

Newton–Cartan (Super)Gravity and Torsion

Dedicated to the memory of Ioannis Bakas

Athanasios Chatzistavrakidis

Rudjer Bošković Institute, Zagreb

Joint work with E. A. Bergshoeff, J. Lahnsteiner, L. Romano, J. Rosseel

arXiv:1708.05414 (JHEP) and work in progress

hep 2018 @ NTU Athens

28 March 2018



Why non-relativistic gravity?

- ✦ Unlike Einstein gravity, Newtonian gravity is not a geometric theory
 - ✦ non-relativistic gravity in an arbitrary frame?
 - ✦ Newton-Cartan: geometric formulation of Newtonian gravity; arbitrary frames
- ✦ New physical applications, mainly in condensed matter physics
 - ✦ Construction of Effective Field Theories, e.g. for the FQHE, chiral superfluids &c.
Hoyos, Son '12; Son '13; Geracie, Son, Wu, Wu '15; Hoyos, Moroz, Son '14; Moroz, Hoyos '15
 - ✦ Universal properties, transport phenomena

Theory of Thermal Transport Coefficients*

J. M. LUTTINGER

Department of Physics, Columbia University, New York, New York

(Received 20 April 1964)

A simple proof of the usual correlation-function expressions for the thermal transport coefficients in a resistive medium is given. This proof only requires the assumption that the phenomenological equations in the usual form exist. It is a "mechanical" derivation in the same sense that Kubo's derivation of the expression for the electrical conductivity is. That is, a purely Hamiltonian formalism with external fields is used, and one never has to make any statements about the nature or existence of a local equilibrium distribution function, or how fluctuations regress. For completeness the analogous formulas for the viscosity coefficients and the heat conductivity of a simple fluid are given.

III. CALCULATION OF THE THERMAL TRANSPORT COEFFICIENTS

Just as the space- and time-varying external electric potential produced electric currents and density variations, so a varying gravitational field will produce, in principle,⁷ energy flows and temperature fluctuations.

large one. In fact, if the gravitational field didn't exist, one could invent one for the purposes of this paper.

Why non-relativistic gravity?

- ✿ Unlike Einstein gravity, Newtonian gravity is not a geometric theory
 - ✿ non-relativistic gravity in an arbitrary frame?
 - ✿ Newton-Cartan: geometric formulation of Newtonian gravity; arbitrary frames
- ✿ New physical applications, mainly in condensed matter physics
 - ✿ Construction of Effective Field Theories, e.g. for FQHE, chiral superfluids &c.
Hoyos, Son '12; Son '13; Geracie, Son, Wu, Wu '15; Hoyos, Moroz, Son '14; Moroz, Hoyos '15
 - ✿ Universal properties, transport phenomena
- ✿ Newton-Cartan geometry is to Newton what Riemann geometry is to Einstein
 - ✿ Coupling of non-relativistic field theories to gravity
 - ✿ Distinguish **geometry** (background fields, symmetries) **vs.** **gravity** (EOMs)

Why torsion?

- ✿ Holography (e.g. Lifshitz) \rightsquigarrow zero torsion is not allowed in CFT

Christensen, Hartong, Obers, Rollier '14; Hartong, Kiritsis, Obers '14-'15

- ✿ To define a non-relativistic energy-momentum tensor, torsion is a key ingredient

Luttinger '64; Gromov, Abanov '14

- ✿ Numerous applications of torsional non-relativistic gravity in condensed matter

Gromov, Abanov '14; Geracie, Golkar, Roberts '14; Geracie, Prabhu, Roberts '16; & c.

Why supersymmetry?

- ✿ Non-relativistic supersymmetric field theories on non-trivial backgrounds?
- ✿ Localization \rightsquigarrow exact partition functions for non-relativistic field theories?
- ✿ But also, what is the supersymmetrization of Newtonian gravity?

To answer such questions, first one has to construct non-relativistic supergravities

see e.g. [Andringa, Bergshoeff, Rosseel, Sezgin '13](#); [Bergshoeff, Rosseel, Zojer '15](#); [Bergshoeff, Rosseel '16](#)

Methods, Goals and Overview

✿ Construct Newton-Cartan Geometry and Gravity with arbitrary torsion

- ✿ Gauging an algebra
- ✿ Null-reduction of GR/conformal gravity
- ✿ Conformal method
- ✓ NC geometry with arbitrary torsion in any dimension
- ✓ NC gravity with arbitrary torsion in 3D
Bergshoeff, A.Ch., Romano, Rosseel '17

✿ Construct off-shell $\mathcal{N} = 2$ Newton-Cartan supergravity with arbitrary torsion

- ✿ Null-reduction of (old-minimal) $\mathcal{N} = 1$ supergravity
- ✿ Supersymmetric curved backgrounds? (only hints in this talk...)
- ✓ Off-shell non-relativistic supergravity with arbitrary torsion in 3D
Bergshoeff, A.Ch., Lahnsteiner, Romano, Rosseel; to appear

Newton-Cartan from gauging

Einstein gravity \sim Gauging of the Poincaré algebra

Newton-Cartan gravity \sim Gauging of the Bargmann algebra

Andringa, Bergshoeff, Panda, de Roo '10

Commutation relations

$$\begin{aligned} [J_{ab}, J_{cd}] &= 4\delta_{[a[c}J_{d]b]}, & [J_{ab}, P_c] &= -2\delta_{c[a}P_{b]}, \\ [J_{ab}, G_c] &= -2\delta_{c[a}G_{b]}, & [G_a, H] &= -P_a, & [G_a, P_b] &= -\delta_{ab}M. \end{aligned}$$

Transformation	Generator	Gauge Field	Gauge Parameter
Time translations	H	τ_μ	ζ
Space translations	P_a	e_μ^a	ζ^a
Spatial rotations	J_{ab}	ω_μ^{ab}	λ^{ab}
Galilean boosts	G_a	ω_μ^a	λ^a
Central charge	M	m_μ	σ

Transformation rules and torsion

The independent fields are $\{\tau_\mu, e_\mu^a, m_\mu\}$ and they transform as 1-forms and as

$$\begin{aligned}\delta\tau_\mu &= 0, \\ \delta e_\mu^a &= \lambda^a_b e_\mu^b + \lambda^a \tau_\mu, \\ \delta m_\mu &= \partial_\mu \sigma + \lambda^a e_{\mu a}.\end{aligned}$$

Inverse vielbeins may be defined using a set of projective invertibility relations

$$e^\mu_a e_\nu^a = \delta_\nu^\mu - \tau^\mu \tau_\nu, \quad e^\mu_a e_\mu^b = \delta_b^a, \quad \tau^\mu \tau_\mu = 1, \quad e^\mu_a \tau_\mu = 0, \quad \tau^\mu e_\mu^a = 0.$$

For each set of generators, there is an associated curvature. In particular

$$R_{\mu\nu}(H) = \partial_{[\mu} \tau_{\nu]} \rightsquigarrow \text{the torsion}$$

$$\tau_{ab} = e^\mu_a e^\nu_b \partial_{[\mu} \tau_{\nu]}$$

$$\tau_{0a} = \tau^\mu e^\nu_a \partial_{[\mu} \tau_{\nu]}$$

geometric constraint	Newton-Cartan
$\tau_{0a} \neq 0, \tau_{ab} \neq 0$	arbitrary torsion
$\tau_{0a} \neq 0, \tau_{ab} = 0$	twistless-torsional
$\tau_{0a} = 0, \tau_{ab} = 0$	zero torsion

Newton-Cartan equations of motion

✦ $R_{\mu\nu}(P^a)$ and $R_{\mu\nu}(M)$ are used to find expressions for $\omega_\mu{}^{ab}(\tau, e, m)$ & $\omega_\mu{}^a(\tau, e, m)$

✦ The remaining two curvatures are related to the equations of motion for NC

✦ For zero torsion

$$\tau^\mu e^\nu{}_a R_{\mu\nu}(G^a) := R_{0a}(G^a) = 0 \quad \rightsquigarrow \quad \nabla^2 \Phi = 0$$

$$R_{c0}{}^c{}_b(J) = 0, \quad R_{ca}{}^c{}_b(J) = 0$$

✦ For arbitrary torsion, it is much harder to establish such equations... see later

Null-reduction of general relativity

Duval, Burdet, Kunzle, Perrin '85; Julia, Nicolai '95

Start with GR in $d + 1$ dimensions, in the second-order formalism $\{E_M^A, \Omega_M^{AB}(E)\}$.

The spin connection and the curvature are given by the standard expressions

$$\begin{aligned}\Omega_M^{BA}(E) &= 2E^{N[A}\partial_{[M}E_{N]}^{B]} - E^{N[A}E^{B]P}E_{MC}\partial_N E_P^C, \\ \hat{R}_{MN}^{AB}(\Omega(E)) &= 2\partial_{[M}\Omega_{N]}^{AB} - 2\Omega_{[M}^{AC}\Omega_{N]C}^B.\end{aligned}$$

In addition, the vielbein transforms under g.c.t.s and local Lorentz transformations as:

$$\delta E_M^A = \zeta^N \partial_N E_M^A + \partial_M \zeta^N E_N^A + \lambda^A_B E_M^B.$$

Assume that we have a **null Killing vector** $\xi = \xi^M \partial_M$ for the metric $g_{MN} \equiv E_M^A E_N^B \eta_{AB}$:

$$\mathcal{L}_\xi g_{MN} = 0 \quad \text{and} \quad \xi^2 = 0.$$

We choose adapted coordinates and split indices as $M = \{\mu, \nu\}$ and $A = \{a, +, -\}$

Notably, this implies that the metric is degenerate (the Killing vector is now $\xi = \xi^\nu \partial_\nu$):

$$g_{\nu\nu} = 0$$

Reduction Ansatz

The suitable Ansatz for the Vielbein and its inverse is cf. Julia, Nicolai '95

$$E_M^A = \begin{matrix} & \mu & & \\ & \begin{matrix} a & - & + \\ \mathbf{e}_\mu^a & \tau_\mu & -m_\mu \end{matrix} & \\ \nu & \begin{pmatrix} 0 & 0 & 1 \end{pmatrix} & \end{matrix}, \quad E^M_A = \begin{matrix} & \mu & & \\ & \begin{matrix} a & v \\ \mathbf{e}^\mu_a & \mathbf{e}^\mu_a m_\mu \end{matrix} & \\ & \begin{pmatrix} \tau^\mu & \tau^\mu m_\mu \\ 0 & 1 \end{pmatrix} & \end{matrix}.$$

Off-shell reduction yields the NC transformations for $\{\mathbf{e}_\mu^a, \tau_\mu, m_\mu\}$ along with

$$\Omega_\mu^{ab}(E) \equiv \omega_\mu^{ab}(\tau, \mathbf{e}, m) = \dot{\omega}_\mu^{ab}(\mathbf{e}, \tau, m) - m_\mu \tau^{ab},$$

$$\Omega_\mu^{a+}(E) \equiv \omega_\mu^{a+}(\tau, \mathbf{e}, m) = \dot{\omega}_\mu^{a+}(\mathbf{e}, \tau, m) + m_\mu \tau^a.$$

On-shell, however, one obtains the NC EOMs **and** $\tau_{0a} = \tau_{ab} = 0$.

Null-reduction of GR: NC geometry with arbitrary torsion, but NC gravity without torsion

Null-reduction of conformal gravity

Bergshoeff, A.Ch., Romano, Rosseel '17

Motivated by the conformal construction of Poincaré gravity

see the book of Freedman and van Proeyen for all details

Poincaré = Conformal + Scalar + gauge-fixing

The full set of gauge fields is $\{E_M^A, \Omega_M^{AB}, b_M, f_M^A\}$; the independent ones transform as

$$\begin{aligned}\delta E_M^A &= \lambda^A_B E_M^B + \lambda_D E_M^A, \\ \delta b_M &= \partial_M \lambda_D + \lambda_K^A E_{MA}.\end{aligned}$$

Null-reduction $\rightsquigarrow z = 2$ **Schrödinger gravity with arbitrary torsion**; anisotropic scaling:

$$\begin{aligned}\delta_D \tau_\mu &= 2\lambda_D \tau_\mu, \\ \delta_D e_\mu^a &= \lambda_D e_\mu^a,\end{aligned}$$

Equivalent to gauging the Schrödinger algebra (conformal extension of Bargmann)

Torsion revisited

Originally $(d+1)$ λ_K^A s. Gauge-fixing $b_\nu = 0$ fixes λ_K^- . Fixing $(d-1)$ more, requires

$$\mathcal{R}_{0a}(H) := \tau^\mu e^\nu{}_a (2\tau_{\mu\nu} - 4b_{[\mu}\tau_{\nu]}) \stackrel{!}{=} 0 \Rightarrow b_a = -\tau_{0a} .$$

This is a **conventional constraint**. Zero torsion is not compatible with dilatations...

geometric constraint	Schrödinger
$\tau_{0a} \neq 0, \tau_{ab} \neq 0$	arbitrary torsion
$\tau_{0a} \neq 0, \tau_{ab} = 0$	twistless-torsional
$\tau_{0a} = 0, \tau_{ab} = 0$	N/A

Newton-Cartan Gravity with Arbitrary Torsion?

Bergshoeff, A.Ch., Romano, Rosseel '17; cf. Afshar et al. '16 for twistless-torsion

Use a non-relativistic version of the conformal construction of relativistic gravity

- Scalar CFT with $\mathcal{L} = \frac{1}{2}\phi\partial^\mu\partial_\mu\phi$
- Coupling to conformal gravity $\partial \rightarrow D_{\text{conf}}$
- Gauge-fix $\phi = 1 \rightsquigarrow \mathcal{L} = R$

In our (arbitrary torsion) case

- SFT for two real scalars (one for dilatations, one for central charge trafos):

$$\partial_0\partial_0\varphi - \frac{2}{M}(\partial_0\partial_a\varphi)\partial_a\chi + \frac{1}{M^2}(\partial_a\partial_b\varphi)\partial_a\chi\partial_b\chi = 0.$$

- Coupling to Schrödinger gravity $\partial \rightarrow D_{\text{Schr}}$
- Gauge-fixing $\varphi = 1$ and $\chi = 0$, and restricting to $d = 3$, leads to a set of EOMs:

$$\mathcal{R}'_{0a}(G^a) + \text{terms}(\tau_{0a}, \tau_{ab}) = 0, \\ \mathcal{R}_{0b}{}^{ba}(\Omega) = 0, \quad \mathcal{R}_{ac}{}^{cb}(\Omega) = 0, \quad \Omega = \dot{\omega} + \text{terms}(\tau_{ab}).$$

Non-relativistic Supergravity

Apply the technique of null-reduction to off-shell (old minimal) $\mathcal{N} = 1, D = 4$ sugra

Bergshoeff, A.Ch., Lahnsteiner, Romano, Rosseel; to appear

The off-shell multiplet comprises the fields $\{E_M^A, \Psi_M, A_M, F\}$

For the bosons, we use the same null-reduction Ansatz as before. For the fermions:

$$\begin{aligned}\Psi_\mu &= (\psi_{\mu-} - m_\mu \psi_{v-}) \otimes \chi_+ + (\psi_{\mu+} - m_\mu \psi_{v+}) \otimes \chi_- , \\ \Psi_v &= \psi_{v-} \otimes \chi_+ + \psi_{v+} \otimes \chi_- .\end{aligned}$$

Nullity of the Killing vector ξ leads to a closed set of constraints:

$$\xi^2 = 0 \quad \Rightarrow \quad E_v^- = 0 \quad \Rightarrow \quad \psi_{v+} = 0 \quad \Rightarrow \quad \epsilon^{ab} \tau_{ab} = 2\sqrt{2} \bar{\psi}_{v-} \psi_{v-} + 12A_v$$

The reduced, non-relativistic off-shell supergravity multiplet contains the fields

$$\{\tau_\mu, e_\mu^a, m_\mu, A_\mu, A_v, F, \psi_{\mu\pm}, \psi_{v-}\}$$

Taking into account the constraint, the supersymmetry algebra closes, as it should

Non-relativistic supersymmetric curved backgrounds?

Relativistic: starting from off-shell sugra, rigid susy theories on curved backgrounds

Pestun '07; Festuccia, Seiberg '11; & c.

- ✿ Classified by studying Killing spinor equations in curved space

Non-relativistic (to date): studied only for torsionless off-shell Newton-Cartan sugra

Knodel, Lisboa, Liu '16

- ✿ “Small” off-shell multiplet $\{\tau_\mu, e_\mu^a, m_\mu, \psi_{\mu\pm}, \mathcal{S}\}$
- ✿ Killing spinor equations \rightsquigarrow solutions exist, but on flat space

In order to ameliorate this, the full off-shell multiplet (allowing torsion) is necessary

Bergshoeff, A.Ch., Lahnsteiner, Romano, Rosseel; work in progress

Epilogue

- ✿ New developments in non-relativistic (super)gravity, in many different directions
- ✓ We constructed Newton-Cartan (super)gravity with arbitrary torsion in 3D
- ✿ Systematic methods, null-reduction and conformal construction
- ✿ Curved supersymmetric backgrounds might exist, from off-shell torsional NC sugra
- ✿ A host of potential applications
 - ✿ Non-relativistic holography
 - ✿ Localization
 - ✿ Condensed matter systems, EFTs
 - ✿ $D = 4$?