

New D-term uplift

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Supersymmetry

- \rightarrow Motivated from BSM physics.
- \rightarrow Motivated from UV complete theories.

BUT

→ SUSY is (spontaneously) broken on de Sitter vacuum.

How is N=1 SUSY broken?

 \rightarrow F-term (chiral multiplet).

 \rightarrow D-term (U(1) gauge multiplet).

 $\,\,\,\,\,\,\,\,\,\,\,\,\,$ Other methods (complex linear, etc).

Fayet-Iliopoulos D-term

→ A gauge multiplet contains the component fields

$$V_{WZ} = -\theta \sigma^m \overline{\theta} A_m - i \overline{\theta}^2 \theta^\alpha \lambda_\alpha + i \theta^2 \overline{\theta}_{\dot{\alpha}} \overline{\lambda}^{\dot{\alpha}} + \frac{1}{2} \theta^2 \overline{\theta}^2 D.$$

→ The simplest model with a Fayet–Iliopoulos term is

$$\begin{split} \mathcal{L} &= \frac{1}{4} \left(\int d^2 \theta \, W^2(V) + c.c. \right) - \frac{2\xi}{4} \int d^4 \theta \, V \\ &= -\frac{1}{4} F^{mn} F_{mn} - i \lambda \sigma^m \partial_m \overline{\lambda} + \frac{1}{2} \mathsf{D}^2 - \xi \, \mathsf{D} \, . \end{split}$$

Fayet, Iliopoulos '74

Supersymmetry is broken spontaneously

- \rightarrow The auxiliary field gets a vev: $\langle D \rangle = \xi$.
- \rightarrow The goldstino is: $\delta \lambda_{\beta} = -i\xi \, \epsilon_{\beta} + \cdots$
- \rightarrow The vacuum energy is: $\langle V \rangle = \frac{1}{2} \xi^2$.

How do we couple to supergravity?

Plan

 \rightarrow Freedman model and R-symmetry gauging.

→ New model without R-symmetry gauging.

 \rightarrow Properties of the new D-term.

The Freedman model

The Noether method

▶ The gauge field of SUSY is the gravitino ψ_m^{α} , with supersymmetry transformation

$$\delta\psi_{\mathbf{m}\alpha} = -\mathbf{2}\,\partial_{\mathbf{m}}\epsilon_{\alpha} + \cdots$$

▶ We start from $-e\xi D$ and perform the Noether procedure

$$\partial_m \epsilon_{\alpha}(\mathbf{x}) \to \psi_{m\alpha}$$
.

At some point we have to cancel

$$-ie\xi \, \epsilon^{klmn} \left(\overline{\psi}_k \overline{\sigma}_l \epsilon - \overline{\epsilon} \, \overline{\sigma}_l \psi_k \right) \partial_n A_m \,.$$

Therefore we have to add

$$\mathcal{L}_{\psi\psi\mathsf{A}} = rac{i}{2}e\,\xi\,\,\epsilon^{klmn}\overline{\psi}_{k}\overline{\sigma}_{l}\psi_{n}\mathsf{A}_{m}\,.$$

- ▶ This means that the gravitino is charged under the U(1).
 - \rightarrow The FI term in supergravity requires R-symmetry gauging by A_m .
 - → This has impact on model building. (Constraints on the gravitino mass.) *Barbieri, Ferrara, Nanopoulos, Stelle '82, Villadoro, Zwirner '05*

▶ Once we complete the procedure we have (on-shell)

$$\begin{split} e^{-1}\mathcal{L}\Big|_{\lambda=0} &= -\frac{1}{2}R + \frac{1}{2}\epsilon^{klmn}\left(\overline{\psi}_{k}\overline{\sigma}_{l}\mathcal{D}_{m}\psi_{n} - \psi_{k}\sigma_{l}\mathcal{D}_{m}\overline{\psi}_{n}\right) \\ &- \frac{1}{4}F_{mn}F^{mn} + \frac{i}{2}e\xi\epsilon^{klmn}\overline{\psi}_{k}\overline{\sigma}_{l}\psi_{n}A_{m} - \frac{1}{2}\xi^{2}\,. \end{split}$$

In superspace this would read

$$\mathcal{L} = -3\int d^4 \theta \ E \, e^{rac{2}{3}\xi \, V} + rac{1}{4} \left(\int d^2 \Theta \, 2 \mathcal{E} \ W^2(V) + c.c.
ight) \, . \label{eq:lambda}$$

A new supergravity embedding Cribiori, Tournoy, FF, Van Proeyen '17 • We observe that λ_{α}/D has a very nice property

$$\delta\left(i\frac{\lambda_{\alpha}}{\mathsf{D}}\right) = \epsilon_{\alpha} + \cdots$$

To cancel the term

$$-ie\xi \, \epsilon^{klmn} \left(\overline{\psi}_k \overline{\sigma}_l \epsilon - \overline{\epsilon} \, \overline{\sigma}_l \psi_k \right) \partial_n A_m$$

during the Noether procedure, we introduce instead the term

$$- \mathbf{e} \xi \, \epsilon^{klmn} \left(\overline{\psi}_{\mathbf{k}} \overline{\sigma}_{\mathbf{l}} \frac{\lambda}{\mathbf{D}} - \frac{\overline{\lambda}}{\mathbf{D}} \, \overline{\sigma}_{\mathbf{l}} \psi_{\mathbf{k}} \right) \partial_{\mathbf{n}} \mathbf{A}_{\mathbf{m}} \, .$$

Supersymmetry has to be broken and by assumption: $\langle D \rangle \neq 0$.

A new FI D-term

▶ In superspace the full U(1) sector is

$$\mathcal{L}_{NEW} = \frac{1}{4} \left(\int d^2 \Theta \, 2\mathcal{E} \, W^2(V) + c.c. \right)$$

$$+8 \, \xi \int d^4 \theta \, E \, \frac{W^2 \overline{W}^2}{\mathcal{D}^2 W^2 \overline{\mathcal{D}}^2 \overline{W}^2} \, \mathcal{D}^{\alpha} W_{\alpha} \, .$$

And in components

$$e^{-1}\mathcal{L}_{NEW} = -rac{1}{4}F_{mn}F^{mn} - i\overline{\lambda}\overline{\sigma}^{m}\mathcal{D}_{m}\lambda + rac{1}{2}\mathsf{D}^{2} \ - \xi\mathsf{D} + \xi\,\mathcal{O}ig(\lambda,\overline{\lambda}ig) \; .$$

▶ The gravitino is not charged under the U(1).

The new term coupled to pure supergravity

In superspace the coupling is

$$\begin{split} \mathcal{L} &= -3 \left(\int d^2\Theta \, 2\mathcal{E} \, \mathcal{R} + c.c. \right) + \left(\int d^2\Theta \, 2\mathcal{E} \, \textit{W}_0 + c.c. \right) \\ &+ \mathcal{L}_{\textit{NEW}} \, . \end{split}$$

And in components (on-shell)

$$\begin{split} e^{-1}\mathcal{L}\Big|_{\lambda=0} &= -\frac{1}{2}R + \frac{1}{2}\epsilon^{klmn}\left(\overline{\psi}_{k}\overline{\sigma}_{l}\mathcal{D}_{m}\psi_{n} - \psi_{k}\sigma_{l}\mathcal{D}_{m}\overline{\psi}_{n}\right) \\ &- \frac{1}{4}F_{mn}F^{mn} - \left(\frac{1}{2}\xi^{2} - 3|W_{0}|^{2}\right) \\ &- \overline{W}_{0}\psi_{a}\sigma^{ab}\psi_{b} - W_{0}\overline{\psi}_{a}\overline{\sigma}^{ab}\overline{\psi}_{b} \,. \end{split}$$

New D-terms

- → No R-symmetry gauging.
- \rightarrow Arbitrary gravitino mass.
- → No smooth supersymmetric limit.
- \rightarrow Cosmological constant: $\frac{1}{2}\xi^2 3|m_{3/2}|^2$.
- → Electric-magnetic duality of the full theory:

$$F_{mn} \rightarrow \epsilon_{mnkl} F^{kl}$$
.

Introducing matter

Chiral superfields

$$\Phi^I = A^I + \sqrt{2}\Theta \chi^I + \Theta^2 F^I.$$

The coupling to standard supergravity reads

$$\mathcal{L}_{K,W} = \int d^2\Theta \, 2\mathcal{E} \left[\frac{3}{8} (\overline{\mathcal{D}}^2 - 8\mathcal{R}) e^{-\frac{1}{3}K(\Phi^i,\overline{\Phi}^i)} + \textit{W}(\Phi^i) \right] + \textit{c.c.}$$

In component form (the bosonic sector) reads

$$\begin{split} e^{-1}\mathcal{L}_{K,W} &= -\frac{1}{2}R - K_{i\bar{j}}\partial A^{i}\partial \overline{A}^{\bar{j}} \\ &- e^{K}\left[(K^{-1})^{i\bar{j}}(W_{i} + K_{i}W)(\overline{W}_{\bar{j}} + K_{\bar{j}}\overline{W}) - 3W\overline{W} \right]. \end{split}$$

Adding the new D-term

For

$$\mathcal{L} = \mathcal{L}_{K,W} + \mathcal{L}_{NEW}$$

Then the scalar potential is

$$V = V_{\text{Standard}} + \frac{\xi^2}{2} e^{2K/3}$$
.

- There is generically a positive definite uplift.
- For constant ξ Kähler invariance is broken.
- SUSY is off-shell linearly realized.

Outlook

- → New embedding of the FI term in supergravity.
- → Is there more unknown new supergravity embeddings?
- → Applications to inflation/cosmology? *Aldabergenov, Ketov,* '17, Antoniadis, Chatrabhuti, Isono, Knoops '18
- → Relation to string theory / brane origin ?

A KKLT-type uplift!

 \rightarrow For

$$K = -3 \ln(T + \overline{T}), W = W(T),$$

we find

$$\mathcal{V} = \mathcal{V}_{\text{Standard}} + rac{\xi^2}{2} rac{1}{(T + \overline{T})^2}.$$

- → This uplift usually constructed in SUGRA with non-linear local SUSY realization. Ferrara, Kallosh, Linde '14, Kallosh, Wrase '14, Bandos, Heller, Kuzenko, Martucci, Sorokin '16
- → Here this construction is straightforward and achieved without introducing non-linear realizations off-shell.

Thank you