \frac{1}{2}(\partial\phi)^2 + V(\phi) - \frac{1}{2}Z(\phi)(\partial\chi)^2 \right].$$

The eoms read

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = \frac{1}{2}\partial_\mu\phi\partial_\nu\phi + \frac{1}{2}Z(\phi)\partial_\mu\chi\partial_\nu\chi - \frac{1}{4}g_{\mu\nu}(\partial\phi)^2 - \frac{1}{4}g_{\mu\nu}Z(\phi)(\partial\chi)^2 + \frac{1}{2}g_{\mu\nu}V(\phi),$$

$$\frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu\phi) = \frac{1}{2}\partial_\phi Z(\phi)(\partial\chi)^2 - V'(\phi),$$

$$\frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu\chi) = 0.$$

Where

$$V(\phi) = 12 \cosh(\sigma\phi) + \left( \frac{m(\Delta)^2}{2} - 6\sigma^2 \right) \phi^2, \quad Z(\phi) = e^{2\gamma\phi},$$

(Gursoy, Kiritsis, Nitti, 2007; (Gubser, Nellore), Pufu, Rocha 2008a,b)

**Remark:** For  $\sigma = 0, \gamma = 1, m(\Delta) = 0$  the action and the solution of eoms, are reduced of IIB supergravity.

# Solutions of the Generalized Einstein-Axion-Dilaton Action

- The background solution

$$ds^2 = \frac{1}{u^2} \left( -\mathcal{F}(u)\mathcal{B}(u) dt^2 + dx_1^2 + dx_2^2 + \mathcal{H}(u)dx_3^2 + \frac{du^2}{\mathcal{F}(u)} \right),$$

$$\chi = \alpha x_3, \quad \phi = \phi(u),$$

- Solutions: e.g. For  $\Delta = 4$ :  
AdS in UV flowing to Hyperscaling Lifshitz-like violation geometries in IR :

$$ds^2 = u^{-\frac{2\theta}{3}} \left( -u^{2z} (f(u) dt^2 + dx_{1,2}^2) + \tilde{\alpha} u^2 dx_3^2 + \frac{du^2}{f(u)u^2} \right),$$

# Axion-Dilaton Coupling and Potential, rule the Scaling Coefficients

- The values of  $(\theta, z)$  dependence on  $(\gamma, \sigma)$

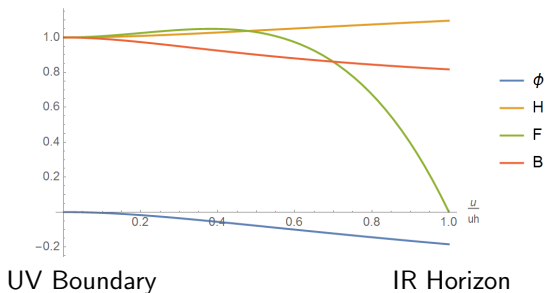
$$z = \frac{4\gamma^2 - 3\sigma^2 + 2}{2\gamma(2\gamma - 3\sigma)} , \quad \theta = \frac{3\sigma}{2\gamma} .$$

- **Special case:**  $(\sigma = 0, \gamma = 1)$  **supergravity truncated action** with a **single** solution  $(\theta = 0, z = 3/2)$ . *(Mateos, Trancanelli, 2011)*
- The scaling factors  $z$  and  $\theta$  are determined by the constants in the **Axion-Dilaton Coupling** and the **Potential**. This is the reason that in the particular setup the supergravity solutions have them fixed.

# Solution : The Full Flow

- Fixing  $(\gamma, \sigma)$  and  $\alpha$  and  $u_h$  we get the **metric flow** from boundary to horizon:

$$ds^2 = \frac{e^{-\frac{1}{2}\phi(u)}}{u^2} \left( -\mathcal{FB} dt^2 + dx_1^2 + dx_2^2 + \mathcal{H} dx_3^2 + \frac{du^2}{\mathcal{F}} \right),$$



# An exact solution

The potential and the axion-dilaton coupling

$$V(\phi) = 6e^{\sigma\phi}, \quad Z(\phi) = e^{2\gamma\phi}.$$

A Lifshitz-like anisotropic hyperscaling violation background which may accommodate a black hole

$$ds_s^2 = \alpha^2 C_R e^{\frac{\phi(u)}{2}} u^{-\frac{2\theta}{3z}} \left( -u^2 (f(u) dt^2 + dx_i^2) + C_Z u^{\frac{2}{z}} dx_3^2 + \frac{du^2}{f(u)\alpha^2 u^2} \right),$$

where

$$f(u) = 1 - \left( \frac{u_h}{u} \right)^{3+(1-\theta)/z}, \quad e^{\frac{\phi(u)}{2}} = u^{\frac{\sqrt{\theta^2 + 3z(1-\theta)} - 3}{\sqrt{6z}}},$$

$$C_R = \frac{(3z - \theta)(1 + 3z - \theta)}{6z^2}, \quad C_Z = \frac{z^2}{2(z - 1)1 + 3z - \theta},$$

$$z = \frac{4\gamma^2 - 3\sigma^2 + 2}{2\gamma(2\gamma - 3\sigma)}, \quad \theta = \frac{3\sigma}{2\gamma}.$$

We have obtained the theories, are they **physical**  
and **stable**?



✓ Energy Conditions Analysis

$$T_{\mu\nu} N^\mu N^\nu \geq 0, \quad N^\mu N_\mu = 0,$$

✓ Local Thermodynamical Stability Analysis

$$c_\alpha = T \left( \frac{\partial S}{\partial T} \right)_\alpha \geq 0, \quad \Phi' = \left( \frac{\partial \Phi}{\partial \alpha} \right)_T \geq 0$$



YES!

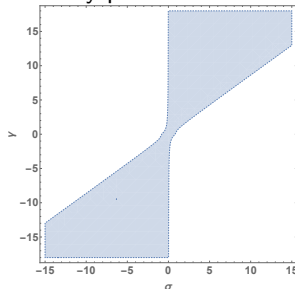
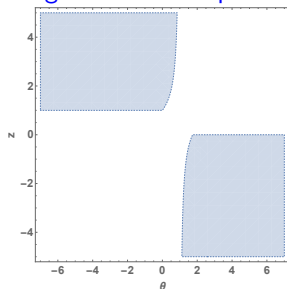
Three conditions that constrain  $(z, \theta)$  and as a result  $(\gamma, \sigma)$ .

$$(z - 1)(1 - \theta + 3z) \geq 0, \quad z = \frac{2 + 4\gamma^2 - 3\sigma^2}{2\gamma(2\gamma - 3\sigma)},$$

$$\theta^2 - 3 + 3z(1 - \theta) \geq 0, \quad \theta = \frac{3\sigma}{2\gamma},$$

$$1 - \theta + 2z \geq 0.$$

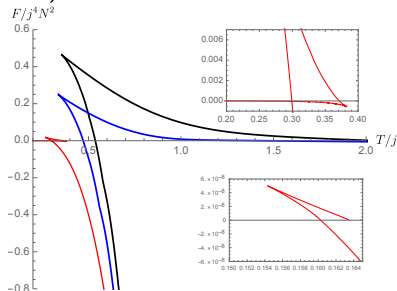
The blue region is the acceptable for the theory parameters.





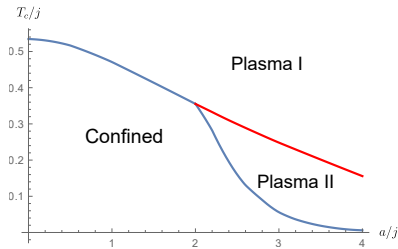
# Confinement/Deconfinement Phase transitions

- The free energy of the theories vs the temperature  $T$  for different anisotropy ( $\alpha/j=0,1,3$ ):



- Horizontal Axis: **Confining Phase**.
- Upper Branch: **Black hole A: Deconfining Plasma Phase**.
- Lower Branch: **Black hole B: Deconfining Plasma Phase**.
- $\alpha/j \simeq 2$ : A critical value above which a richer structure in the phase diagram exist.

- The **Critical Temperature** of the theories vs the **anisotropy** gives:



- The  $T_c$  is **reduced** in presence of anisotropies of the theory.

# The Proposal

- The  $T_c(\alpha)$  decrease with  $\alpha$ , resembling the phenomenon of **inverse magnetic catalysis** where the **confinement-deconfinement** temperature decreases with the magnetic field  $B$  (where an anisotropy is introduced as in our plasma).
- **No charged fermionic degrees** of freedom in our case; our plasma is neutral.
- Our findings suggest that the **anisotropy by itself** could instead be the cause of lower  $T_c$  in presence of **anisotropies**.

# $\eta/s$ for our theories

- Shear Viscosity over Entropy Density

$$\eta_{ij,kl} = - \lim_{\omega \rightarrow 0} \frac{1}{\omega} \text{Im} \int dt dx e^{i\omega t} \langle T_{ij}(t, x), T_{kl}(0, 0) \rangle$$

$$s = \frac{2\pi}{\kappa^2} A .$$

The **two-point function** is obtained by calculating the **response** to turning on suitable metric perturbations in the bulk.

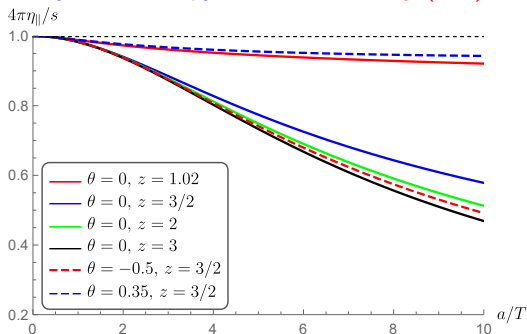
- The relevant part of the perturbed action is mapped to a **Maxwell system with a mass term**.

$$S = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} \left( -\frac{1}{4g_{\text{eff}}^2} F^2 - \frac{1}{4} m^2(u) A^2 \right) ,$$

where

$$m^2(u) = Z(\phi + \frac{1}{4} \log g_{33}) \alpha^2 , \quad \frac{1}{g_{\text{eff}}^2} = g_{33}^{3/2}(u) , \quad A_\mu = \frac{\delta g_{\mu 3}}{g_{33}}$$

- The shear viscosity over entropy ratio for arbitrary  $(z, \theta)$ .

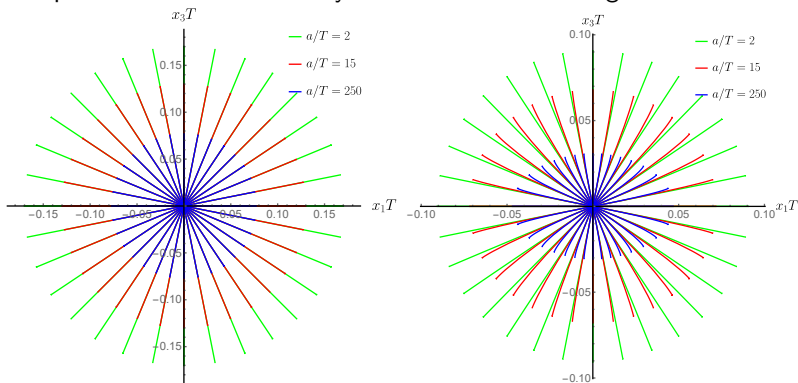


- The ratio depends on the temperature at  $a/T \gg 1$  as

$$4\pi \frac{\eta_{||}}{s} = \frac{g_{11}}{g_{33}} \sim \left( \frac{T}{\tilde{\alpha}|1 + 3z - \theta|} \right)^{2 - \frac{2}{z}}.$$

- The range of the temperature power is  $[0, \infty)$ .

- The quark distributions for baryons in theories with magnetic fields:



- Baryon on the transverse plane and Baryon on the plane that the field lies.
- System of **fundamental strings with a vertex Dp-brane**, in an anisotropic gravity theory.
- The baryon **dissociates at stages**, depending on the proximity angle to the **magnetic field direction**.

# Conclusions

- ✓ We have obtained and studied a) **Confining Anisotropic** theories. b) **Hyperscaling Lifshitz-like Anisotropic black holes** with **arbitrary scalings**. (**1st construction in the literature**)
- ✓ The theories are **physical and stable** for a wide range of **parameters** of the theory.
- ✓ The **Confinement/Deconfinement** phase transitions occur at **lower critical Temperature** as the anisotropy is **increased**!
- ✓ The **anisotropy by itself** could instead be the cause of the **inverse magnetic catalysis**.
  - The **shear viscosity over entropy density** ratio, takes values parametrically **lower** than  $1/4\pi$ , and depends on the **Temperature** as  $(T/\alpha)^{2-2/z}$ .
  - The **diffusion (butterfly velocity) of chaos** occurs **faster** than isotropic systems.
  - **Baryons** dissociate **at stages** in theories with strong fields. (D.G. 2018)

# Thank you!



# Reminding Slide:

- The anisotropic **hyperscaling violation** metric

$$ds^2 = r^{-\frac{2\theta}{dz}} \left( -r^2(dt^2 + dy_i^2) + r^{\frac{2}{z}} dx_i^2 + \frac{dr^2}{r^2} \right).$$

which exhibits a **critical exponent  $z$**  and a **hyperscaling violation exponent  $\theta$** .

- The metric is not scale invariant

$$t \rightarrow \lambda^z t, \quad y \rightarrow \lambda^z y, \quad x \rightarrow \lambda x, \quad r \rightarrow \frac{r}{\lambda^z}, \quad ds \rightarrow \lambda^{\frac{\theta}{d}} ds.$$

# Special case: IIB Supergravity

## Remark:

The **ten dimensional action** gives our generalized model, when the internal space is an  **$S^5$**  supported by fluxes and  $\sigma = 0, \gamma = 1, \Delta = 4$ :

$$S = \frac{1}{2\kappa_{10}^2} \int d^{10}x \sqrt{-g} \left[ R + 4\partial_M \phi \partial^M \phi - e^{2\phi} \left( \frac{1}{2} F_1^2 + \frac{1}{4 \cdot 5!} F_5^2 \right) \right], \quad F_1 := d\chi.$$

where  $M = 0, \dots, 9$  and  $F_1$  is the axion field-strength. The equations of motion for the background are:

$$R + 4g^{MN} (\nabla_M \nabla_N \phi - \partial_M \phi \partial_N \phi) = 0,$$

$$R_{MN} + 2\nabla_M \nabla_N \phi + \frac{1}{4} g_{MN} e^{2\phi} \partial_P \chi \partial^P \chi - \frac{1}{2} e^{2\phi} \left( F_M F_N + \frac{1}{48} F_{MABCD} F_N^{ABCD} \right) = 0.$$

plus the **Bianchi identities** and **self duality** constraints. The **axion field** equation is satisfied trivially for **linear axion**.

# Null Energy Condition

- The averaged radial acceleration between two null geodesics is

$$A_r = -4\pi T_{\mu\nu} N^\mu N^\nu ,$$

if it is negative the null geodesics observe a **non-repulsive gravity** on nearby particles along them.

- This imposes the **Null Energy Condition**

$$T_{\mu\nu} N^\mu N^\nu \geq 0 , \quad N^\mu N_\mu = 0 ,$$

leading to the following constraints:

- For the **Lifshitz-like** space  $z \geq 1$ .
- For the **Hyperscaling violation anisotropic metric** in 3+1-dim spacetime and anisotropic in 1-dim reads

$$(z-1)(1-\theta+3z) \geq 0 ,$$

$$\theta^2 - 3 + 3z(1-\theta) \geq 0 .$$

**Additional** conditions from **thermodynamics**?

# Local Thermodynamic Stability

- The **necessary and sufficient conditions** for **local thermodynamical stability** in the **canonical ensemble** are

$$c_\alpha = T \left( \frac{\partial S}{\partial T} \right)_\alpha \geq 0, \quad \Phi' = \left( \frac{\partial \Phi}{\partial \alpha} \right)_T \geq 0$$

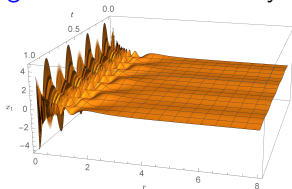
$c_\alpha$  is the **specific heat**: increase of the temperature leads to increase of the entropy.

$\Phi'$  is **derivative of the potential**: the system is stable under infinitesimal charge fluctuations.

- In the **GCE** these conditions should be equivalent of having **no positive eigenvalues** of the **Hessian matrix** of the entropy with respect to the thermodynamic variables. *(Gubser, Mitra 2001)*
- In the IR the **positivity** of **the specific heat** imposes

$$c_\alpha = 1 - \theta + 2z \geq 0$$

- In the linear response theory the response function  $\chi(\omega)$  of the quark to the an external force, is proportional to the **two point correlator of a string fluctuations** divided by the applied force.



- The string fluctuations along  $x_1$  close to the boundary are found by the **monodromy patching method**

$$\delta x_{1\omega}(r) = c_1 \left( 1 + i\omega c_0 g_{11}(r_h) + \frac{i\omega g_{11}(r_h)}{2\kappa\nu} r^{-2\kappa\nu} \right).$$

- The diffusion coefficient  $D = T \lim_{\omega \rightarrow 0} (-i \omega \chi(\omega))$

$$D \sim T^{2(1-\nu_i)}$$

where  $\nu_i$  is determined by the asymptotics of the metric element along the fluctuation direction.

- Fluctuation-Dissipation theorem holds along each direction;** **The noise is white;** **Self energy an thermal mass of the particle** depend on the properties and direction of the system...